

Energy Environment: Science Technology and Management (STEEM)  
Greenhouse gases: Challenges and observations

## Numerical modelling of recent and futur climate changes

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## Outlook

- I. Numerical climate models
- II. Simulating and attributing recent climate changes
- III. Scenarios and future climate changes
- IV. Climate sensitivity and climate feedbacks
- V. Climate variability and past climate changes

## Emergence of the physics of climate

### J. Fourier:

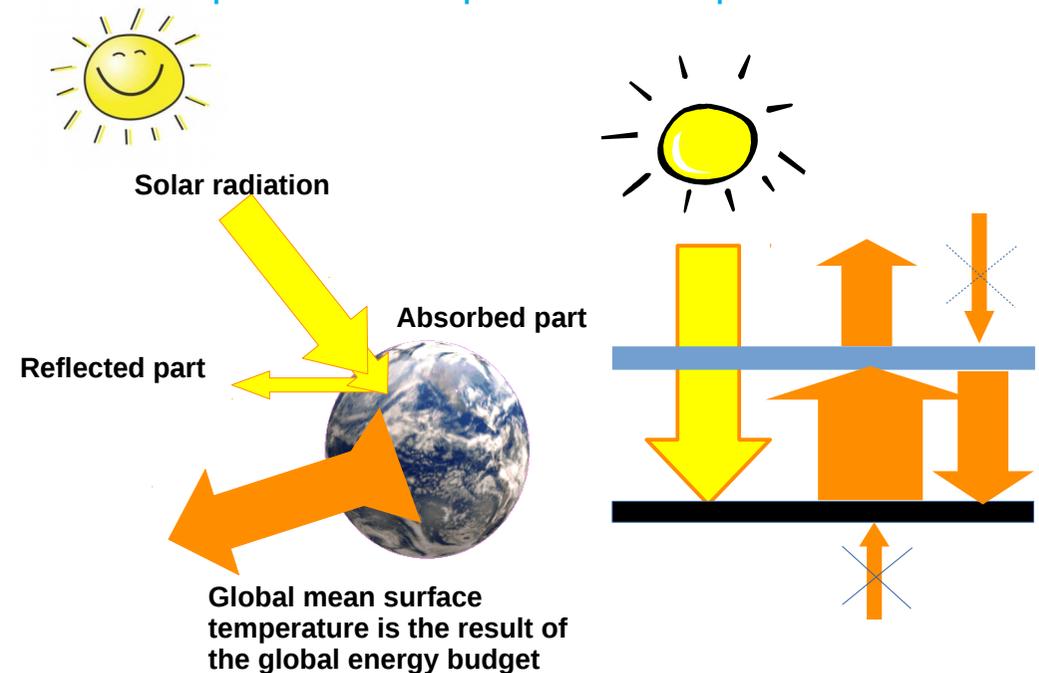
- *Mémoire sur les températures du globe terrestre et des espaces planétaires*, Mémoires de l'Académie des Sciences de l'Institut de France, 1824
- *General remarks on the Temperature of the Terrestrial Globe and the Planetary Spaces*; American Journal of Science, Vol. 32, N°1, 1837.



Joseph Fourier  
(1768-1830)

- He consider the Earth like any other planet
- The energy balance equation drives the temperature of all the planets
- The major heat transfers are
  1. Solar radiation
  2. Infra-red radiation
  3. Diffusion with the interior of Earth

## Equilibrium temperature of a planet



## Emergence of the physics of climate

### J. Fourier:

- *Mémoire sur les températures du globe terrestre et des espaces planétaires*, Mémoires de l'Académie des Sciences de l'Institut de France, 1824
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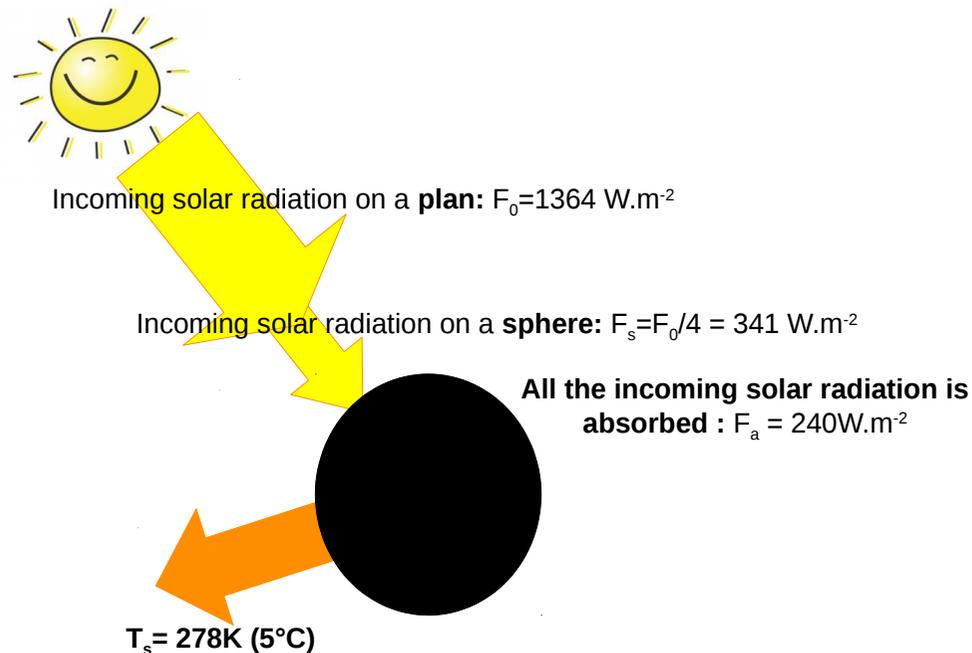
**Joseph Fourier**  
(1768-1830)

➤ He **envisages the importance of any change of the sun** « The least variation in the distance of that body [the sun] from the earth would occasion very considerable changes of temperature. »

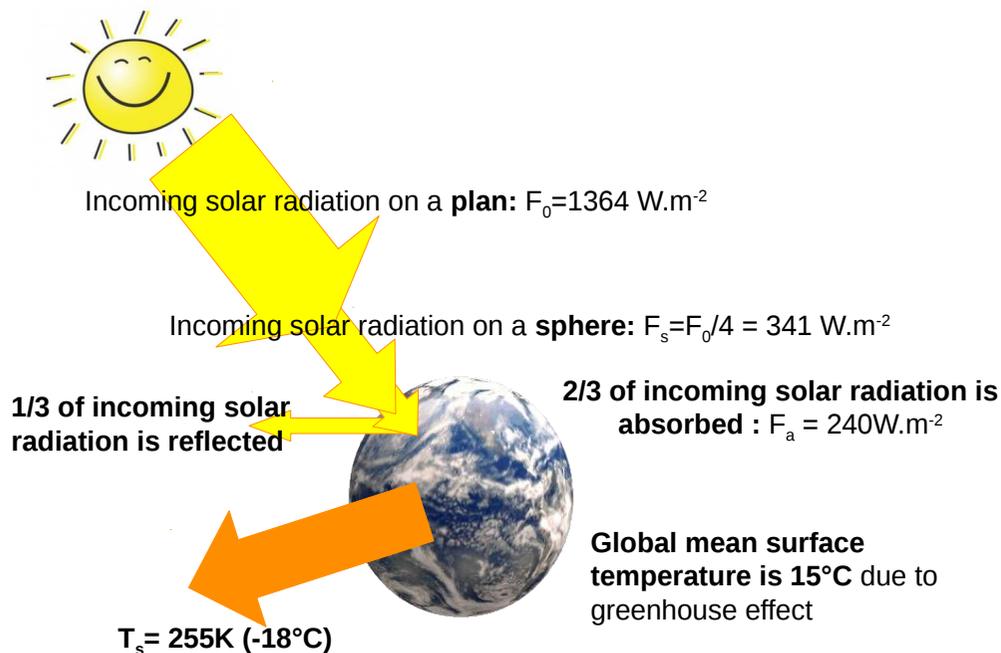
➤ He **envisages that climate may change**: « The establishment and progress of human society, and the action of natural powers, may, in extensive regions, produce remarkable changes in the state of the surface, the distribution of waters, and the great movements of the air. Such effects, in the course of some centuries, must produce variations in the mean temperature for such places ».

