1 2 3	Can parameterizations reproduce the gravity waves momentum fluxes and drag simulated by a global high-resolution model?
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10	keypoints
11	In the lower stratosphere, parameterizations simulate well the regions where gravity waves
12	are large in a high-resolution global model.
13	An attenuation of the resolved gravity waves aloft the subtropical jets is described by
14	the parameterization but is underestimated.
15	The parameterizations give good estimates of the mountain waves but underestimate
-	the convective waves and overestimate the frontal waves
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17 Abstract

We compare the gravity wave (GW) parameterizations used in the IPSLCM6 cli-18 mate model with the GWs resolved in the ICON global model with 5 km horizontal res-19 olution. The parameterizations are run offline using ICON fields coarse-grained to a 100 20 km grid and compared to the GWs with smaller scales that are resolved in ICON. Over-21 all, the drags are comparable, the momentum fluxes align well, and each GW parame-22 terization (fronts, convection, and mountains) plays a role at geographical locations con-23 sistent with ICON. Among the differences, we find that in ICON, GWs are substantially 24 25 attenuated aloft the subtropical jets; this is underestimated by the parameterizations. It could be corrected by tuning the characteristic phase speeds or the breaking criteria 26 in the parameterizations. It also seems that ICON underestimates frontal waves in the 27 mid-latitudes, that the parameterizations underestimate the convective waves in the trop-28 ics, and that the mountain waves are more alike. 29

³⁰ Plain Language Summary

To simulate the middle atmosphere, climate models use parameterizations of small-31 scale gravity waves that are generated in the troposphere and break in the middle at-32 mosphere. The direct in situ observations of these waves are sparse and non-global, while 33 remotely-sensed satellite observations are more global but have quite coarse resolutions. 34 To compensate for these deficiencies, the recent high-resolution global simulations of the 35 earth atmosphere are promising because they explicitly solve a good fraction of the grav-36 ity wave spectra and their dissipation. Here we show that these simulations can help to 37 validate relationships of parameterized gravity waves to their sources and to adjust their 38 characteristic phase speed. Such comparisons can also establish whether high-resolution 39 simulations resolve enough GW momentum flux to simulate the observed middle atmo-40 sphere climate. 41

42 **1** Introduction

Gravity waves (GWs) are generated by various sources, such as flow over moun-43 tains (Lilly & Kennedy, 1973), convection (Fovell et al., 1992), and imbalances in jets 44 and fronts (Plougonven & Zhang, 2014). As GWs propagate vertically, they carry hor-45 izontal momentum and affect the large-scale circulation when they break (Dunkerton, 46 1997; Alexander & Rosenlof, 2003; Fritts & Alexander, 2003; McLandress & Shepherd, 47 2009). Since the horizontal scale of the GWs can be much shorter than the 1° to 2° hor-48 izontal resolution of climate models, they need to be parameterized (Alexander et al., 49 2010). State-of-the-art GW parameterization schemes use quite a distinct treatment for 50 orographic GWs compared to non-orographic GWs. 51

In the orographic GW schemes, quite local and detailed information about the source 52 is explicitly taken into account. Such orographic GW parameterizations have been used 53 for almost 40 years and have proven to be successful in reducing biases in the troposphere 54 (Palmer et al., 1986; Lott, 1999; Scinocca & McFarlane, 2000). They have been validated 55 by dedicated in situ observations during field campaigns (Lott & Miller, 1997; Smith & 56 Kruse, 2018) and high-resolution limited area model simulations, i.e., validations that 57 remain very local in space and time. In the non-orographic GW schemes, the numeri-58 cal and theoretical complexities of treating a large ensemble of waves have led to the de-59 velopment of global spectral schemes (Hines, 1997; Warner & McIntyre, 1999) or to stochas-60 tic methods (Lott et al., 2012). Among these schemes, some include sources from con-61 vection (Beres et al., 2005; Song & Chun, 2005; Richter et al., 2010; Lott & Guez, 2013a) 62 and fronts in mid-latitudes (Charron & Manzini, 2002; Richter et al., 2010; De la Cámara 63 & Lott, 2015). These non-orographic GW parameterizations reduce biases in the mid-64 dle atmosphere (Scinocca, 2003; Song & Chun, 2005; Beres et al., 2005; Orr et al., 2010; 65

Lott & Guez, 2013a; De la Cámara & Lott, 2015; Bushell et al., 2015; Anstey et al., 2016; Serva et al., 2018), which is an indirect proof of their realism.

