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Mountain waves developing inside and aloft stably stratified turbulent boundary layers

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A linear theory of the trapped mountain waves that develop in a turbulent boundary layer is presented. The theory uses a mixing length turbulence model based on Monin-Obukhov similarity theory. First, the backward reflection of a stationary gravity wave (GW) propagating toward the ground is examined. Three parameters are investigated systematically: the Monin-Obukhov length (L_{mo}) , the roughness length (z_0) and the limit value of the mixing length (λ) aloft the "inner" layer. The reflection coefficient appears to strongly depends on the Richardson number aloft the inner layer ($J = \frac{\lambda}{\kappa L_{mn}}$, with κ the von-Karman constant): the reflection decreasing when the stability (J) increases. The influence of the roughness and mixing lengths on the reflection is explained in terms of the depth of a "pseudo"critical level located below the surface: the reflection decreasing when the depth of the "pseudo" critical level decreases. The preferential modes of oscillations occurring in presence of mountain forcing are then analysed, the decay rate of the trapped waves downstream increasing when the reflection decreases. At a certain point nevertheless, when the absorption is large but the boundary layer depth deep enough, there appears trapped modes that interact little with the surface.

^{*}Conceptualization, Formal analysis, Methodology, Writing

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KEYWORDS

Trapped mountain waves, Neutral and stratified boundary layers, Turbulence

1 1 | INTRODUCTION

An early theory for gravity waves developing in the lee of mountain ranges was proposed by Scorer (1949) who
 demonstrated that resonant modes can be excited by mountains when the background wind and stratification vary
 with altitude. More specifically and in the 2D-case, they occur when the Scorer parameter,

$$S(z) = \frac{N^2}{U^2} - \frac{U_{zz}}{U},$$
 (1)

5 decreases with altitude above the surface. Here *N*, *U* and *z* are the background buoyancy frequency, the background

k wind and the altitude, respectively. In this case all the harmonics with horizontal wavenumber k that encounter a

7 turning point h_t where

$$S(h_t(k)) = k^2, \tag{2}$$

are trapped at low levels. A discrete number of them become resonant in the inviscid case, as a result of successive 8 reflections between their turning level and the surface. While Scorer (1949), and more recently Teixeira et al. (2013), ç applied the theory to a two layer atmosphere, cases with smooth variations of S(z) have also been analyzed in 10 numerous studies (see for instance Durran (1986) or Wurtele et al. (1996)). Of particular interest to our study is that 11 of Keller (1994) which includes trapped waves with strong increase of the wind above the surface. Nevertheless and 12 after Scorer (1949), it has been realized that trapped waves can also appear at a sharp density or wind inversion, in 13 which case the interaction with the surface is not as pronounced as when the waves result from multiple reflections 14 between the surface and turning levels (Vosper, 2004; Sachsperger et al., 2017). In the present paper, we will see that 15 near resonant modes that have slight interaction with the surface can also be found when the background atmospheric 16 state varies smoothly in the vertical. 17

If we now return to the initial Scorer (1949)'s theory, one of its weaknesses is that it neglects the dissipative effects 18 occurring within the boundary layer, regardless that the variations with altitude of wind and stratification occurring 19 in the boundary layer are potentially ducting trapped lee waves. This neglect of dissipation yields to overestimate 20 mountain waves amplitude, downslope winds and trapped waves downstream development, as revealed modelling 21 studies (Richard et al., 1989; Miller and Durran, 1991; Georgelin et al., 1994). More recently, Jansing et al. (2022) and 22 Tian et al. (2023) demonstrated that simulations of mountain waves and Foehn strongly depends on the boundary 23 layer parameterizations, parameterizations that are still uncertain in mountainous area (Goger et al., 2019) (see also 24 the review in Serafin et al. (2018)). For completeness note that other deficiencies concerning boundary layer effects 25 over complex terrain are discussed in Tsiringakis et al. (2017), Lehner and Rotach (2018) and Vosper et al. (2018). 26

Beyond numerical models, attempts to understand the more fundamental mechanism at works in the interaction between the trapped lee waves and the boundary layer were motivated by observations. During the Mesoscale Alpine Program (Bougeault et al., 2001), Smith et al. (2002) noticed that a strong absorption by a near stagnant surface layer could inhibit the development of trapped waves, despite favorable conditions aloft. In subsequent papers, they proposed to analyse systematically the absorptive properties of the surface, illustrating that the combination of reduced winds in the boundary layer and dissipation contribute to the absorption (Smith et al., 2006). They characterized a surface reflection coefficient for waves returning to the surface downstream and related it to the spatial decay rate of trapped waves (referred to as q and α respectively). In short, a small reflection is associated with a strong decay.

Nevertheless, in Smith et al. (2006) but also in Jiang et al. (2006), frictional effects are represented by bulk for-35 mula with linear Rayleigh drag, which are not extensively used in models. Indeed the dissipative effects are more 36 often taken into account by introducing turbulence closures based on eddy diffusivity. A fundamental difficulty when 37 using such closures is that the equations have higher order derivative in the vertical (6 compared to 2 for the inviscid 38 Taylor Goldstein equation). In the constant eddy viscosity case, Lott (2007) obtained solutions by using asymptotic 39 techniques where the flow is split into an outer region, where viscosity is negligible, and an inner region, where dissi-40 pative effects compare to disturbance advection (Jackson and Hunt, 1975). This region has depth varying according 41 to an "inner layer" scale h_i that is distinct from the boundary layer depth, and satisfying 42

$$kU(h_i(k)) \approx \frac{\nu'}{h_i^2},\tag{3}$$

where v' is the eddy diffusivity acting on the disturbance produced by the mountain. An important result is that when
 the Richardson number near the surface

$$Ri(z) = \frac{N^2}{U_z^2} \underset{z \to 0}{\approx} J > 0.25,$$
(4)

increases, the surface reflection decreases even in the inviscid limit. Furthermore, in the inviscid limit the reflection is almost total when J < 0.25. This condition is similar to the Richardson-number instability criterion for continuously stratified inviscid parallel flows (Miles, 1961; Howard, 1961). Elaborating on this further, Lott (2007) showed that the neutral modes of Kelvin-Helmholtz instability (Drazin, 1958) can also correspond to trapped lee waves (see also Soufflet et al. (2022)).