[Dufresne, 2006]

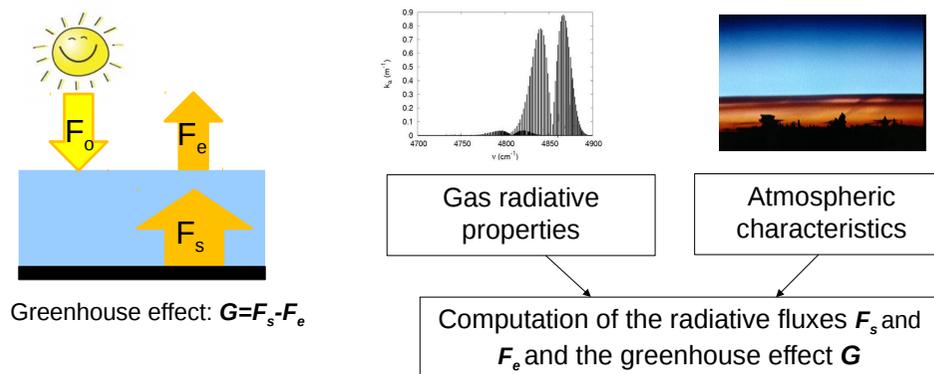
## Equilibrium temperature of a planet



## Equilibrium temperature of a planet

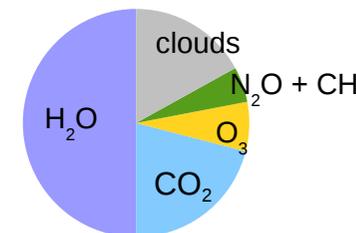


## What radiation heat transfer theory tell us

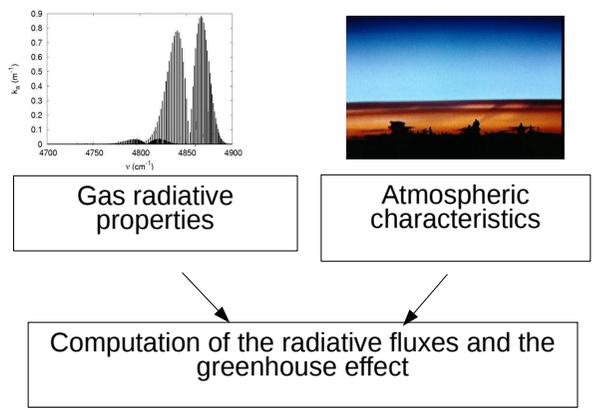


### Current greenhouse effect and the various contributions

	(W.m <sup>-2</sup> )	(%)
Total	150	
Water vapour	75	50
CO <sub>2</sub>	32	21
ozone	10	7
N <sub>2</sub> O+CH <sub>4</sub>	8	5
Clouds	25	17



# What radiation heat transfer theory tell us



## For a doubling of the CO<sub>2</sub> concentration:

- The green house effect increases by  $\approx 3.7 \text{ W.m}^{-2}$
- The temperature increases by  $\approx 1.2 \text{ K}$ , if nothing changes except the temperature

# From radiative transfer computation to climate modelling

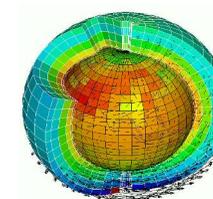
For a doubling of the CO<sub>2</sub> concentration:

- the green house effect increases by  $3.7 \text{ W.m}^{-2}$
- the temperature increases by  $\approx 1.2 \text{ K}$ , if nothing change except the temperature

## But feedbacks exist:

- Snow and sea ice reflect solar radiation; if they decrease, more solar energy will be absorbed  $\Rightarrow$  positive feedback
- Water vapour is the main greenhouse gas; if it increases, the greenhouse effect will be enhanced  $\Rightarrow$  positive feedback
- Clouds reflect solar radiation and contribute to the greenhouse effect; if they change, the energy budget will be modified  $\Rightarrow$  positive or negative feedback

## Need of 3D numerical climate models



# Numerical climate models (numerical weather simulators)



**Wilhelm Bjerknes**  
(1862-1951)



**L. F. Richardson**  
(1881-1953)



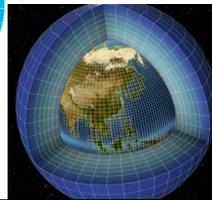
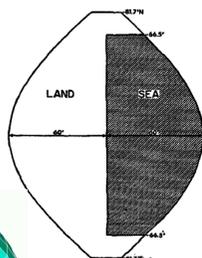
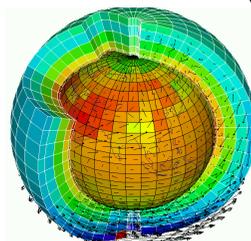
**J. von Neumann**  
(1903-1957)



**Jule Charney**  
(1917-1981)

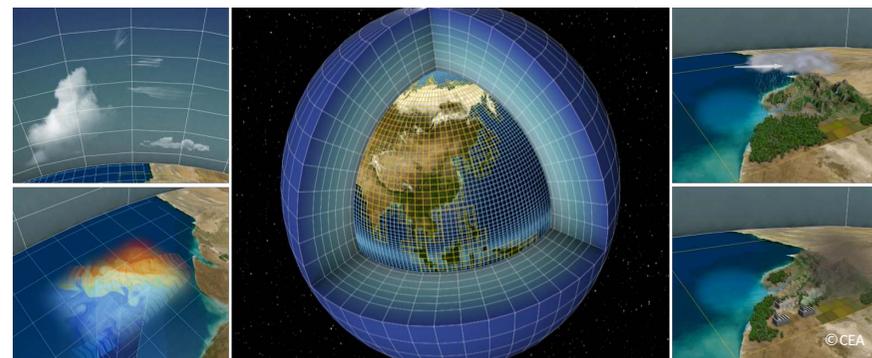


**Syukuro Manabe**  
(1931-)

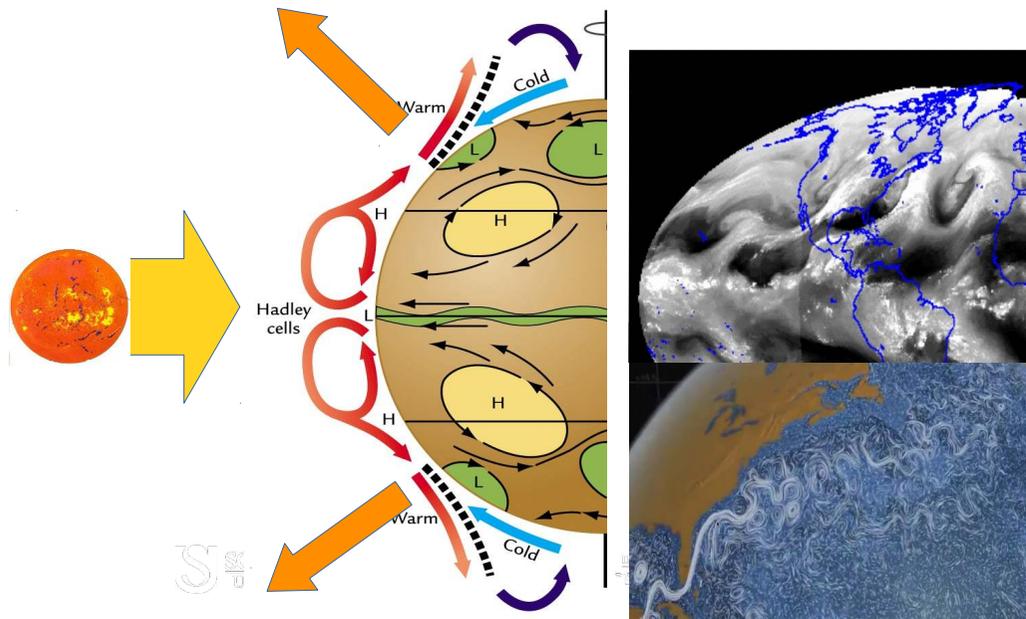


# 3D climate numerical models

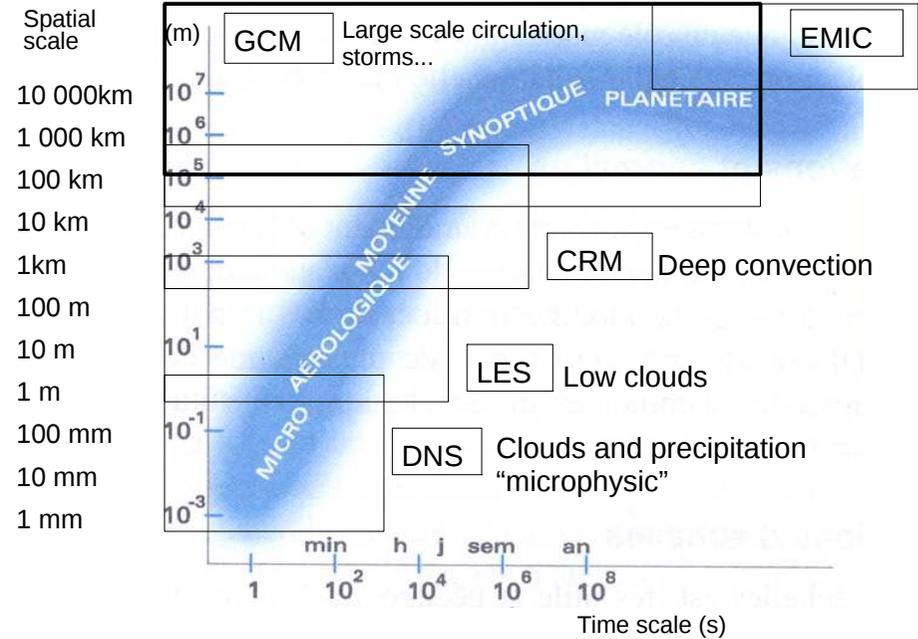
- Physical laws (atmosphere, ocean, sea-ice...)
- Discretization (temporally and spatially)
- Modelling of the sub-grid phenomena, or parameterization



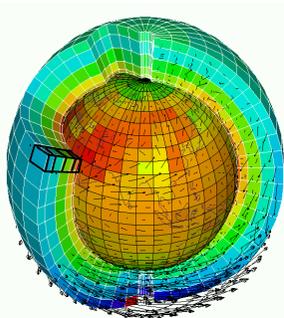
# Large scale circulation: Meridional heat transport and the effects of Earth rotation