With the most recent global satellite observations, it becomes possible to gain a 68 global view of the GW fields. However, these only detect large horizontal- and/or vertical-69 scale GWs and mostly measure temperature fluctuations (Geller et al., 2013a): the mo-70 mentum fluxes are computed indirectly using polarization relations (Alexander et al., 71 2010; Ern et al., 2014). In order to observe the shorter horizontal scales and to have more 72 direct access to the momentum fluxes, in situ observations are provided by constant-level 73 74 long-duration balloons like those made in the Antarctic region during Strateole, Vorcore (Hertzog, 2007) and Concordiasi (Rabier et al., 2010), or in the deep tropics during Pre-75 Concordiasi (Jewtoukoff et al., 2013) and Strateole 2 (Haase et al., 2018). These obser-76 vations have helped improve parameterizations of GWs (Alexander et al., 2021; Lott et 77 al., 2023) but remain regional, and limited to the lower stratosphere (Geller et al., 2013b; 78 Achatz et al., 2024). 79

Global high-resolution models offer a promising avenue to supplement these lim-80 itations. First, in these models, the GW fields start to look rather realistic compared to 81 satellite observations (Kruse et al., 2022; Gupta, Reichert, et al., 2024), with the limit 82 that the GW dynamics in these models remains dependent on the model formulation (Wedi 83 et al., 2020; Stephan et al., 2019). Secondly, these models start to reveal the resolution 84 at which GWs no longer need to be parameterized, which is around 1 km or even finer 85 (Polichtchouk et al., 2023; Gupta, Reichert, et al., 2024). If true, it means that param-86 eterizations will remain necessary in most climate models for the foreseeable future (Achatz 87 et al., 2024). Third, high-resolution models provide valuable recommendations for im-88 proving GW parameterizations. For example, Kruse et al. (2022) and Gupta, Reichert, 89 et al. (2024) highlight the necessity of including lateral propagation, and Polichtchouk 90 et al. (2023) highlight the need to reconsider the partitioning between the different sources. 91 For instance, it is essential to achieve the correct balance between orographic, frontal and 92 convective GWs to simulate well the Northern Hemisphere (NH), the Southern Hemi-93 sphere (SH), and the quasi-biennial oscillation in the tropics (de la Cámara et al., 2016). 94

Based on the success of the GW parameterization schemes to simulate a realistic 95 climate, the purpose of this paper is to evaluate momentum fluxes parameterized in cli-96 mate models by comparing them to those explicitly resolved in a recent state-of-the-art 97 high-resolution model. This comparison is an essential step in bridging the gap between 98 observations, models, and parameterizations and has practical applications. One is to 99 determine if high-resolution models produce the right amount of waves, and another is 100 to assess if the GW dynamics at work in parameterizations are consistent with the GW 101 dynamics in high-resolution models. To reach these goals, we first use a simulation per-102 formed with the ICOsahedral Nonhydrostatic Weather and Climate Model (ICON) at 103 a resolution around 5 km. The fields from this model are "coarse-grained" at horizon-104 tal scales near 1° to mimic the grid scale of a climate model. Then, subgrid-scale fluxes 105 in the full-resolution model are defined as those with scales less than the coarse-grained 106 resolution. We second run the GW parameterizations that are operational in the Insti-107 tut Pierre-Simon Laplace Coupled Model version for CMIP6 (IPSLCM6, Boucher et al. 108 (2020)) in offline mode using the ICON coarse-grained fields of winds, temperature, and 109 precipitation. 110

The plan of the paper is as follows: Data and methods are discussed in section §2. In section §3, the "subgrid-scale" GWs drag and fluxes resolved by ICON are compared to the parameterized GWs drag and fluxes across different sectors and altitudes. Section §4 discusses how such a comparison can help improve parameterizations of GWs and/or help estimate whether a high-resolution model produces the right amount of GWs.