A limit of the constant eddy viscosity case is that in reality, turbulent viscosity over a rough surface is more 50 realistically represented using mixing-length theory, the mixing length decreasing when approaching the surface. This 51 introduces other difficulties, for instance one needs to treat the problem using curvilinear coordinates and re-calculate 52 the viscous solutions. To a certain extent, this was accomplished in the past, guite theoretically in Belcher et al. 53 (1993), Belcher and Wood (1996), and Hunt et al. (1988), and more numerically in Weng et al. (1997). However in 54 these papers the trapping of the waves between the surface (or the top of the inner layer) and the turning levels was 55 not fully considered. Specifically, the authors did not capture trapped wave dissipation that may occur in atmospheric 56 boundary layers. In a recent paper, Lott et al. (2023) derived such a theory (hereinafter referred to Part 1), and focused 57 on the nature of the transition from form drag to wave drag, and from downstream sheltering to upstream blocking. 58 During the transition, which is shown to occur when the Richardson number value aloft the surface layer J is close to 59 1, it was noticed that trapped waves deeply affect the dynamics. It was also noticed that for small J the turning levels 60 are too close to the surface for trapped modes to emerge, and that for large J the surface absorption is too large for 61 trapped modes to develop horizontally. Lott et al. (2023) also describe briefly the sensitivity to the Froude number, 62

$$F = \frac{U(\infty)}{N(\infty)L},\tag{5}$$

where *L* is a characteristic length of the mountain. This number controls the significance of the non-hydrostatic effects
 (Yu and Teixeira, 2014). The purpose of the present paper is to assess further these issues, first by calculating the

reflection coefficient and second by trying to relate the flow response to a mountain forcing in terms of this reflection 65 coefficient. As we shall see, the decay rate of the trapped waves is well explained by the absorptive properties of the 66 surface and their horizontal wavenumber is quite controlled by the more conventional inviscid trapped wave dynamics. 67 The plan of the paper is as follows. Section 2 recalls aspects of the formalism used in Part I and needed here. 68 Section 3 evaluates the reflection coefficients. Section 4 analyses the response of the flow to an idealized mountain 60 using the full model presented in Part I, with a focus on the spatial decay rate of trapped modes. Section 5 concludes 70 and the appendix gives solutions to the homogeneous inviscid problem that were not given in Part I and that are 71 required to evaluate the surface reflection coefficients. 72

73 2 | TURBULENCE CLOSURE AND BACKGROUND FLOW PROPERTIES

⁷⁴ Although the evaluation of solutions in the presence of a mountain necessitates curved coordinates, they are not ⁷⁵ needed to calculate surface reflection downstream so we will lighten the formalism and consider cartesian coordinate ⁷⁶ (x, z) in the following. For completeness, we recall here that as in part I, the vertical fluxes of horizontal momentum ⁷⁷ and buoyancy are parameterized by an eddy diffusivity *v* based on mixing length theory,

$$\tau_{xz} = v \partial_z u, q_z = v \partial_z b, \quad v = \Lambda_0^2 \left\| \frac{\partial u}{\partial z} \right\|, \tag{6}$$

where *u* is the zonal wind and $b = -g \frac{\theta - \theta_s}{\theta_s}$ is the buoyancy, θ being potential temperature and θ_s a reference value. For simplicity, we slightly modify the Blackadar formulation for the mixing length - common for neutral flows - but keep the same asymptotes,

$$\Lambda_0 = \lambda \tanh\left(\kappa \frac{z+z_0}{\lambda}\right),\tag{7}$$

where z_0 , κ , and λ are the roughness length, the von Karman constant, and the limit value of the mixing length respectively. The formulation for the mixing length in (7) gives background wind and buoyancy profiles with uniform fluxes,

$$U_{V}(z) = \frac{u_{*}}{\kappa} \ln\left(\frac{\sinh(\kappa \frac{z+z_{0}}{\lambda})}{\sinh(\kappa \frac{z_{0}}{\lambda})}\right), \qquad B_{V}(z) = \frac{b_{*}}{\kappa} \ln\left(\frac{\sinh(\kappa \frac{z+z_{0}}{\lambda})}{\sinh(\kappa \frac{z_{0}}{\lambda})}\right), \tag{8}$$

where $u_* = \sqrt{\tau_s/\rho_s}$ is the friction velocity, and $b_* = g H_s/(\rho_s c_p u_* \theta_s)$ is the buoyancy scale, with τ_s and H_s for surface stress and heat flux and c_p for the air heat capacity per unit mass at constant pressure. The background wind in (8) is going to infinity when $z \to \infty$, which makes that the Scorer parameter in (1) goes to zero: all the harmonics are trapped which is not representative of the real atmosphere. For this reason, we also considered cases where the background wind smoothly becomes constant beyond a boundary layer depth *d* writing

$$U(z) = \frac{u_*d}{\lambda} \tanh\left[\frac{\lambda}{u_*d}U_V(z)\right], B(z) = B_V(z),$$
(9)

keeping in mind that in the limit $d \to \infty U$ and U_{γ} coincide.

Figure 1a shows the wind profile in (9) for $d = \infty$ and d = 1km and for the same parameter values as in Fig. 1 in Part I. When $d = \infty$ the wind shear is almost constant everywhere whereas it is approximately constant below ⁹³ a log character, as expected. Figure 1b, also shows the location of the inner layer, evaluated for the case where the ⁹⁴ mountain has a characteristic length L = 1km (i.e. $h_i(1/L)$ according to Eq. 3, and anticipating for v' the values given in (16a). Figures 1b) and 1c) also show the linear asymptote of U and B when $\lambda \ll z \ll d$, illustrating that

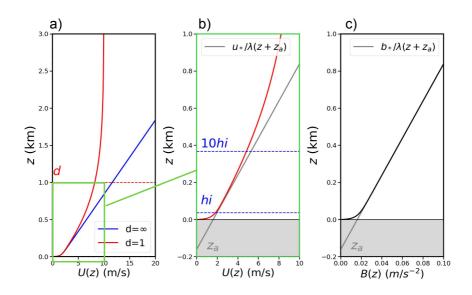


FIGURE 1 Background profiles for $z_0=1$ m, $\lambda=20$ m, $u_*=0.2$ m/s, $b_*=5$ m/s⁻², d=1km, L=1km: a) Background wind U in the constant shear case (blue, see (8)) and the variable shear case (red, see (9). b) Zoom for the wind profile in the boundary layer and in grey its linear asymptote in (10). Are also shown the depth "critical level" depth z_a and the inner layer scale $h_i(1/L)$. c) Buoyancy profile and its linear fit.

95

$$U(z) \underbrace{\approx}_{\lambda \ll z \ll d} \frac{u_*(z+z_a)}{\lambda}, \quad B(z) \underbrace{\approx}_{\lambda \ll z} \frac{b_*(z+z_a)}{\lambda}$$
(10)

⁹⁶ where the parameter,

$$z_a = z_0 - \frac{\lambda}{\kappa} \log\left(2\sinh\frac{\kappa z_0}{\lambda}\right). \tag{11}$$

⁹⁷ At least in the boundary layer and above the inner layer, Figs. 1b)-1c) show that these asymptotes approximate U and ⁹⁸ B in (9) quite well. As we shall see, the parameter z_a can be viewed as the critical level depth for the inviscid part of

⁹⁹ the response.