# Relevant spatial and time scales



# General circulation models (GCMs)



**Dynamical core** : discretized version of the equations of fluid mechanics

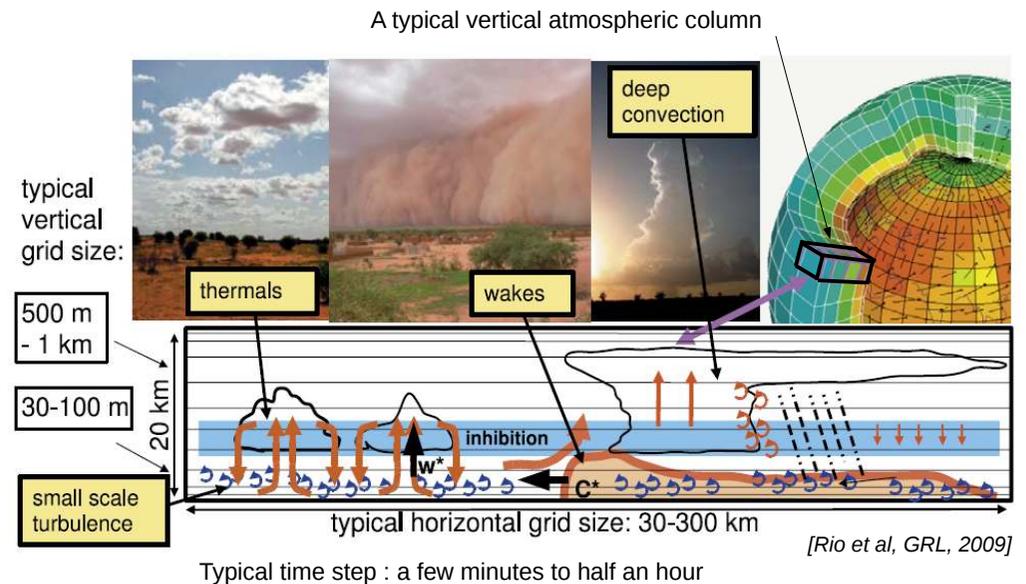
- Mass Conservation  
 $D\rho/Dt + \rho \text{div}\underline{U} = 0$
- Energy Conservation  
 $D\theta/Dt = Q/C_p (\rho_0/p)^\kappa$
- Momentum Conservation  
 $D\underline{U}/Dt + (1/\rho) \text{grad}p - g + 2 \underline{\Omega} \wedge \underline{U} = \underline{E}$
- Conservation of Water (and other species)  
 $Dq/Dt = S_q$

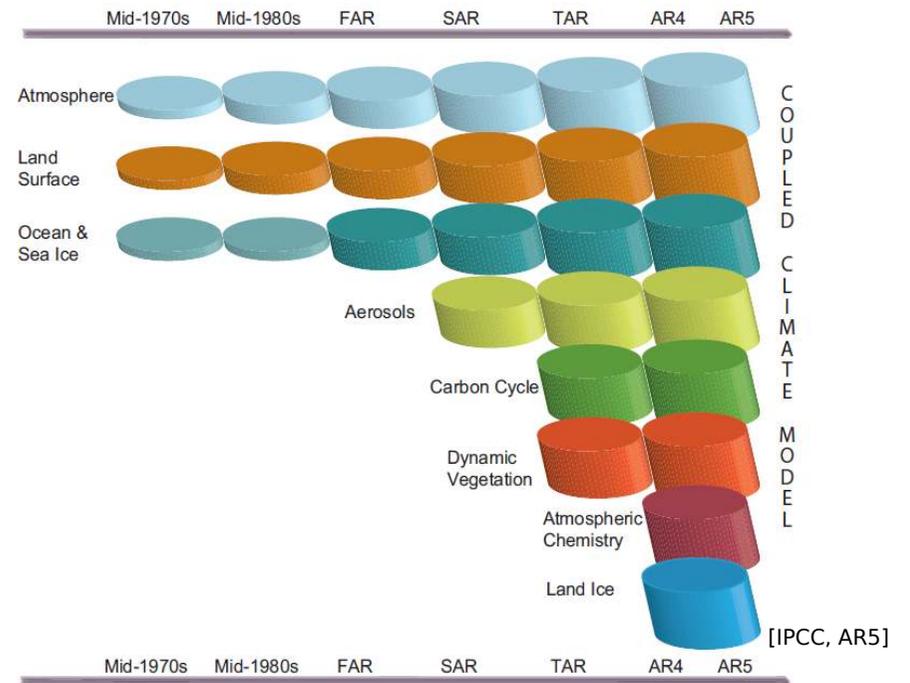
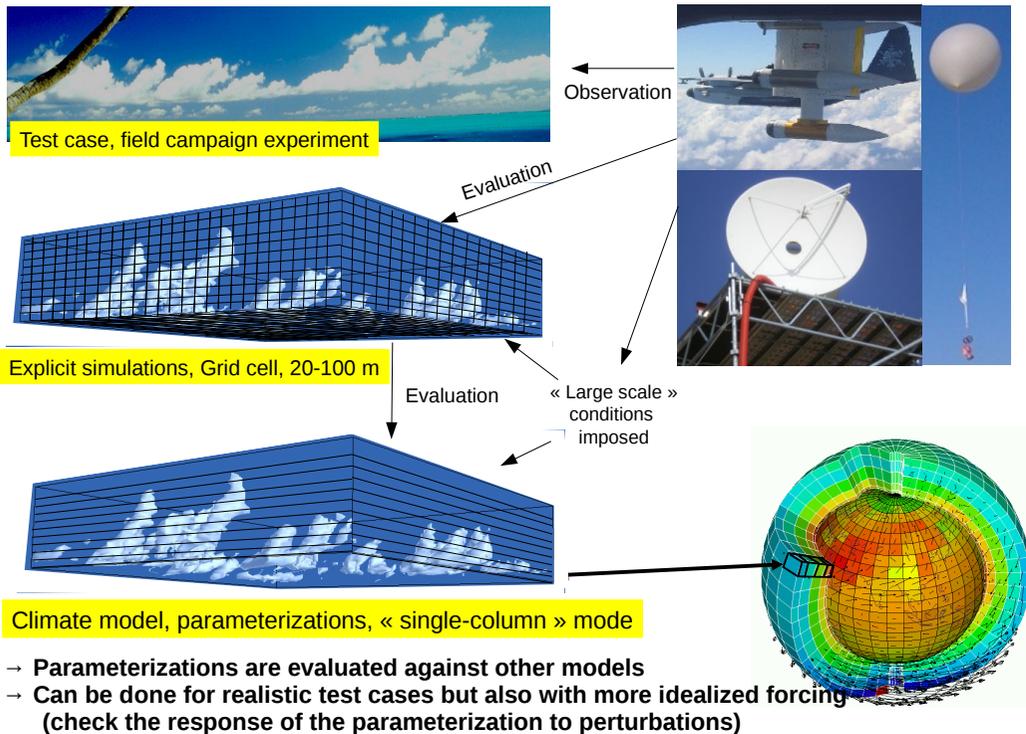
In red, source terms : other than fluid mechanics and unresolved scales

## General Circulation Models

- Developed in the 60s for the purpose of weather forecast
- Based on a discretized version of the « primitive equations of meteorology »
- On the Earth but also very rapidly on other planets
- A number of important process are subgrid scale and must be parameterized

# Modeling of unresolved scales Development of parameterization



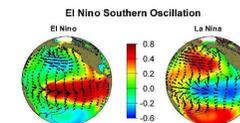


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## What drives climate variations and changes ?