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¹¹⁶ 2 Data and methods

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2.1 Resolved momentum flux

To calculate momentum fluxes due to GWs with horizontal scales smaller than cli-118 mate models' grid scales, we take 3-hourly instantaneous output from a 5 year ICON run 119 with atmosphere-ocean coupling performed within cycle 3 of the nextGEMS project (Koldunov 120 et al., 2023). We choose ICON because at the resolution of climate models it reasonably 121 simulates the middle atmosphere (Borchert et al., 2019). At the resolution of around 5 122 km used here, and without deep convection and GW parameterizations, it simulates pre-123 cipitation realistically (Stevens et al., 2019), it is capable of simulating GWs explicitly 124 (Hohenegger et al., 2023), and its GW fields show similarities when compared to balloon 125 and satellite observations (Köhler et al., 2023; Stephan et al., 2019). These comparisons 126 were carried out for the tropics, where the 5 km resolution is sufficient for an explicit sim-127 ulation of realistic deep convection (Stevens et al., 2019). The run was initialized from 128 ERA5 on January 20, 2020 and we analyze 3 weeks of data between March 20 and April 129 10, 2020. Selecting dates near equinox was motivated by the desire to have comparable 130 dynamical filtering conditions for GWs between the two hemispheres in the low strato-131 sphere (e.g., mid-latitude stratospheric jets with positive zonal mean zonal winds in both 132 hemispheres). The horizontal resolution of the icosahedral grid is R2B9 (~ 5 km), but 133 the outputs are written on a Hierarchical Equal Area isoLatitude Pixelisation (Healpix) 134 grid with resolution $N_{\rm side} = 1024$, corresponding to ~ 6.3 km. In the vertical, the model 135 has 90 unequally spaced levels up to z = 75 km. In the following, we only consider data 136 up to the top of the stratosphere, z = 50 km, below the sponge layer. 137

To extract GWs, we apply to the horizontal wind field a Helmholtz decomposition between a divergent and a rotational part (Lindborg, 2015), and assume that only the divergent part is associated with GWs (Stephan et al., 2022; Gupta, Sheshadri, & Anantharaj, 2024). We use it to compute cross products, such as uw, with u and w being the zonal and vertical winds, respectively. We then average the original wind fields and cross products to a coarser Healpix grid with a resolution of $N_{\text{side}} = 64$, corresponding to ~100 km. This mapping allows the calculation of the momentum flux,

$$F^{z} = \rho \left[u'w' \right] + \rho f \frac{\left[v'\theta' \right]}{\left[\theta \right]_{z}},\tag{1}$$

where [] is the average over the $N_{\text{side}} = 64$ Healpix coarse-grained grid boxes and $u^{'} = u - [u]$. In addition, ρ is the density, θ is the potential temperature, v is the meridional velocity and f is the Coriolis parameter. In (1), we found that the contribution of the thermal flux $[v'\theta']$ to F^z was always negligible compared to that of [u'w']. To quantify the contribution of F^z to the general circulation, we also evaluate its zonal mean (indicated by overlines) and refer to

$$\overline{F}^{z} = a\cos\phi\left(-\overline{\rho\left[u'w'\right]} + \overline{\rho f\frac{\left[v'\theta'\right]}{\left[\theta\right]_{z}}}\right),\tag{2}$$

as the contribution of the subgrid-scale waves to the vertical component of the EP flux, with *a* being Earth's radius and ϕ latitude. For completeness, note that when using ICON data, we always include the meridional component of the Eliassen-Palm flux,

$$\overline{F}^{\phi} = -a\cos\phi\,\overline{\rho\,[u'v']},\tag{3}$$

to calculate the zonal mean GW drag, i.e.

$$\frac{1}{\rho a \cos \phi} \left(\frac{\partial \cos \phi \,\overline{F}^{\phi}}{\partial \phi} + \frac{\partial \overline{F}^z}{\partial z} \right). \tag{4}$$

Note that we found the contribution of \overline{F}^{ϕ} to be small, leaving $\overline{u'w'}$ making the largest

contribution to the drag and momentum fluxes. Note also that Procházková et al. (2023)

and Sun et al. (2023) found that $\overline{u'w'}$ is not much sensitive to the method for extracting GWs from high-resolution model data.