As said in the introduction, an important parameter of the flow is the Richardson number (4), for the boundary

101 layer profiles in (9)

$$Ri(z) = \frac{N^2}{U_z^2} \approx \begin{cases} 0 & \text{for } z \to 0\\ J = \frac{b_* \lambda}{u_*^2} = \frac{\lambda}{\kappa L_{mo}} & \text{for } \lambda << z << d\\ \infty & \text{for } d << z \end{cases}$$
(12)

where $N^2 = B_z$ and L_{mo} is the Monin-Obukhov length. Still for the boundary layer flow (9) the Froude number (5)

$$F = \frac{U(\infty)}{N(\infty)L} = \sqrt{\frac{u_*^2}{b_*\lambda}} \frac{d}{L} = J^{-1/2} \frac{d}{L}.$$
 (13)

A difficulty is that F changes when J changes. In the present paper, as we are focused on the absorbing properties of 103 the boundary layer we make the choice to systematically vary the dissipative parameters λ and z_0 and the parameter 104 J, F will in most cases take the two contrasting values $F = 1, \infty$ ($d = \sqrt{JL}, \infty$). In these two cases, the fraction of 105 harmonics that stay trapped compared to the number of harmonics excited by the mountain remains constant when 106 the other parameters change. Note nevertheless that to keep F constant we are forced to increase the shear layer 107 depth d when J increases. J will therefore have two contrasting effect: on the one hand trapped lee waves will be 108 more attenuated at the surface when J increases, but on the other the ducting region will increase enabling more 100 modes to develop. We will see that some of them interact little with the surface and can propagate downstream 110 substantially even when surface absorption is large. 111

112 3 | LINEAR SOLUTIONS

¹¹³ To analyse the gravity wave absorption at the surface, we use the boussinesq equations linearized around U and B, ¹¹⁴ and evaluate the behaviour of harmonics with wavenumber k > 0, i.e. considering disturbance fields under the form

$$(u', w', p', b') = (\mathbf{u}, \mathbf{w}, \rho_s \mathbf{p}, \mathbf{b}) e^{ikx},$$
(14)

where u', and w', are horizontal and vertical wind disturbances, whereas p' and b' are disturbances in pressure and buoyancy and ρ_s a reference density. In this Fourier space, the equations we solve are

$$ikU\mathbf{u} + \mathbf{w}\partial_z U + ik\mathbf{p} = \partial_z 2\Lambda_0 u_* \partial_z \mathbf{u}, \tag{15a}$$

117

$$ikU\mathbf{b} + N^2\mathbf{w} = \partial_z \Lambda_0 \left(u_* \partial_z \mathbf{b} + b_* \partial_z \mathbf{u} \right), \qquad (15b)$$

118

$$ikU\mathbf{w} + \partial_z \mathbf{p} - \mathbf{b} = 0, \tag{15c}$$

119

$$i\mathbf{k}\mathbf{u} + \partial_z \mathbf{w} = 0,$$
 (15d)

which are the dimensional form of the homogeneous part of Eq. 29 in Part I, the dissipation being neglected in (15c)
 consistent with the Prandtl approximation. Also, the dissipative terms in (15a) and (15b) result from the linearizations

$$\Lambda_0^2 \left\| \frac{\partial u}{\partial z} \right\| \frac{\partial u}{\partial z} \approx \Lambda_0^2 \left(\frac{dU}{dz} \right)^2 + 2\Lambda_0^2 \left(\frac{dU}{dz} \right) \frac{\partial u'}{\partial z} = u_*^2 + \underbrace{2\Lambda_0 u_*}_{v'} \frac{\partial u'}{\partial z},$$
(16a)

123

$$\Lambda_0^2 \left\| \frac{\partial u}{\partial z} \right\| \frac{\partial b}{\partial z} \approx u_* b_* + \Lambda_0 \left(u_* \partial_z b' + b_* \partial_z u' \right), \tag{16b}$$

124 respectively.

125 3.1 | Outer solution

The scale analysis of the various terms in Eqs. 15 have been done systematically in Part I and for the case where the limit value of the mixing length is much smaller than the characteristic horizontal scale of the waves $\lambda \ll L$. In this case it was shown that all dissipative terms on the RHS of (15) are small and can be neglected at the leading order. This simplification does not allow to satisfy all the boundary conditions and can only be applied in an "outer region" defined by $z \gg h_i$. After verification that $h_i > \lambda$ in the cases we consider, in this outer region the background wind and stratification can be approximated by,

$$U \approx \frac{u_* d}{\lambda} \tanh\left(\frac{z+z_a}{d}\right), \ N^2 \approx \frac{b_*}{\lambda}.$$
 (17)

With these simplified form, the "outer solutions" in the constant shear case $d = \infty$, are expressed in terms of Hankel functions whereas in the variable shear case solutions are expressed in terms of hypergeometric functions all of them exhibiting a critical level at $-z_a$ (i.e below the surface, see appendix for details). What is important is that in all cases we can use analytical solutions with asymptotic behaviours,

$$\mathbf{w}_{I}(z) \underset{z/L \to \infty}{\sim} e^{-m(z+z_{a})}, \ \mathbf{w}_{D}(z) \underset{z/L \to \infty}{\sim} e^{+m(z+z_{a})}$$
 (18a)

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$$\mathbf{w}_{I}(\bar{z}) \underset{z/L \to 0}{\sim} a_{1} (z + z_{a})^{1/2 + i\mu} + a_{2} (z + z_{a})^{1/2 - i\mu},$$

137

$$\mathbf{w}_{D}(z) \underbrace{\approx}_{z/L \to 0} a_{3} \left(z + z_{a}\right)^{1/2 + i\mu} + a_{4} \left(z + z_{a}\right)^{1/2 - i\mu}, \tag{18b}$$

with more details in the appendix. In (18a) and (18b),

$$m = \sqrt{|k^2 - F^{-2}L^{-2}|}, \text{ and } \mu = \sqrt{|J - \frac{1}{4}|},$$
 (19)

$z_0(m)$ $\lambda(m)$	0.5	1	2	$z_0(m)$ $\lambda(m)$	0.5	1	2
5	32	24	16	5	22 (29)	23 (29)	24 (29)
20	196	162	128	20	36 (73)	40 (73)	43 (73)
50	604	518	432	50	50 (135)	52 (135)	57 (135)

TABLE 1 Values of z_a (left) and of $h_i(1/L)$ ($\delta(1/L)$) (right) for different values of λ and z_0 (see Eqs. (11), (3) with ν' in (16a) and (21))

respectively, μ being changed in $i\mu$ when J < 1/4 and m in -im when $k^2 < F^{-2}L^{-2}$. With these conventions, w_I exponentially decays as z increases or is an upward propagating wave when m is imaginary, whereas w_D grows as zincreases or is a downward propagating gravity wave.