Internal variability



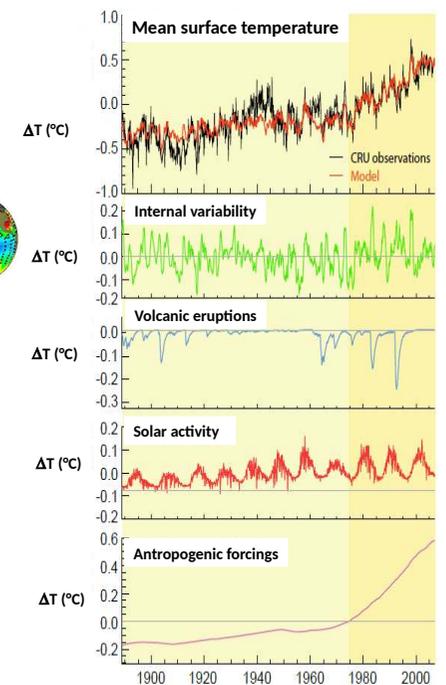
Volcanic activity



Solar activity



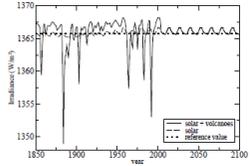
Anthropogenic factors



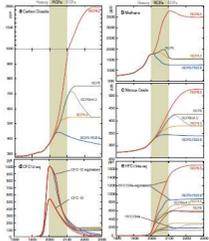
# The IPSL Earth System Model

## Natural and anthropogenic forcings

### Solar and volcanoes

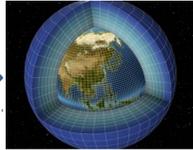


### Green house gases and active gases

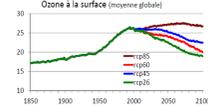


### CO<sub>2</sub> concentration

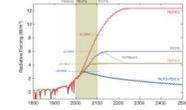
IPSL-CM5A-LR



## Atmospheric composition

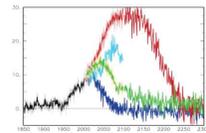


## Radiative forcings

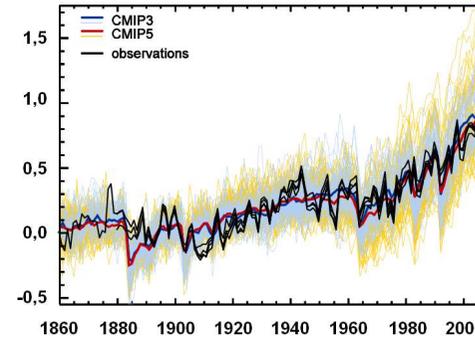


## Climate changes

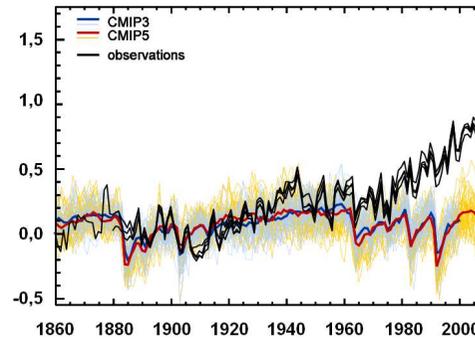
## Authorized CO<sub>2</sub> emissions



# Human activities and recent climate change



Mean surface temperature anomalies **observed** and **computed** considering *both the natural and anthropogenic perturbations* (observed volcanic eruptions and changes of the solar activity, the green house gases, the aerosols...)

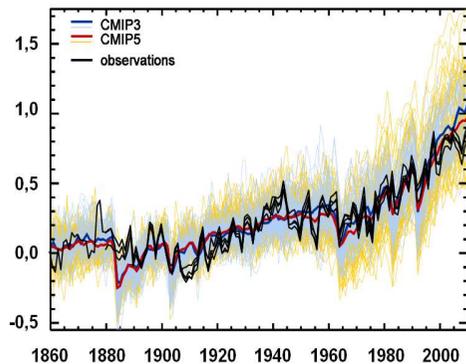


Mean surface temperature anomalies **observed** and **computed** considering *only the natural perturbations* (observed volcanic eruptions and changes of the solar activity...)

(IPCC, 2013)

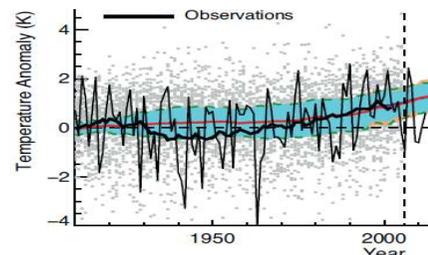
# Surface temperature evolution: observation and models

## Annual global mean

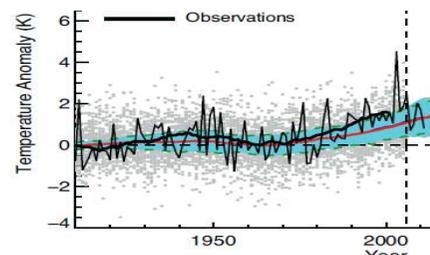


[IPCC, 2013]

## Winter mean over France

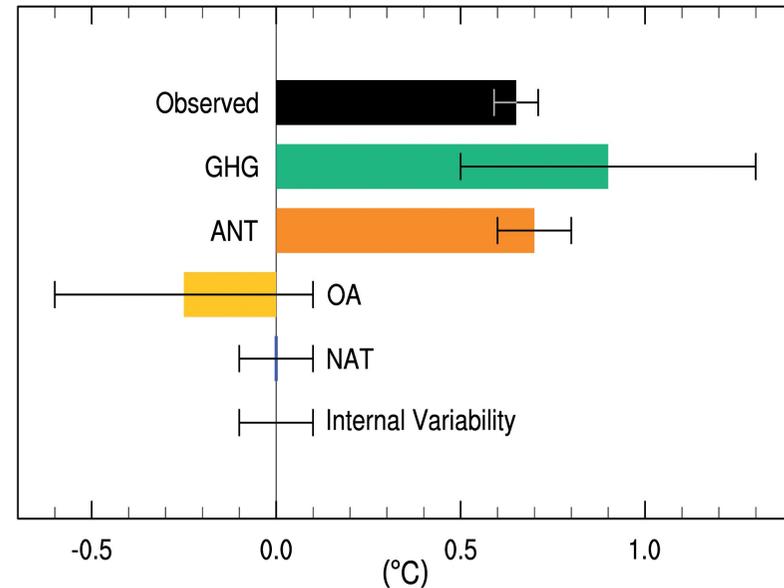


## Summer mean over France



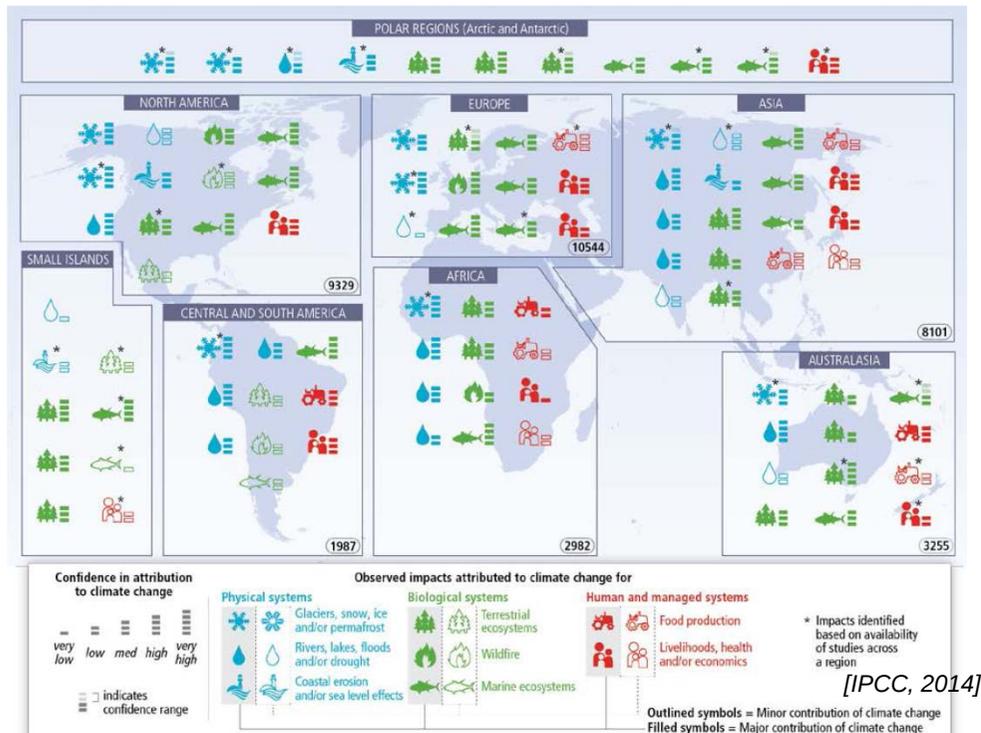
[Terry et Boé, 2013]

# Origin of the global warming trends over the 1951-2010 period



[IPCC, 2013]

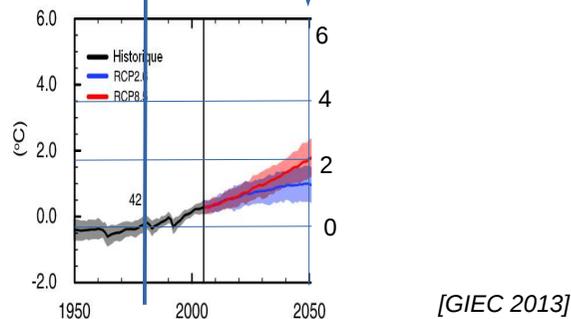
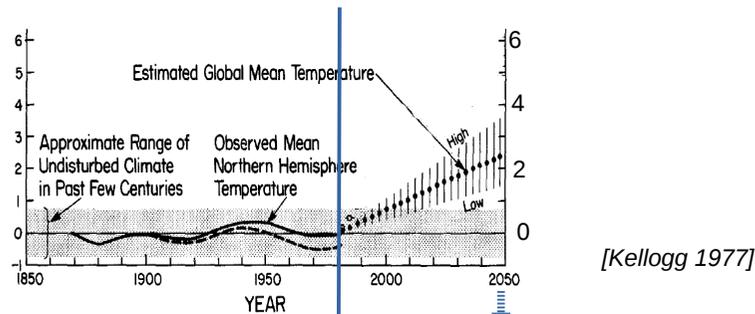
## Observed impact attribute to climate change