¹⁵⁹ 2.2 Parameterized momentum flux

The subgrid-scale orography parameterization we use is described in Lott and Miller 160 (1997), the subgrid scale fields being calculated on the IPSLCM6 grid. The non-orographic 161 GW schemes, which include as sources convection and fronts, are described in Lott and 162 Guez (2013b) and De la Cámara and Lott (2015), respectively. None of these schemes 163 include lateral propagation or rotation; they assume $F^y = 0$ and parameterize F^z ne-164 glecting the thermal flux. Note that we performed no tuning and kept the setup of IP-165 SLCM6 for a 143×142 lat-lon grid. The parameterizations are then run offline from 166 March 20 to April 10, 2020, using the coarse-grained three-dimensional fields of ICON 167 with calculations performed on the coarse-grained grid. Note that Lott et al. (2023) dis-168 cuss to what extent such an offline approach gives useful insight into what occurs on-169 line. 170



Figure 1. Zonal mean 21-day mean of the resolved (a) and parameterized (b) GW drag. Solid and dashed contour lines represent eastward and westward zonal mean 21-day mean zonal wind, respectively. The gray lines indicate the three altitudes z = 8, 16, 23 km at which we analyse the momentum fluxes and the latitudes $\phi = \pm 15^{\circ}$, $\phi = \pm 45^{\circ}$, we use to dynamically separate the tropics from the subtropics, and the subtropics from the mid-latitudes and polar regions.

Fig. 1 compares the GW drag from the resolved and parameterized waves. In the 172 stratosphere, Fig. 1(a) shows that the resolved waves produce negative drag where the 173 winds are positive with negative vertical shears for instance below the stratopause in the 174 NH on the upper flank of the stratospheric jet (30 km < z < 50 km, $\phi > 45^{\circ}$), and in 175 the subtropical regions above the tropospheric jets in both hemispheres (13 km < z <176 20 km, $15^{\circ} < |\phi| < 35^{\circ}$). There are also negative drags in the SH stratospheric jet, 177 again in locations of positive zonal wind (below $z \approx 50$ km, $\phi < -45^{\circ}$). In the tropi-178 cal and subtropical upper stratosphere (z > 30 km, $-45^{\circ} < \phi < 45^{\circ}$), the drag is posi-179

171 **3 Results**

tive, mostly in locations of negative or low wind speeds with positive shear. Interestingly, 180 the parameterizations in Fig. 1(b) reproduce comparable patterns concerning the sign 181 of the drag, at least in the upper stratosphere. Regarding the amplitudes, resolved and 182 parameterized drags in the upper stratosphere compare well above the NH stratospheric 183 jet core, but the parameterizations underestimate the positive drag in the tropics and 184 overestimate the negative drag in the SH stratospheric jet. Another pronounced differ-185 ence is that above the tropospheric subtropical jet ($z \approx 16$ km, $15^{\circ} < |\phi| < 45^{\circ}$), the 186 parameterizations strongly underestimate the negative drags. 187

It is encouraging that the signs of parameterized drags are largely consistent with those due to resolved waves. It implies that the resolved wave dynamics is well captured by parameterizations representing upward-propagating GWs with origins in the troposphere. The parameterizations are formulated such that negative (positive) intrinsic phase speed waves yield negative (positive) momentum fluxes and break more easily in negative (positive) zonal winds. In the following, we refer to this mechanism as "dynamical filtering".



Figure 2. Global distribution of the 21-day mean net zonal momentum fluxes resolved by ICON (left), predicted by the parameterization schemes using ICON meteorological fields (middle), and their difference (right) at (a,b,c) z = 8 km, (d,e,f) z = 16.4 km, and (g,h,i) z = 23.9 km. The correlations between the resolved and parameterized momentum fluxes are presented in the adjacent column.