As we shall see and to calculate surface reflections when J < 0.25 we will need to analyse the outer solutions in the far field and in the variable shear case rather than just above the inner layer. In this case and to "de-activate" the ducting region we will use hydrostatic solutions. In the outer region this is simply done by changing *m* in (19) by

$$m = -\frac{i}{FL} :$$
 (20)

there is no turning levels and trapped harmonics. For completeness, note that the hydrostatic approximation is always
 made in the very thin inner region.

147 3.2 | Inner solution

To find solutions that match the outer solution and satisfy the 3 surface boundary conditions on the wind (\mathbf{u}, \mathbf{w}) and buoyancy **b** we evaluated numerically in Part I 4 solutions of the inner equations. Note nevertheless that for mathematical convenience, the inner equations we derive do not use vertical distances scaled by the inner scale h_i but by the scale

$$\delta(k) = \left(\frac{\lambda^2}{k}\right)^{1/3}.$$
(21)

¹⁵² We have verified that in all cases we analyse the two scales compare in amplitude h_i always being 1 – 3 times smaller ¹⁵³ than δ (see table 1). Beyond this technical issue, the most important result is that near aloft the inner layer, the ¹⁵⁴ background wind shear is almost constant and the 4 inner solutions needed have asymptotic behaviour

$$\mathbf{w}_{v1} \underbrace{\approx}_{z/\delta \to \infty} (z + z_a)^{1/2 - i\mu}, \ \mathbf{w}_{v2} \underbrace{\approx}_{z/\delta \to \infty} (z + z_a)^{1/2 + i\mu}, \tag{22a}$$

$$\mathbf{w}_{v3} \underbrace{\approx}_{z/\delta \to \infty} (z+z_a)^{-\frac{9+2J}{4}} e^{-\frac{2}{3}\sqrt{j}\left(\frac{z+z_a}{\delta}\right)^{3/2}}, \mathbf{w}_{v4} \underbrace{\approx}_{z/\delta \to \infty} (z+z_a)^{-\frac{5-2J}{4}} e^{-\frac{2}{3}\sqrt{\frac{j}{2}}\left(\frac{z+z_a}{\delta}\right)^{3/2}}.$$
 (22b)

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The interest here is that when J > 0.25, we know from Booker and Bretherton (1967), that the asymptotic behaviour of \mathbf{w}_{v1} in (22a) corresponds to that of a downward propagating wave and that of \mathbf{w}_{v2} to an upward propagating wave (in the convention k > 0). When J < 0.25, the two solutions need to be combined to build vertically propagating disturbances which complicates the analysis (see next section). Finally, \mathbf{w}_{v3} and \mathbf{w}_{v4} decay exponentially when zincreases, they are entirely "dissipative" solutions as indicate their exponential decay rate scaled by the inner layer scale δ .

162 4 | ABSORPTION BY THE SURFACE LAYER

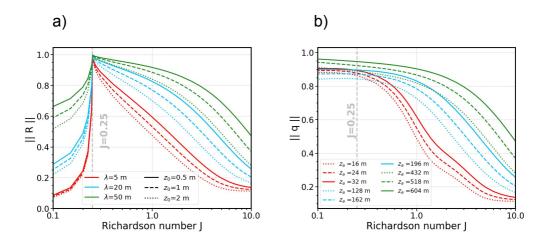
The absorbing properties of the surface can be estimated using the solution above the inner layer. Following Lott (2007), the solution can be expressed as the superposition of the four viscous solutions with asymptotic behaviour given in (22), i.e.

$$\mathbf{w} = \mathbf{w}_{v1} + R\mathbf{w}_{v2} + c\mathbf{w}_{v3} + d\mathbf{w}_{v4},$$
(23)

with similar expressions for the horizontal wind and buoyancy. From the numerical integration of the four viscous solutions to the surface described in Part I we then evaluate the coefficients R, c, and d so as to satisfy the three surface boundary conditions,

$$w(z = 0) = u(z = 0) = b(z = 0) = 0,$$
(24)

while imposing "unit" amplitude upward propagating solution (w_{v1} see (23)).



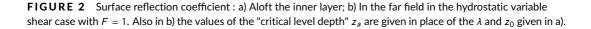


Figure 2a) shows the amplitude of the reflection R for different roughness length (z_0) and mixing length (λ) as a

¹⁷¹ function of *J* and for the dominant wavenumber $k = 1/L = 10^{-3} \text{m}^{-1}$. When J > 0.25, we see that this coefficient ¹⁷² decreases with stability, which is consistent with the fact that solutions with larger μ oscillate more rapidly in the ¹⁷³ vertical when *J* increases and are therefore more absorbed. This behavior is illustrated in Fig. 3: for a given and ¹⁷⁴ quite small z_a , the inviscid solution (black curves) oscillates much more when *J* is large, at least in the region where ¹⁷⁵ dissipative effects start to be significant (region arbitrarily placed below $5h_i$ here).

176 We can also notice that the reflection coefficient decreases with the roughness length, consistently with the fact that enhanced roughness length leads to more dissipation. This has been shown and quantified by Jiang et al. (2006) 177 in a simple model of wave reflection. Quite surprisingly nevertheless, one also sees that absorption decreases (R 178 increases) when the limit value of the mixing length λ increases as if more dissipation results in less absorption. To 179 interpret this, one needs to recall that the oscillatory behaviour in the solutions \mathbf{w}_{v1} and \mathbf{w}_{v2} is more pronounced near 180 the surface when the apparent critical depth is not too large (z_a in Eq. (11)). As shown in table 1, this depth decreases 181 when z_0 increases but increases when λ increases, which explains the absorbing behaviour seen in Fig. 2a. Again, the 182 fact that absorption becomes less sensitive to J when z_a is large is also illustrated by the red curves in Fig. 3: when 183 $z_a = 162$ m, the differences in oscillatory behaviour oscillations between J = 0.3 and J = 4, are not as pronounced 184 when z_a is smaller. Accordingly the decrease in reflection when J increases is therefore less pronounced (compare 185 the dashed blue and red dot curves in Figs. 2). Finally, it is important to notice that |R| is not sensitive to the choice 186 of k because the inner equations we solved in Part I become independent of k when inner varibles are used. This of 187 course stays true in the limit of validity of our analysis, that is when: 188