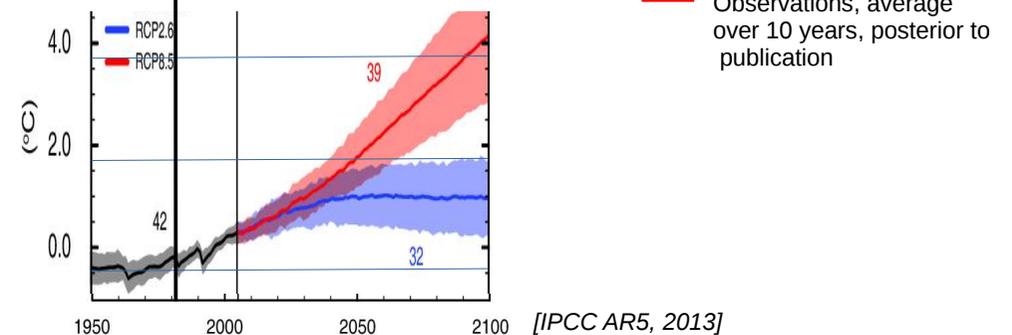
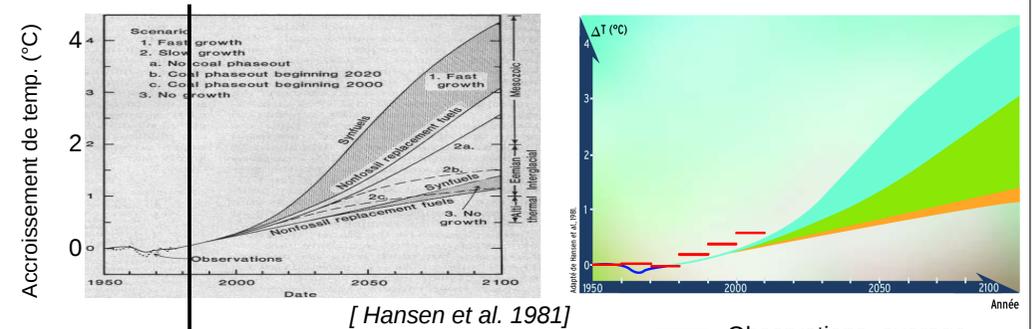
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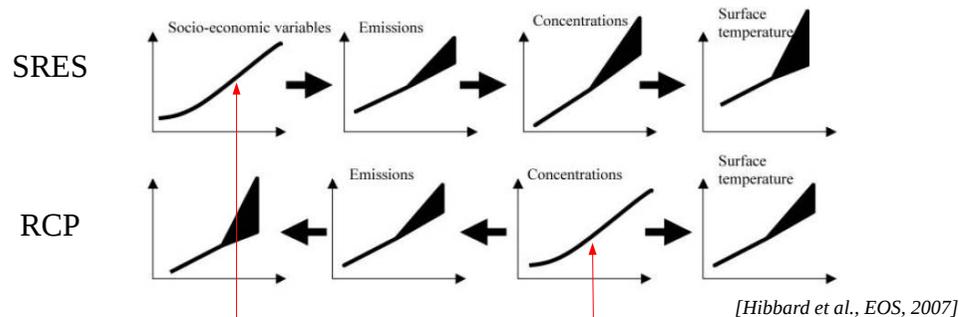
## First climate projections before global warming has been observed



## First climate projections before global warming has been observed



## Scenarios for future climate change projections



Scenario as the starting point of the full chain

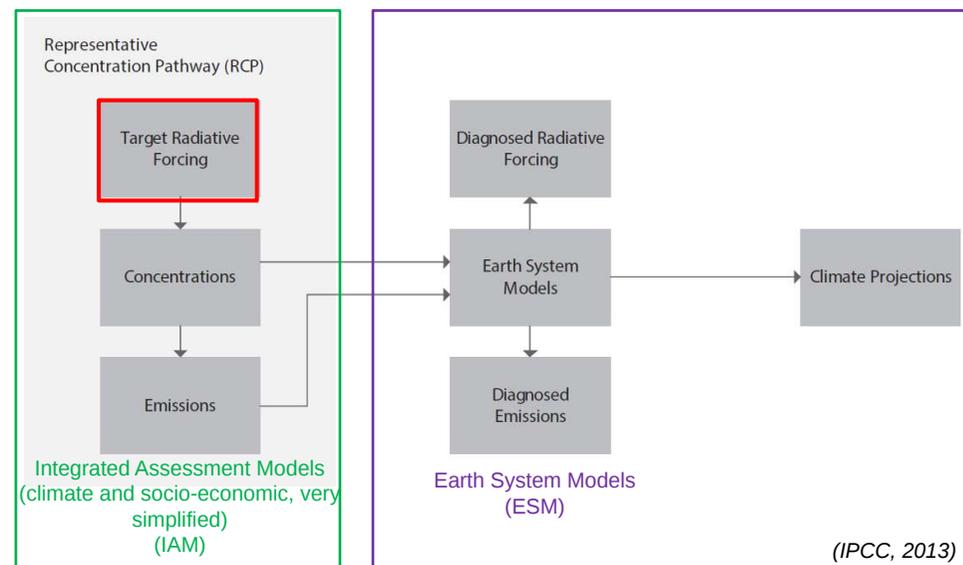
Scenario as the gathering place between socio-economic and climate sciences

[Hibbard et al., EOS, 2007]

## Scenarios for future climate change projections

Scenarios are defined with a target value of the **radiative forcing ( $W.m^{-2}$ )**

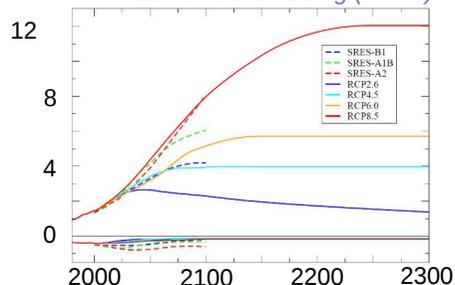
**RCP : « Representative concentration pathways »**



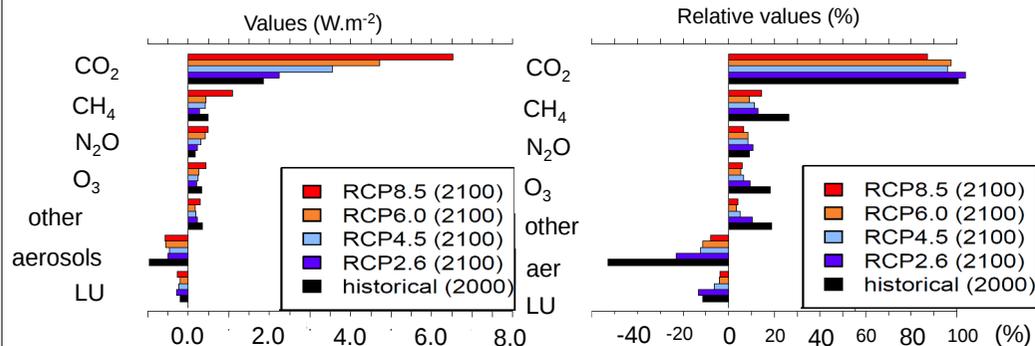
(IPCC, 2013)

## Radiative forcing of future scenarios

Total radiative forcing ( $W.m^{-2}$ )

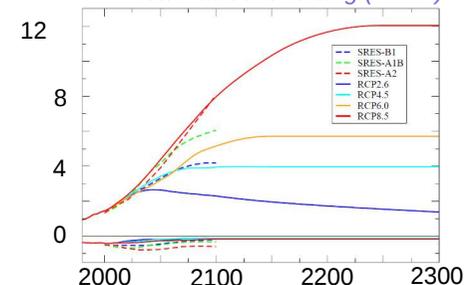


Contribution of individual forcings to total forcing relative to 1850

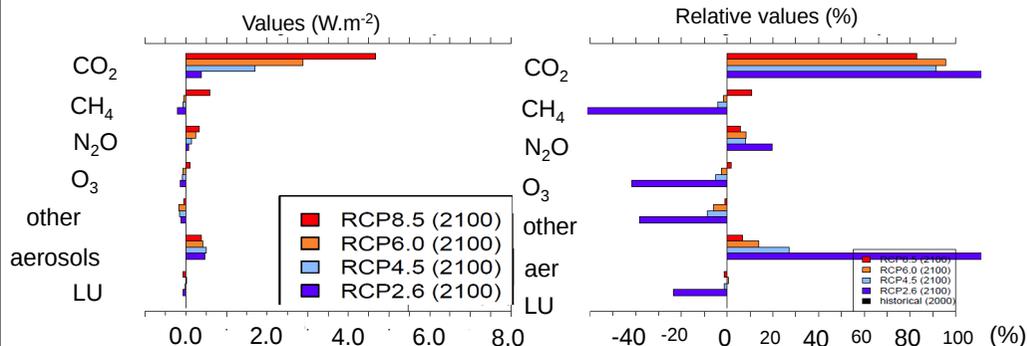


## Radiative forcing of future scenarios

Total radiative forcing ( $W.m^{-2}$ )