To test if the resolved and parameterized waves entering the stratosphere originate from the same regions, Fig. 2 shows the geographic distribution of the 21-day averaged momentum fluxes, F^z , at z = 8 km (a,b), 16 km (c,d), and 23 km (e,f). We choose these three levels because we found more substantial differences in GWs absorptions in the lower stratosphere than above (see S2). To better evaluate the geographic correspondences, the differences between the resolved and parameterized momentum fluxes are shown in a separate column (Fig. 2(c,f,i)) and the correlations,

$$C(\phi) = \overline{(F_I^z - \overline{F}_I^z)(F_P^z - \overline{F}_P^z)} / \sqrt{(F_I^z - \overline{F}_I^z)^2} \overline{(F_P^z - \overline{F}_P^z)^2},$$
(5)

are shown as a function of latitude in the adjacent column. In (5), I and P indices denote the resolved and parameterized momentum fluxes, respectively.

In the troposphere at z = 8 km in Figs. 2(a,b,c), there are resemblances over the 204 major mountain ranges (Rockies, Andes, and from the Alps to the Tibetan plateau), with 205 the resolved fluxes larger in amplitude than parameterized ones. Over the Rockies, Hi-206 malayas, and Tibetan plateau, the difference map in Fig. 2(c) shows changes in sign above 207 the same mountain range. This qualitative difference is certainly due to the fact that the 208 orographic GW scheme does not include lateral propagation (Kruse et al., 2022). In other 209 210 regions, the patterns agree to a varying extent. Over the oceans and outside the tropics, the resolved and parameterized fluxes have the same sign, the resolved fluxes being 211 substantially larger in the region including the subtropics $(15^{\circ} < |\phi| < 45^{\circ})$ and quite 212 comparable more poleward $(|\phi| > 45^{\circ})$. In the tropics, the differences are much more 213 pronounced. The resolved fluxes are much larger and often of opposite sign compared 214 to the parameterized ones and the correlations are small. In these regions, we recover 215 the results in Wei et al. (2022) or Köhler et al. (2023), who have shown that when there 216 is deep convection, the circulations that are diabatically forced produce momentum fluxes 217 that cannot be explained in terms of upward propagating GWs only. 218

In the lower stratosphere at z = 16 km in Figs. 2(d,e,f), the momentum fluxes show 219 better agreement. They are still quite large and predominantly negative over the moun-220 tain ranges in the mid-latitudes and over the oceans away from the tropics. Over the oceans, 221 there is a global shift of the parameterized fluxes toward the polar regions. In the NH, 222 for instance, and over the Atlantic Ocean east of Newfoundland along the storm track, 223 the parameterizations predict intense and negative fluxes whereas the resolved fluxes are 224 small. In the SH, the resolved fluxes have a band of maximum strength around $\phi = -30^{\circ}$ 225 (Fig. 2(d)), whereas the parameterized fluxes have a comparable band but centered 10° 226 poleward (Fig. 2(e)). In the tropics ($|\phi| < 15^{\circ}$), the pattern and signs of the fluxes agree 227 well, with positive correlation, although the parameterized fluxes tend to be slightly weaker 228 in amplitude. More specifically, in the eastern Pacific, resolved and parameterized fluxes 229 are strong and negative above the two branches of the Intertropical Convergence Zone 230 (ITCZ), and positive over Brazil, equatorial Africa, the Indian Ocean and the Maritime 231 Continent. It appears that at this altitude and above convective regions, upward prop-232 agating GW dynamics are better able to explain momentum fluxes compared to the tro-233 posphere. 234

Whereas the parameterized and resolved fluxes show resemblance at z = 16 km, 235 the agreement deteriorates at z = 23 km in Figs. 2(g,h,i). The major differences are 236 above the subtropical jets $(15^{\circ} < |\phi| < 45^{\circ})$, where the resolved fluxes are positive 237 and the parameterized ones negative. In these regions, Fig. 2(g) indicates that the re-238 solved westward waves contributing to the negative fluxes in Fig. 2(d) have been con-239 siderably attenuated, whereas the parameterized westward waves continue to dominate. 240 Accordingly, the correlations between the fluxes become small. In the mid-latitudes and 241 polar regions ($|\phi| > 45^{\circ}$) over oceans and land, resolved and parameterized fluxes are 242 consistently positive and mostly correlated, but the resolved fluxes are substantially smaller 243 in amplitude. 244