$$\lambda < \delta(k) < d. \tag{25}$$

The results for |R| when J < 0.25 are more problematic to interpret because in this case \mathbf{w}_{v1} and \mathbf{w}_{v2} in Eq. (23) cannot be associated with downward and upward gravity waves. To circumvent this difficulty one needs to look further aloft for instance where the wind becomes constant in the variable shear case. There, as the two solutions \mathbf{w}_D and \mathbf{w}_I in (18) only represent downward and upward propagating waves when *m* is imaginary, we make the hydrostastic approximation to ensure that it is always the case (see (20)). With this approximation, the harmonics are no longer trapped and we can analyse the wave reflection in $z \to \infty$ by writing the solution in the outer region under the form,

$$\mathbf{w} = U_{p}\mathbf{w}_{I} + D_{o}\mathbf{w}_{D} \underbrace{\approx}_{z \to 0} (U_{p}a_{1} + D_{o}a_{3}) (z + z_{a})^{1/2 + i\mu} + (U_{p}a_{2} + D_{o}a_{4}) (z + z_{a})^{1/2 - i\mu},$$
(26)

the matching with the inner solution in being done by neglecting the viscous solutions in (23) that are exponentially small when $z/\delta \rightarrow \infty$. This yields the total reflection coefficient,

$$q = \frac{U_p}{D_o} = \frac{a_4 - a_3 R}{a_2 - a_1 R}.$$
 (27)

As shown in Fig. 2b), when F = 1 ($d = \sqrt{JL}$), the amplitude of this coefficient |q| behaves almost as |R| when J > 0.25and becomes near 1 (near total reflection) when J < 0.25. Again, as in the outer layer the hydrostatic solutions are such that the a_i s do not depend on k and as |R| does not depend on k, |q| does not vary much with k either. There is nevertheless a weak sensitivity: keeping values in the range 0.5/L < k < 2L we found that slightly more absorption occurs when k decrease, i.e. when the inner layer h_i get closer from the boundary layer depth d (not shown).

This interpretation that the surface reflection is near 1 when J < 0.25 nevertheless has a limit. Indeed, the coefficient *q* is more than a surface reflection coefficient since partial reflections of the incident wave can occur where

- ²⁰⁵ We verified that these partial reflections have small qualitative impacts by making sensitivity tests of our results to the
- ²⁰⁶ Froude number *F* (and hence *d*, not shown). This weak sensitivities to the curvature around *d* probably follows that
- the tanh function used to stop the infinite growth with altitude of the boundary layer wind U_{ν} in (9) is very smooth. This tanh profile was chosen in Lott (2007) to minimize these partial reflections. It permits to say with confidence that
- the amplitude of the hydrostatic approximation of q essentially measures surface reflection.

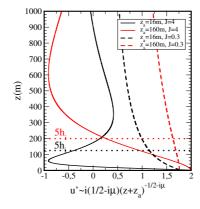


FIGURE 3 Schematic representation of the horizontal velocity field associated with the Booker and Bretherton (1967)'s downward solutions $\mathbf{u}_1 = \frac{i}{k} (1/2 - i\mu) (z + z_a)^{-1/2 - i\mu}$ for z_a =16m, 162m, and J = 0.3, 4. Only the real part is shown for conciseness.

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210 5 | TRAPPED LEE WAVE DEVELOPMENT

To determine the manner in which the trapped waves change as flow stability and dissipations change, we next use the model presented in Part I, and where the mountain is represented by a 2D gaussian ridge of characteristic horizontal length *L*:

$$h(x) = He^{-\frac{x^2}{2L^2}}$$
(28)

where *H* is the maximum mountain height. The ratio S = H/L is fixed to 0.2 keeping L = 1 km. With these parameters, the gaussian mountain forces harmonics with dominant wavenumber around k = 1km⁻¹. In this model where disturbances come from the surface, and in contrast with the previous section, we recall that each harmonics $\mathbf{w}(k, z)$ varies in the vertical and when z >> d according to

$$e^{-m(z+z_{g})} \text{ where } m = \begin{cases} -i\sqrt{F^{-2}L^{-2}-k^{2}} & \text{for } k < k_{c} = (FL)^{-1} \\ \sqrt{k^{2}-F^{-2}L^{-2}} & \text{for } k > k_{c} \end{cases}$$
(29)

where k_c is the cutoff wavenumber separating trapped and freely propagating harmonics.

219 **5.1** | Wave field

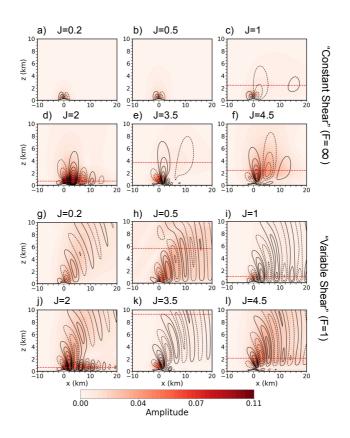


FIGURE 4 Vertical velocity fields w(x, z) for a roughness length $z_0 = 1$ m and a mixing length $\lambda = 20$ m ($z_a = 162$ m). The 6 top figures are for the "constant shear case" ($F = \infty$), and the 6 bottom figures are for the "variable shear case" with F = 1. Between panels in each two case only varies the Richardson number J. In all panels, the contour interval is fixed to 0.01 m.s⁻¹ and the color represents the amplitude of w. The red dashed lines give the altitude where the characteristics of the dominant trapped waves are extracted (see section 5.2).

The vertical velocity fields are plotted in Figure 4 for different values of the Richardson number J and for $z_0 = 1$ m and $\lambda = 20$ m. The 6 top panels show the wave field for the constant shear case, i.e. for $F = \infty$, and thus m = k (see (29)). The wave fields obtained with the variable shear (F = 1) are sketched in the 6 bottom panels.