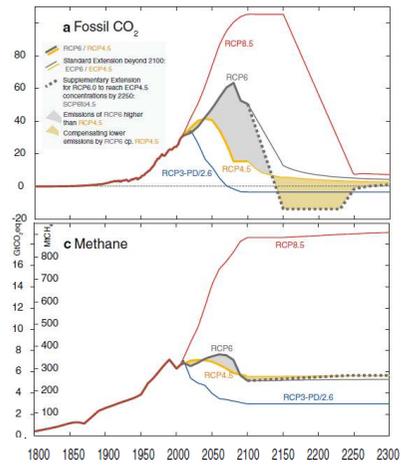


Contribution of individual forcings to total forcing relative to 2000

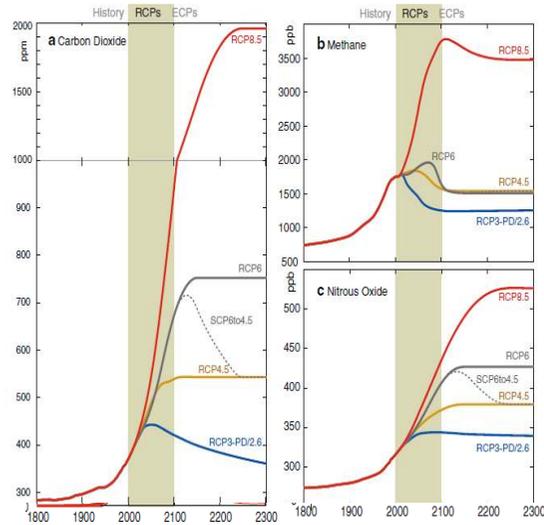


## Future scenarios: gaz emission and concentration

emission

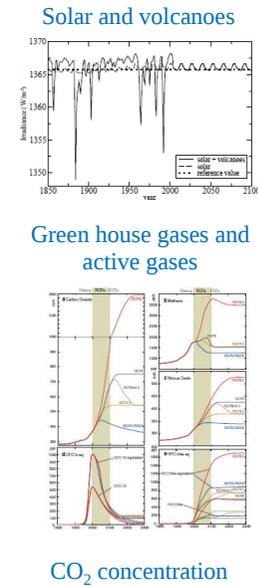


concentration

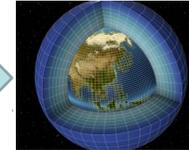


## The IPSL Earth system model

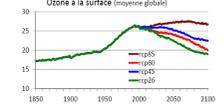
Natural and anthropogenic forcings



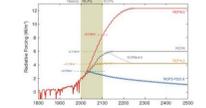
IPSL-CM5A-LR



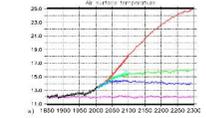
Atmospheric composition



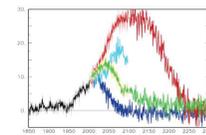
Radiative forcings



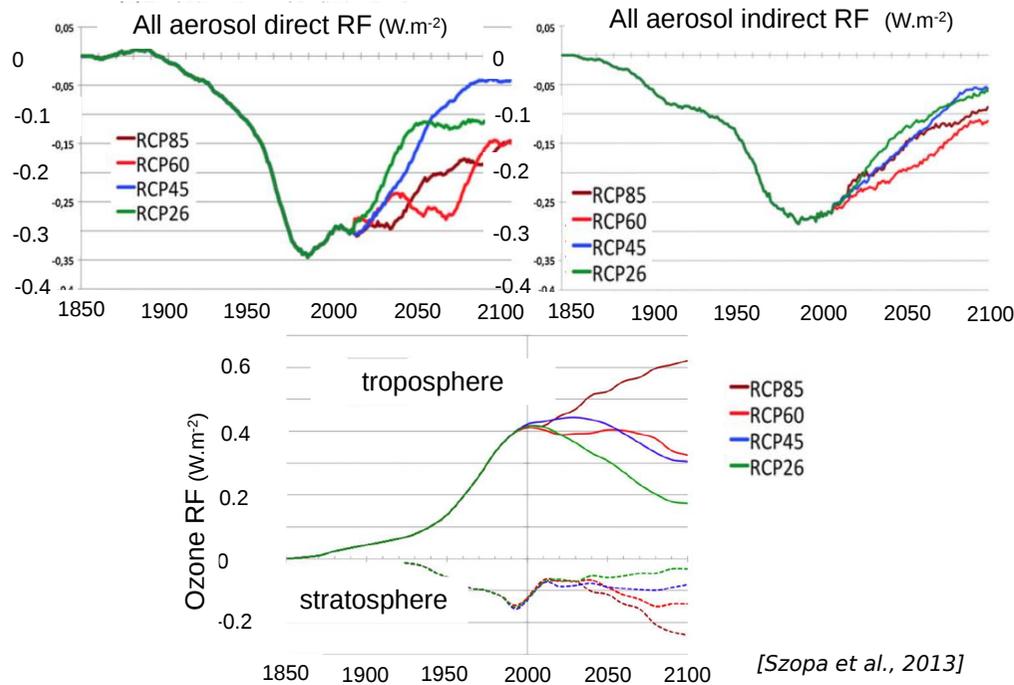
Climate changes



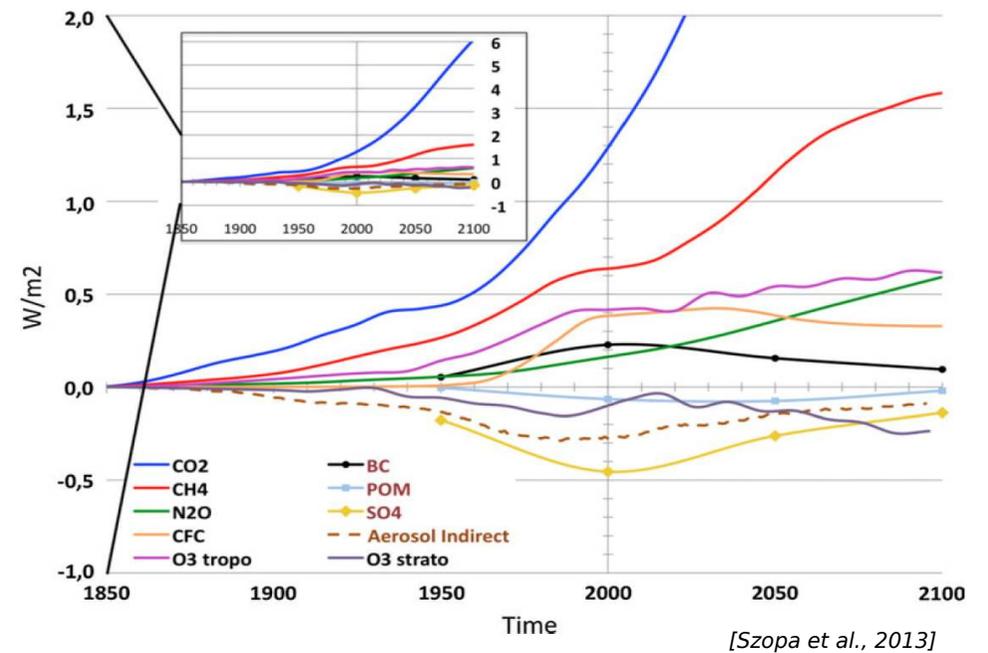
Authorized CO<sub>2</sub> emissions



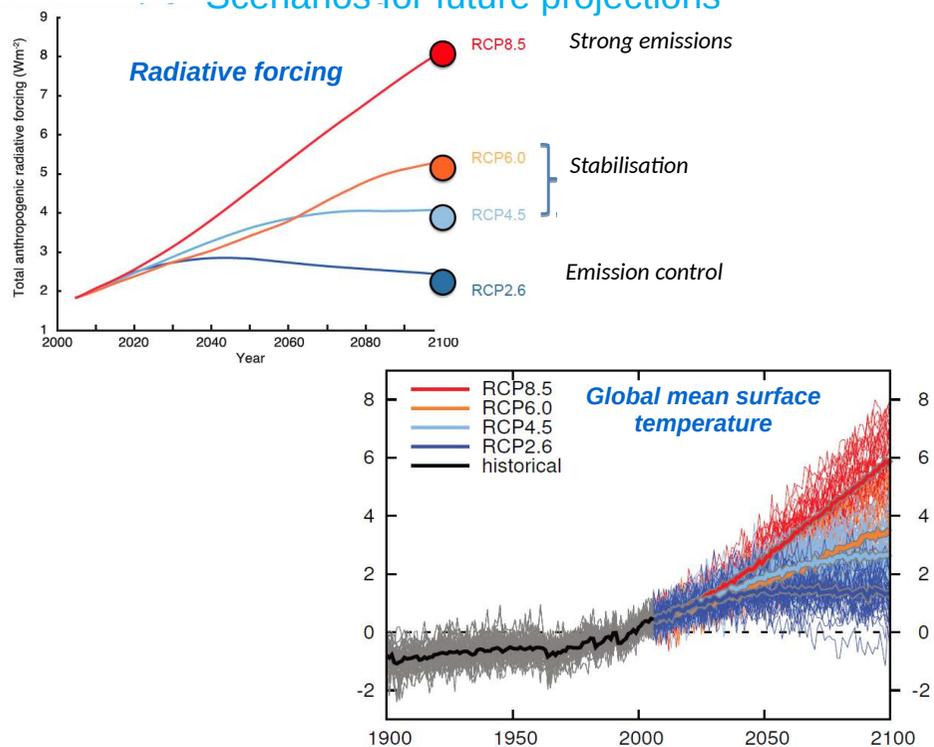
## Ozone and aerosols computations (IPSL-CM5A-LR model)