To shed light on the causes of similarities and differences between the resolved and 245 parameterized fluxes, Fig. 3 shows the contribution of the three GW parameterizations 246 at z = 16 km and z = 23 km. Over mountainous regions, the orographic GW param-247 eterization is responsible for the major fraction of the negative fluxes. They are substan-248 tially attenuated between 16 km and 23 km in the subtropical bands $(15^{\circ} < |\phi| < 45^{\circ})$ 249 250 (above the subtropical tropospheric jet) when the zonal wind in the stratosphere becomes small or negative (Fig. 1). For these waves of zero absolute phase speed, this attenua-251 tion is caused by "near"-critical level situations and a comparable attenuation is seen 252 in the resolved fluxes in Figs. 2(d) and 2(g). 253



Figure 3. Global distribution of the 21-day mean net zonal parameterized momentum fluxes using ICON meteorological fields: orographic GWs (left), convective GWs (middle), and frontal GWs (right), at z = 16 km (first row) and z = 23 km (second row).

The frontal GW parameterization (Figs. 3(b) and (e)) produces large negative fluxes 254 from the subtropics to the polar regions ($|\phi| > 15^{\circ}$), and is responsible for the exces-255 sively large parameterized fluxes at z = 16 km in the mid-latitudes (Fig. 2(e)). These 256 westward fluxes are not much attenuated between 16 km and 23 km because in the mid-257 latitudes and polar regions $(|\phi| > 45^{\circ})$, the stratospheric wind is positive in both hemi-258 spheres (see Fig. 1). In the subtropics $(15^{\circ} < |\phi| < 45^{\circ})$, the parameterized frontal 259 waves are not much attenuated between z = 16 km and z = 23 km, presumably be-260 cause they have too large phase speeds. The facts that the resolved westward waves (i) 261 have smaller amplitudes in the mid-latitudes and (ii) are more filtered in the subtrop-262 ics explain most of the positive differences between the resolved and parameterized mo-263 mentum fluxes in the subtropics and mid-latitudes seen at 23 km in Fig. 2(i) where $|\phi| >$ 264 15^{o} . 265

The convective GW parameterization in Figs 3(c) and 3(f) explains most of the pa-266 rameterized fluxes in the tropics and subtropics, for instance over the Pacific ITCZ, the 267 Maritime Continent and the monsoon regions (Figs. 2(d) and 2(e), respectively). Around 268 the subtropical zones at $\phi = \pm 30^{\circ}$, the parameterized convective GW flux is consid-269 erably attenuated and changes sign between z = 16 km and z = 23 km. Such a change 270 of sign is also present in the resolved fluxes around these latitudes (Figs. 2(d) and 2(g)), 271 although much more pronounced. Again and like for orographic waves, this attenuation 272 of the negative intrinsic phase speed waves occurs aloft the tropospheric subtropical jet 273 center, consistent with dynamical filtering. Although the convective GW fluxes have smaller 274 amplitudes than the resolved fluxes, their change in sign between z = 16 km and z =275 23 km and their patterns at z = 23 km look qualitatively comparable in the tropics and 276 subtropics (for the resolved waves see Figs. 2(d) and 2(g), and the peak in correlation 277 compared to the subtropics). 278

Considering net fluxes hides the fact that negative and positive phase speed waves can lead to cancelling momentum fluxes but result in drags of opposite signs at different altitudes. With the approach of identifying momentum fluxes in ICON taken here, it is impossible to separate eastward and westward waves. As supporting information, we therefore propose to assume that the net fluxes of a given sign at a given time are associated with GWs of corresponding intrinsic phase speed sign. The method is not ideal,
 we can underestimate the fluxes when there are phase speed symmetries, but easy to implement.

Fig. S1 shows that the temporal averages of these sign definite fluxes have ampli-287 tudes comparable with the net fluxes in Fig. 2. It also shows that the resolved and parametrized 288 fluxes continue to be comparable. More specifically, away from the tropics $(|\phi| > 15^{\circ})$, 289 Figs. S1(a,b,e,f) show that at z = 16 km, the westward fluxes strongly dominate the 290 eastward fluxes as expected for mountain waves and also because the tropospheric zonal 291 winds are predominantly positive and favouring westward waves above. Comparing Figs. 292 S1(a) and S1(c) show that the resolved westward waves are strongly absorbed above the 293 subtropical jet $(15^{\circ} < |\phi| < 45^{\circ})$, and much more than the parameterized ones in S1(b) 294 and S1(d). In the tropics $(|\phi| < 15^{\circ})$, there is a strong longitudinal variations with the 295 westward fluxes dominating over the Pacific ITCZ at z = 16km (S1(c) and S1(d)), and 296 the eastward fluxes dominating elsewhere (S1(e) and S1(f)). The westward fluxes seem 297 more attenuated than the eastward fluxes between z = 16 km and z = 23 km, presum-298 ably because the wind shear is predominantly positive in the tropical lower stratosphere 299 (Fig. 1). 300