When $F = \infty$ the 6 top panels show that the wave field decreases rapidly with altitude, as a result of the presence 223 of turning heights (and hence real m's) for each wavenumber. For small J, these turning heights $h_t(k)$ in (2) are close 224 to the surface compared to the horizontal scale (i.e. $h_t(1/L)/L < 1$), whereas they are quite far for large J. Accordingly 225 and for small J the trapping region is too narrow vertically compared to the horizontal scale of the waves and trapped 226 modes do not emerge as Figs. 4a), 4b) show. Some signal downstream the mountain start to appear at J = 1 (Fig. 4c)), 227 downstream decaying trapped waves dominating the response for J > 1. For J = 2, we observe one dominant mode 228 confined at low level and substantially decaying downstream (Fig. 4d)) whereas for J = 3.5 two modes coexist (Fig. 4e)): 229 the longest mode appears at higher altitude than for smaller J whereas the second much shorter mode emerges from 230 the lower part of the flow, immediatly downstream the mountain and confined near the surface. Note that for J = 4.5231 this low level mode is extremely attenuated consistent with the fact that the surface is strongly absorbing. The fact 232 that the responses present multiple modes and that the dominant one is confined near the surface for small J and 233 developing higher for larger J's are studied further in section 5.2. 234

When F = 1, the six lowest panels in Fig. 4 show that an important feature of the wave field arises from the 235 harmonics that do not encounter a turning altitude (i.e. for which we have $k < k_c$). These harmonics are free to 236 propagate in the far-field and combine to form a system of upward propagating waves. When J = 0.2, in Fig. 4g) this 237 far-field component dominates aloft, there is very small downstream signal at low level (somehow reminiscent of the 238 very small low-level signal in the constant shear case in Fig. 4a)). As when $F = \infty$ in Fig. 4a) the turning levels are 239 too close from the surface for trapped modes to emerge. Low level trapped waves become more substantial when J 240 increases (4h), 4i), 4j), due to the fact that the ducting region thickens. Interestingly, for J = 0.5, the trapped wave 241 merges with the system of upward propagating waves above the mountain. As we shall see in the next section, this 242 occurs because the dominant trapped mode that first arise when J increases have horizontal wavenumber near the 243 "cut-off" value separating trapped and propagating harmonics (e.g. k = 1/(FL), see (19)). As J increases further, the 244 wave signal near the surface becomes distinct from that in the far field. It also decays downstream faster when J 245 increases from J = 1 to J = 2 (see Fig. 4i) and 4j) below z = 2km). This low level signal becomes very small for J = 3.5246 (Fig. 4k)), we presume that in this case the surface absorption is too large for low level trapped modes to develop 247 downstream. For even larger J in Fig. 4l), and similarly to the case with $F = \infty$ in Fig. 4j) a second mode emerges 248 above the first one, which is no longer confined near the surface but quite significant and much less attenuated. 249

250 5.2 | Trapped waves

The downstream evolution of dissipated trapped mountain waves can be characterized using a description in term of wavepacket of complex wavenumber (Teixeira and Argaín (2022)). To follow this approach in a diagnostic context we will use surrogates of the form

$$\mathbf{w}_{S} = A e^{ikx - \alpha x} \tag{30}$$

where the amplitude *A* is a complex, α denotes the downstream decay rate and *k* the horizontal wavenumber. To evaluate α and *k*, we minimize the misfit between $\mathbf{w}(x, z_{\text{max}})$ and the surrogate, where z_{max} is the altitude at which the trapped waves present a relative maximum in the amplitude of \mathbf{w} . This altitude is obtained by averaging between $k_{\text{struct}} < k_{\text{struct}} < k_{\text{str$

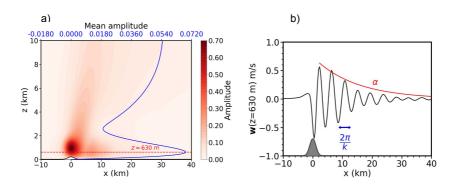


FIGURE 5 Extraction of α and k for $z_a = 162$ m and J = 2. (a) Amplitude of the vertical velocity perturbations obtained by combining w and its Hilbert transform (colors), amplitude averaged between x = 5 km and x = 40 km (blue) and altitude z_{max} of the corresponding maximum amplitude (red). (b) vertical velocity w at z_{max} and mountain (grey).

Figure 6 shows how the decay rate α and the wavenumber k are affected by the Richardson number J and the 260 critical level depth z_a . For a fixed $z_a = 162m$, the decay rate in Fig. 6a) tends to increase with J consistent with the 261 fact that the waves are more absorbed by the surface (see Fig. 2b)). Nevertheless the approach for tracking the low 262 level wave properties produces a jump for k(J) and $\alpha(J)$ as J increases. This behaviour is related to the co-existence 263 of two damped modes and indeed, when the transition occurs, the pair (k, α) that minimizes $w - w_S$ captures the 264 wave that is less absorbed. In general, the less absorbed wave is that which is less confined at low level. This upper 265 level trapped wave is clearly apparent when J = 4.5 in both the constant shear and variable shear case in Figs. 4f) 266 and 41) respectively. This is in contrast with the cases at lower J where the lower level trapped wave is dominant (for 26 J = 2 see Figs. 4d) and 4l)). Note that after the jump in α , the decay rate continues to increase when J increases, but 268 the increase is less pronounced because for such waves the interaction with the surface is not as strong. Note also 269 that the decay rates are more pronounced in the constant shear cases ($F = \infty$) then when F = 1, again because the 270 waves are more confined and return faster to the surface where they are absorbed. Note also that in the cases with 271 $F = \infty$ the dominant modes often have smaller k's than in the variable shear case, which simply follows that waves 272 with k < 1/L cannot be trapped when F = 1. Similarly, Figs. 6c), 6d) show α and k respectively and as functions of z_a 273 for fixed values of J. It is interesting to emphasize here that for a given J, the damping rate is larger for small z_a , as 274 a result of more dissipation in the lower layer. This is consistent with the results of section 4 showing that waves are 275 more absorbed when the critical level is near below the surface (z_a small). 276

Some prediction of the preferred wavenumbers can also be obtained using the reflection coefficient resonances, 277 that is from |q(k)| in (27) but without making the hydrostatic approximation: for trapped waves and when |q(k)|278 presents a maximum when k varies beyond the cut-off wavenumber ($k_c = 1/(FL)$ see (19)), the exponentially decaying 279 inviscid solution (w_I in (18a)) largely dominates the exponentially growing one (w_D in (18a)). As the Richardson number 280 at the surface Ri(z = 0) = 0 we follow Lott (2007) and assume that the resonance that corresponds to trapped 281 waves are those occurring near the cut-off wavenumber ($k_c = 0$ in the constant shear case and $k_c = 1 \text{ km}^{-1}$ in the 282 variable shear case). In the following, we therefore only keep the first trapped wave modes with values of k larger 283 than the cut-off wavenumber and giving maxima in |q(k)|. Figure 6b) shows that in the constant shear case there is 284 good agreement between the first two modes obtained from the resonances of |q| and the mode captured using the 285

model. In this case when multiple modes are possible, the diagnostics from the model presented before capture the mode with smaller wavenumber k since it is less confined near the surface. In the variable shear case, we find some correspondence as well, at least when J < 3 and for the first maximum of |q|, but proliferation of adjacent resonances in |q| when J is larger make the correspondence less straightforward (not shown).