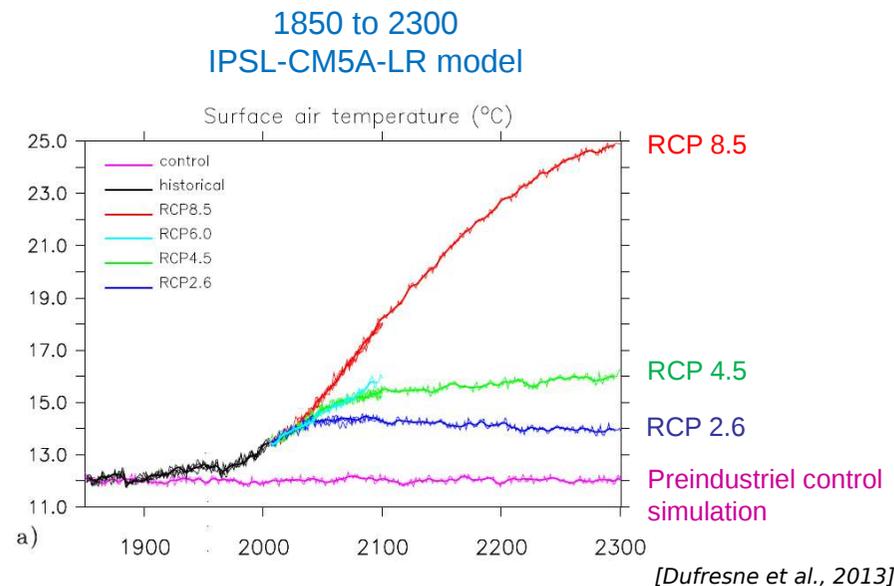
## Radiative forcings for the historical period and the future RCP8.5 scenario (IPSL-CM5A-LR model)



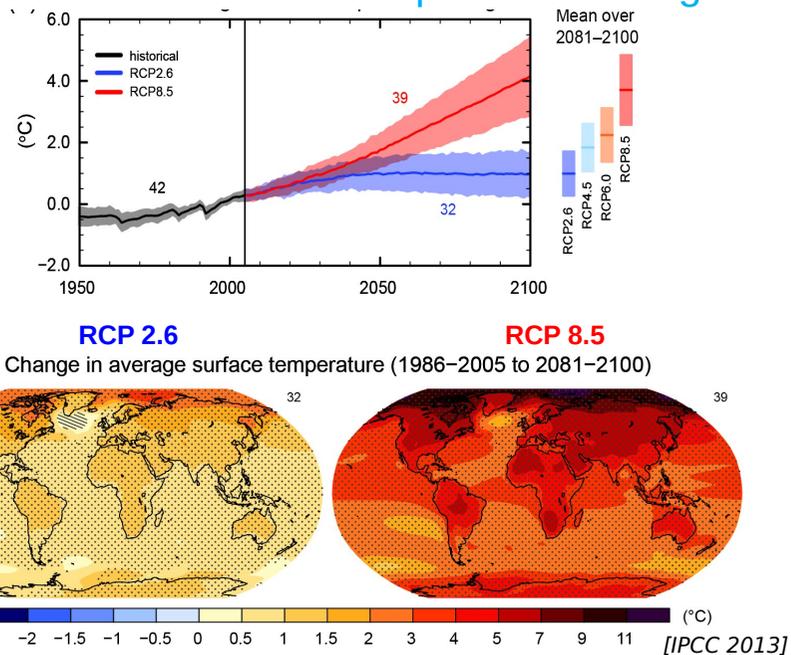
## Scenarios for future projections



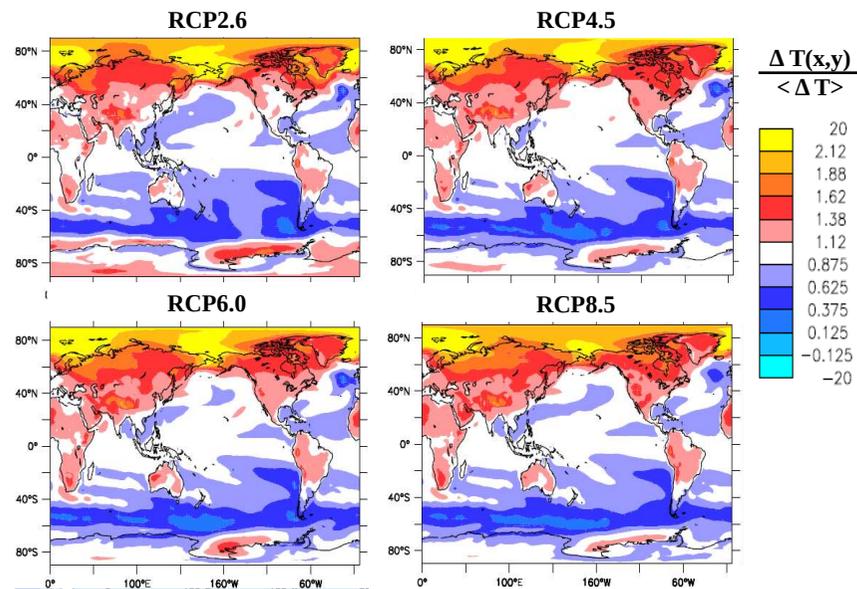
## Global mean surface temperature



## Global mean surface temperature change

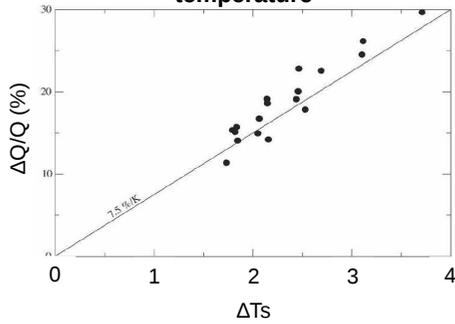


## Spatial distribution of the normalized air surface temperature change $\Delta T(x,y)/\langle \Delta T \rangle$ in 2100



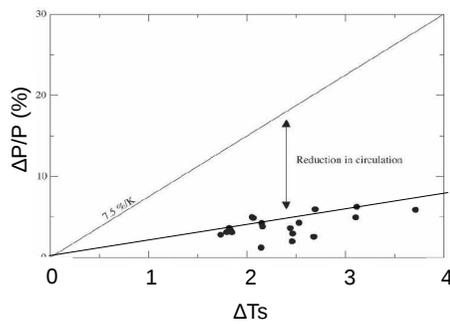
## Precipitation changes

Change of the amount of **water vapor** H<sub>2</sub>O vs change of the average surface temperature



$$\Delta Q/Q (\%) = 7.5 \Delta T_s$$

Change of **precipitation** vs change of the average surface temperature

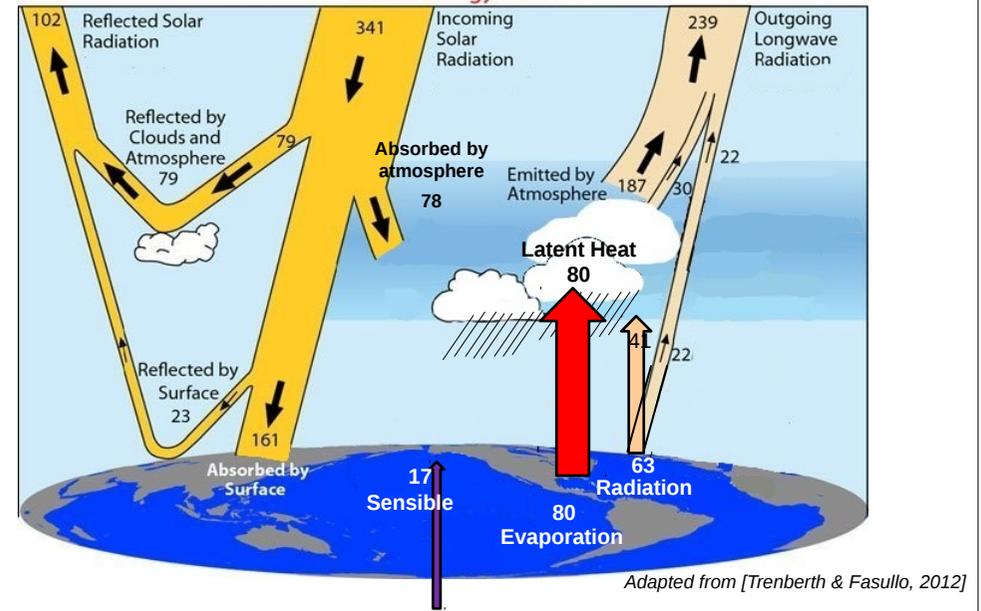


$$\Delta P/P (\%) = 1.5 \Delta T_s$$

The change of the global average precipitation does not depend directly from the change of global average water vapor

(Vecchi & Soden, 2007)

## Global Energy Flows W m<sup>-2</sup>

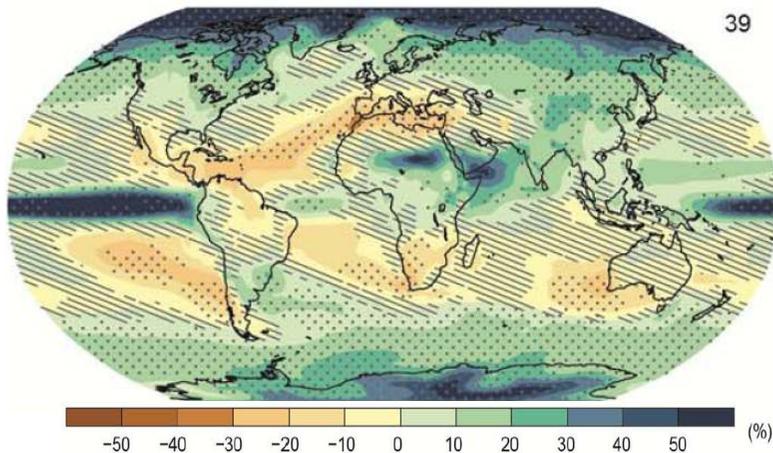


Adapted from [Trenberth & Fasullo, 2012]