For a more quantitative depiction, Fig. S2(a,b,c,d) present their vertical profiles 301 averaged over the four sectors we have used so far to characterize the momentum flux 302 behaviour, i.e. mountainous regions, defined as standard deviation of subgrid-scale orog-303 raphy $\sqrt{[h'^2]} > 100$ m, and non-mountainous regions (land and ocean) in the mid-latitudes 304 $(|\phi| > 45^{\circ})$, the subtropics $(15^{\circ} < |\phi| < 45^{\circ})$, and the tropics $(|\phi| < 15^{\circ})$. In all pan-305 els, and as expected from the comparisons at the three levels analyzed before (z = 8, 16, 23306 km), the resolved fluxes (solid lines) are systematically larger than the parameterized 307 ones (dashed lines) in the upper troposphere but decay rapidly with altitude in the lower 308 stratosphere. The $z \simeq 16$ km level is where the fluxes almost intersect. More specifi-309 cally, the fluxes over the mountain regions are those for which there are the best corre-310 spondences, dominated by westward waves, and for the parameterized waves, the con-311 tribution of the subgrid-scale orographic waves make the larger contribution. In the mid-312 latitudes, the parameterized waves dominate in the lower stratosphere and are slightly 313 less attenuated than the resolved ones between z = 16 km and z = 23 km. In the sub-314 tropics between z = 16 km and z = 23 km, the westward resolved waves are strongly 315 attenuated. In the tropics, the resolved fluxes are greater at all levels. Aloft z = 23 km, 316 the decay rates of all the fluxes with altitude become more comparable. 317

318 4 Conclusion

This study uses high-resolution simulations of the ICON model to estimate the mo-319 mentum fluxes due to disturbances with horizontal scales shorter than ≈ 100 km, rep-320 resentative of the grid of a climate model. These momentum fluxes are then predicted 321 by the GW parameterization schemes used in the IPSLCM6 climate model, the input 322 fields for the parameterizations being the ICON fields on an appropriately coarse-grained 323 grid. A key finding is that the momentum fluxes and drag in the stratosphere are quite 324 well predicted in terms of amplitude and geographical distribution. When we split the 325 analysis between sectors and parameterizations, we found that the high-resolution model 326 and the parameterizations show the best agreement over mountainous regions, where the 327 Lott and Miller (1997) parameterization is most relevant. In the mid-latitudes, the pa-328 rameterizations overestimate the fluxes, which we attributed to the De la Cámara and 329 Lott (2015) parameterization of frontal waves, whereas in the tropics, the parameteri-330 zations underestimate the fluxes, which we attribute to the Lott and Guez (2013a) pa-331 rameterization of convective GWs. The non-orographic parameterizations also under-332 estimate the dynamical filtering of the westward propagating GWs above the subtrop-333 ical jet. 334

If we take the high-resolution model as the truth, these discrepancies could be cor-335 rected, for instance, by decreasing the tuning parameter that controls the drag ampli-336 tude in the frontal scheme and increasing the one that controls the convective GW am-337 plitude. To improve the dynamical filtering aloft the subtropical jet, we could decrease 338 the parameter that controls the intrinsic phase speeds of the non-orographic GWs. Pre-339 liminary offline tests show that these changes indeed improve some aspects of the com-340 parison. Improving the statistical distributions is more challenging since in parameter-341 izations the intermittency of the momentum fluxes is in good part related to the way we 342 relate the GWs with their sources (de la Cámara et al., 2014). In the frontal scheme, they 343 are directly related to the square of grid-scale vorticity and in the convective GW schemes 344 to the square of grid-scale precipitation. Reconsidering these relations certainly calls for 345 more advanced analysis, like ENKF parameter estimation of the launch momentum fluxes 346 to better determine what causes large momentum fluxes in the lower stratosphere (Tandeo 347 et al., 2015). In any case, such parameter estimations are limited by the fact that our 348 parameterizations neglect lateral propagation (Voelker et al., 2024). Taking it into ac-349 count is numerically costly and may require approaches based on machine learning (Matsuoka 350 et al., 2020; Espinosa et al., 2022). 351