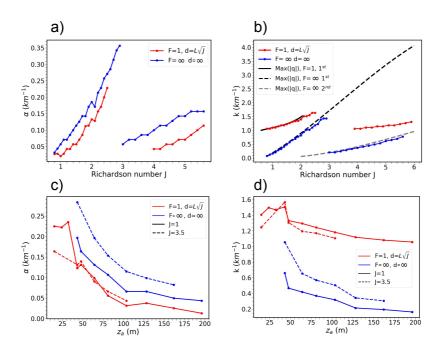


FIGURE 6 Spatial decay rate α (a,c) and wavenumber k (b,d) extracted from the model. (a,b) for $z_a = 162$ m and as function of the Richardson number J; (c,d) For two values of J and as a function of the critical level depth z_a . The thin solid lines with dots give the results obtained from the minimization problem when the Froude number $F = \infty$ (blue) and F = 1 (red). The thick solid lines in shades of grey in b) give the first two modes from the maxima of |q(k)|, and for k increasing beyond $k_c = 1/(FL)$, i.e. $k_c = 0$ when $F = \infty$ and $k_c = 1$ km⁻¹ when F = 1

290 5.3 Agreement with the inviscid theory

Figure 6b)a and 6d) also shows that the dominant wavenumber of the trapped modes have a tendency to increase when *J* increases and to decrease when z_a increases. To a large extent this is more related to inviscid dynamics than to dissipations. To establish this, we next ask ourselves if the much simpler inviscid theories developed in the past stay valid. For this purpose, we consider solutions to the inviscid Taylor Goldstein equation taking for incident flow (17). For such flow where the surface "log"-layer is absent, the potentially resonant inviscid solutions for $F = \infty$ and F = 1 are given by the w_I in (33a) and (36a)) respectively and for wavenumbers *k* such that $w_I(k, z = 0) = 0$.

Figure 7a) and 7b) plots the wavenumbers extracted from the full dissipative theories as a function of the wavenumbers predicted by the inviscid theory and in the constant shear case (F = 0). In these two figures all the parameters of interest are changed (z_0 , λ , J) but we managed to distinguish two quite separate regimes. On the left panel (Fig. 7a),

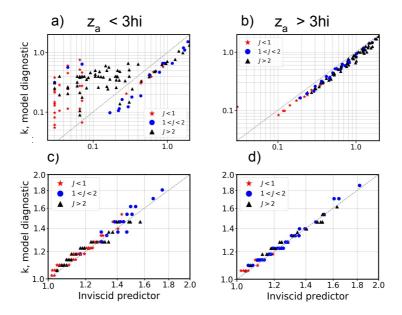


FIGURE 7 Wavenumbers evaluated with the full dissipative theories as a function of the wavenumbers predicted by the inviscid theory. The upper panels show the constant shear cases $d = \infty$, the lower panels show the variable shear case, $d = \sqrt{J}L$. On the left (right) are cases with critical level near (far) below the surface (i.e. $z_a < 3h_i$ and $z_a > 3h_i$ respectively).

the depth of the critical level $z_a < 3h_i$ on the right (Fig. 7b) $z_a > 3h_i$. It therefore shows that the inviscid theory is quite right when the critical level is far below the surface compared to the inner layer depth, but fails when it is quite near. The same comparison in the variable shear case (F = 1) in Figs. 7c) and 7c) deliver about the same message except maybe that the inviscid predictor is more accurate than in the constant shear case since the inviscid predictor work quite well when z_a is quite near below the surface (compare Figs. 7a) and 7c)).

305 6 | SUMMARY

The purpose of this paper is to analyse the trapped waves that can occur when a turbulent boundary layer interacts 306 with a low mountain ridge. We use for that a linear theory developed in a companion paper (Part I), and where 307 turbulence is represented by a viscosity which amplitude varies according to the mixing length theory. We are aware 308 that such a theory oversimplifies the interaction between turbulence and the obstacle, for instance neglecting the 309 impact of the disturbance on the mixing length, or the dependence of the mixing length on flow stability. It also 310 neglects that the disturbances produce turbulence in the outer part of the flow, i.e. that turbulence and dissipative 311 effects penetrate in the outer layer. Much more fundamentally it also neglects that at the horizontal scales we analyse 312 the turbulent eddies can backscatter on the large scale, effect that is entirely absent when representing turbulence 313 with eddy viscosity (see Sun et al. (2015)). This being said, we believe that our theory stays more realistic than the 314 theories developed so far to describe the interaction between mountains and boundary layers (Smith et al., 2002; 315 Lott et al., 2020a). Also, the closure we analyse is somehow representative of the turbulence closures adopted in the 316 atmospheric mesoscale models that are used to do large-eddy simulations in mountainous areas (Doyle et al., 2011). 317 In this respect, our theory could help interpreting what occurs in these models. 318

The first message is that near the surface the undisturbed boundary layer flow has null Richardson number by 319 construction, and is less absorptive than in the constant viscosity case analysed for instance in Lott (2007) (compare 320 Figs. 2a) and 2b) with and Figs. 2 and 3 in Lott (2007) respectively). We interpret this by the fact that the surface 321 critical level absorption at work in Lott (2007) is less effective because the critical level migrates below the surface 322 in the cases considered here. As an illustration, we find that absorptions only compare to those in Lott (2007) when 323 the critical level depth z_a is small. We also find that this depth is large and absorption is small when the limit value 324 of the mixing length is large, as if more dissipation resulted in less absorption! This is because at fixed roughness 325 length, increasing λ yields a larger critical level depth z_a . Due to the central role of this parameter, we could suggest 326 to diagnose it in practice, for instance by fitting a linear function to the boundary layer winds above the surface layer 327 and identify at which depth this function is null below the surface. We also found that absorption increase with the 328 Richardson number value of the background flow in the boundary layer, and this occurs because more oscillatory 329 behaviours near the surface result in more absorption, as in Lott (2007). 330

We next relate the absorptive properties of the inner layer to the decay rate of the trapped waves downstream 331 and found that they are indeed well related. We also found that the trapped waves often has horizontal wavenumbers 332 quite near the cut-off value $\sqrt{N(\infty)}/U(\infty)$, a behaviour we also found when the Richardson number at the surface 333 is null in Lott (2007), as is always the case here. For quite large Richardson numbers aloft the inner layer J, which 334 are cases where the surface absorption is strong, the analysis also reveals the presence of gravity waves propagating 335 downstream. Their decay rate is quite small despite the fact that the surface is supposedly absorbing them substan-336 tially. This new class of waves are characterized by the fact that they have small amplitude near the surface, and large 337 amplitude at the boundary layer height and above. As these waves appear when J is quite large, the fact that the 338 ducting region depth becomes quite large is presumably central. Indeed they are reflected back toward the surface 339

at turning points much higher than when J is small and we can presume that they reach the surface at much larger
 distance downstream than when J is smaller. These waves are therefore absorbed at the surface at a much longer
 distance from the obstacle than those that stay confined near the surface.

We have also tried to test if the more classical inviscid theories can still be applied to predict trapped waves, and when considering boundary layer flow without surface "log"-layer. We found that it is general the case, at least for predicting the horizontal wavenumbers. Some discrepancies can nevertheless appear when dissipations are large, which in our case means that the critical level depth z_a is small.

347 Data availability statement

The theoretical model used to support the findings of this study is available from the corresponding author upon request.

350 A | OUTER SOLUTIONS IN THE CONSTANT SHEAR CASE

In the outer region, the solutions are the inviscid solutions of the Taylor Goldstein equation,

$$\frac{d^2\mathbf{w}}{dz^2} + \left(\frac{N^2}{U^2} - \frac{U_{zz}}{U} - k^2\right)\mathbf{w} = 0$$
(31)

³⁵² which in the constant shear case approximates into the Bessel's equation

$$\frac{d^2 \mathbf{w}}{dz^2} + \left(\frac{J}{(z+z_a)^2} - k^2\right) \mathbf{w} = 0.$$
 (32)

A solution in term of Hankel function and for exponentially decaying disturbances in the farfield is developed in Lott et al. (2020b), it is extended here to include exponentially growing solutions, and writes when k > 0:

$$\mathbf{w}_{I} = i\sqrt{\frac{\pi k(z+z_{a})}{2}}e^{-\mu\pi/2}H_{i\mu}^{(1)}(ik(z+z_{a})) \underset{z \to \infty}{\approx} e^{-k(z+z_{a})},$$
(33a)

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$$\mathbf{w}_{D} = \sqrt{\frac{\pi k (z + z_{a})}{2}} e^{+\mu \pi/2} H_{i\mu}^{(2)} (ik(z + z_{a})) \underbrace{\approx}_{z \to \infty} e^{+k(z + z_{a})},$$
(33b)

where μ is given in (19). In the far field, the asymptotic behaviours of (33) are as in (18) keeping in mind that m = kin (19) when $d = \infty$. Near the surface, (33) are as (18) as well (see 9.1.3, 9.1.4, and 9.1.7 in AS), and by taking,

$$a_{1} = \frac{-i\sqrt{\pi}}{\sinh(\mu\pi)\Gamma(1-i\mu)} \left(\frac{k}{2}\right)^{1/2-i\mu}, \quad a_{2} = a_{1}^{*},$$
(34a)

$$a_{3} = \frac{\sqrt{\pi}e^{\mu\pi}}{\sinh(\mu\pi)\Gamma(1-i\mu)} \left(\frac{k}{2}\right)^{1/2-i\mu}, \quad a_{4} = -\frac{\sqrt{\pi}e^{-\mu\pi}}{\sinh(\mu\pi)\Gamma(1+i\mu)} \left(\frac{k}{2}\right)^{1/2+i\mu}.$$
 (34b)

359 B | OUTER SOLUTIONS IN THE VARIABLE SHEAR CASE

In the variable shear case $d \neq \infty$, the solution for upward propagating waves was developed in the appendix of Soufflet et al. (2022) and the calculation is extended again here to include downward propagating waves. More precisely, and when taking $r = \tanh^2((z + z_a)/d)$, the Taylor-Goldstein equation (31) becomes Eq. A1 in Soufflet et al. (2022)) it is there transformed into an hypergeometric equation when accounting for the change of variable:

$$\mathbf{w} = r^{\frac{1}{4} + i\frac{\mu}{2}} (1 - r)^{-\frac{md}{2}} W,$$
(35)

where μ and *m* are in (19) again (see also Eq. 8 in Lott et al. (1992)). Two inviscid solutions are :

$$\mathbf{w}_{I} = 2^{-md} r^{1/4 + i\mu/2} (1 - r)^{-md/2} W_{2(1)} \underset{z \to \infty}{\approx} e^{-m(z + z_{a})}$$
(36a)

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$$\mathbf{w}_{D} = 2^{+md} r^{1/4 + i\mu/2} (1 - r)^{-md/2} W_{1(1)} \underbrace{\approx}_{z \to \infty} e^{+m(z + z_{\theta})}$$
(36b)

And where \mathbf{w}_I and \mathbf{w}_D are unique amplitude waves propagating upward and downward respectively when *m* is imaginary. The solutions $W_{1(1)}$ and $W_{2(1)}$ are expressed with the hypergeometric function *F*:

$$W_{1(1)} = r^{-i\mu}F\left(-\frac{1}{4} - \frac{i\mu}{2} - \frac{md}{2}, \frac{5}{4} - \frac{i\mu}{2} - \frac{md}{2}; 1 - md; 1 - r\right)$$
(37a)

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$$W_{2(1)} = (1-r)^{md} F\left(\frac{i\mu}{2} + \frac{5}{4} + \frac{md}{2}, \frac{i\mu}{2} - \frac{1}{4} + \frac{md}{2}; 1 + md; 1 - r\right)$$
(37b)

Then, to evaluate the solution near the surface, we use relations (15.3.6) given in Abramowitz and Stegun (1964), to link $W_{1(1)}, W_{2(1)}, W_{1(0)}, W_{2(0)}$ with the help of A_1, A_2, A_3, A_4 calculated in the appendix of Soufflet et al. (2022):

$$W_{1(0)} = A_1 W_{1(1)} + A_3 W_{2(1)}, \qquad W_{2(0)} = A_2 W_{1(1)} + A_4 W_{2(1)}.$$
 (38)

That leads to a total solution (combination of upward and downward propagating waves),

$$D_o \mathbf{w}_D + U_\rho \mathbf{w}_I = r^{\alpha_1} (1 - r)^{\gamma_1} (D_o(a_4 \mathbf{w}_{1(0)} + a_3 \mathbf{w}_{2(0)}) + U_\rho(a_2 \mathbf{w}_{1(0)} + a_1 \mathbf{w}_{2(0)})),$$
(39)

 $_{372}$ $\,$ where $w_{1(0)}$ (w_{2(0)}) is related to $\mathcal{W}_{1(0)}$ ($\mathcal{W}_{2(0)})$ by using Eq. 35 and

$$a_{j} = \begin{cases} (-1)^{j-1} \frac{2^{-md}A_{j}}{A_{1}A_{4} - A_{3}A_{2}} d^{-1/2 + (-1)^{j-1}i\mu} & \text{for } j=1,2 \\ (-1)^{j} \frac{2^{md}A_{j}}{A_{1}A_{4} - A_{3}A_{2}} d^{-1/2 + (-1)^{j-1}i\mu} & \text{for } j=3,4 \end{cases}$$
(40)

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