The similarities and differences we find here are certainly model and parameter-352 ization dependent. This is a deliberate choice: the parameterization schemes used have 353 not been tuned prior to the comparison, and the middle-atmosphere climatology of the 354 ICON high-resolution simulations we use has not yet been thoroughly validated. Nev-355 ertheless, the stratospheric jets shown in Fig. 1 are quite strong compared to climatol-356 ogy, a quite systematic error during all seasons of the run analysed (not shown), suggest-357 ing that the resolved GW drag is too small. In the tropics, the QBO rapidly fades away 358 (at the initial time of the run, the zonal winds in the lower tropical stratosphere are much 359 stronger than shown in Fig. 1), which may also indicate improper GWs forcing. How-360 ever, note that Giorgetta et al. (2022) used a comparable version of ICON and found GWs 361 forcing in the QBO region that are quite reasonable. In this respect, our results can also 362 serve as guidance to judge wave forcing in high-resolution models. After all, the param-363 eterizations have been tuned to reproduce a correct mean climate in the middle atmo-364 sphere. For example, we have seen that the frontal waves parameterization predicts larger 365 momentum fluxes than ICON simulates, but it is possible that the resolved waves are 366 underestimated in ICON. Indeed, high-resolution simulations have not yet converged when 367 it comes to GWs in the extratropics (Polichtchouk et al., 2023; Gupta, Reichert, et al., 368 2024) and should not be considered true. Other parameterizations or numerical imple-369 mentations may also affect the GW impact (e.g., dissipation). At least for convective waves 370 there is substantial evidence that their amplitude in high-resolution models is very sen-371 sitive to model formulation (Stephan et al., 2022). Accordingly, and in all regions, the 372 GW drags predicted by the parameterizations become an important upper bound to val-373 idate the high-resolution models. Systematically tuning the GW parameterizations to 374 improve their fit with high-resolution simulations could be a fruitful endeavour to learn 375 about deficits in the parameterizations as well as deficits in resolved dynamics. 376

377 Open Research

All data and routines needed to run the parameterizations offline and to compare the results with the coarse-grained ICON fields can be downloaded from Zenodo (Toghraei, 2024).

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Supporting Information for "Can parameterizations reproduce the gravity waves momentum fluxes and drag simulated by a global high-resolution model?"

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- 1. Figure S1
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Introduction

This supporting information provides two additional figures to the main article. Figure S1 shows the global distribution of the "sign definite" positive and negative zonal momentum fluxes resolved by ICON and predicted by the parameterization schemes. Figure S2 compares vertical profiles of east and west momentum fluxes resolved by ICON and predicted by the parameterization schemes.



Figure S1. Global distribution of the 21-day mean of the "sign definite" positive (eastward) and negative (westward) zonal momentum fluxes resolved by ICON (left) and predicted by the parameterization schemes (right). The eastward fluxes are in the top four panels, the westward ones in the bottom four panels. The altitude is z = 16 km in (a,b,e,f) and z = 23 km in (c,d,g,h).

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Figure S2. Vertical profile of the "sign definite" eastward (red) and westward (blue) zonal momentum fluxes resolved by ICON (solid lines) and predicted by the parameterization schemes (dashed lines) in 4 different sectors we have used so far to characterize the momentum flux behaviour, i.e. mountainous regions, defined as standard deviation of subgrid-scale orography $\sqrt{[h'^2]} > 100$ m, and non-mountainous regions (land and ocean) in the mid-latitudes ($|\phi| > 45^{\circ}$), the subtropics ($15^{\circ} < |\phi| < 45^{\circ}$), and the tropics ($|\phi| < 15^{\circ}$).Over mountain regions, the fluxes from the orographic GW parameterization scheme are shown with a dashed-dotted line.