The change of the global average precipitation is constrained by the radiative cooling of the atmosphere

## Precipitation changes: Geographical distribution

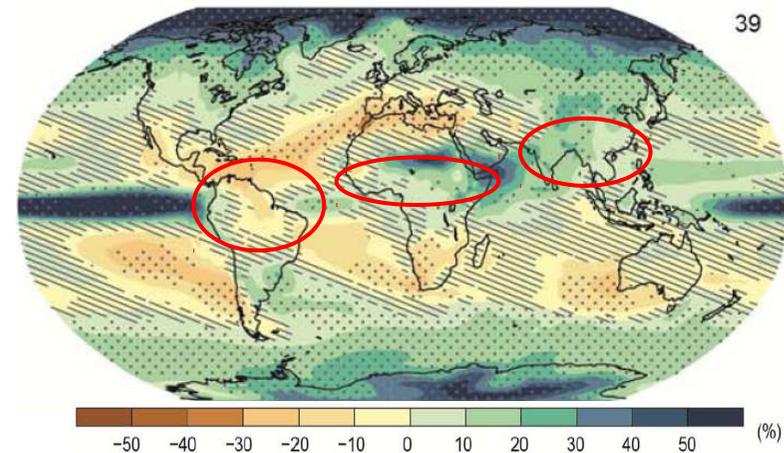
Relative change in average precipitation, RCP8.5 scenario (2081-2100)



[IPCC 2013]

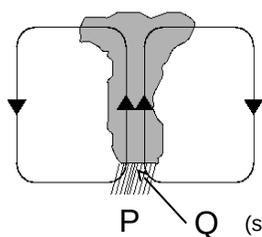
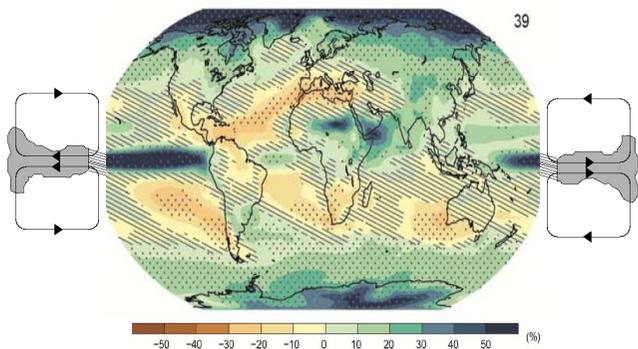
## Precipitation changes: Geographical distribution

Relative change in average precipitation, RCP8.5 scenario (2081-2100)



[IPCC 2013]

## Precipitation changes



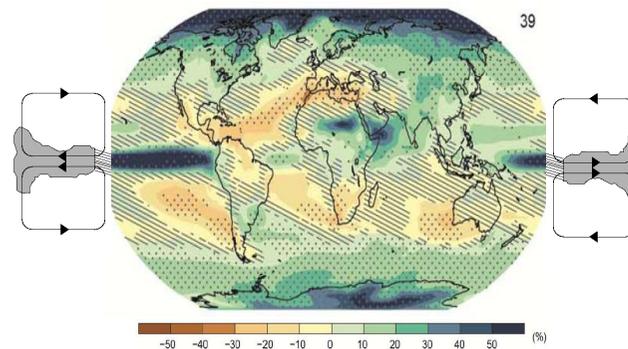
Precipitations changes

$$\Delta P \approx \omega \Delta Q + Q \Delta \omega$$

Thermodynamic response

Dynamic response

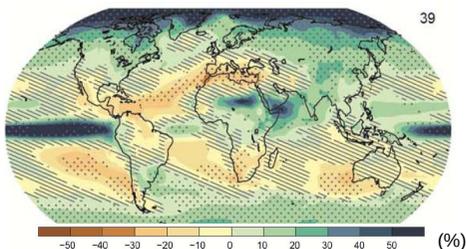
## Precipitation changes



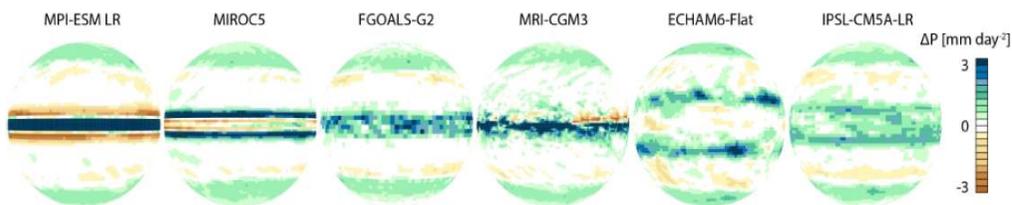
At the global scale:

- Precipitation increases in some regions while decreasing in others
- the contrast between wet and dry regions is expected to increase
- same with the contrast between wet and dry seasons

## Precipitation changes



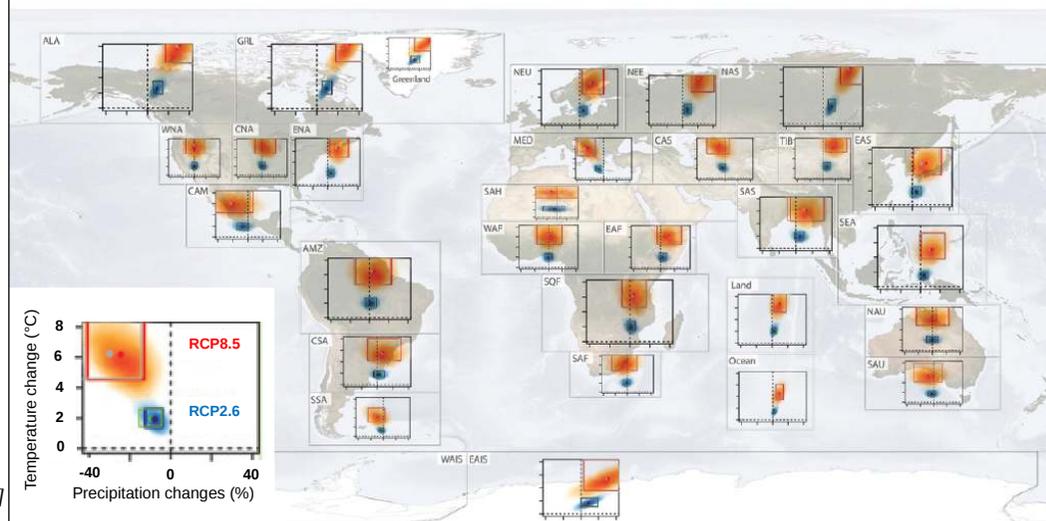
And in a simpler world? Precipitation changes in response to a uniform increase of temperature of 4K for aqua-planets



[Stevens & Bony, 2013]

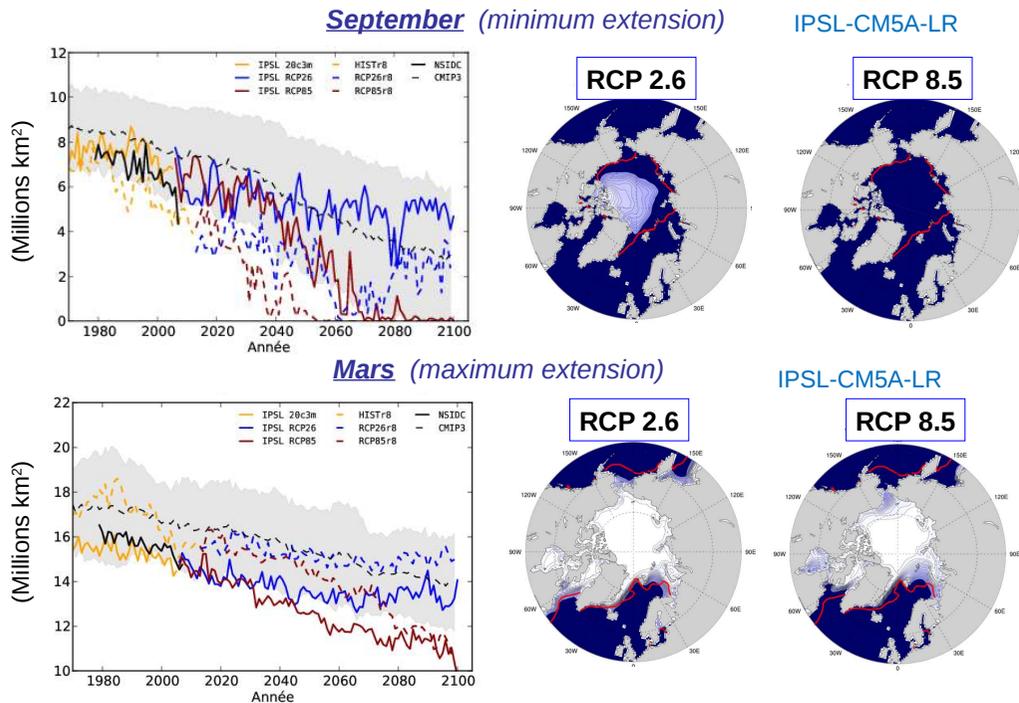
A large fraction of the spread in precipitation changes originates from fundamental problems in water-vapor-temperature-circulation interactions

## Regional precipitation changes using a scaling approach

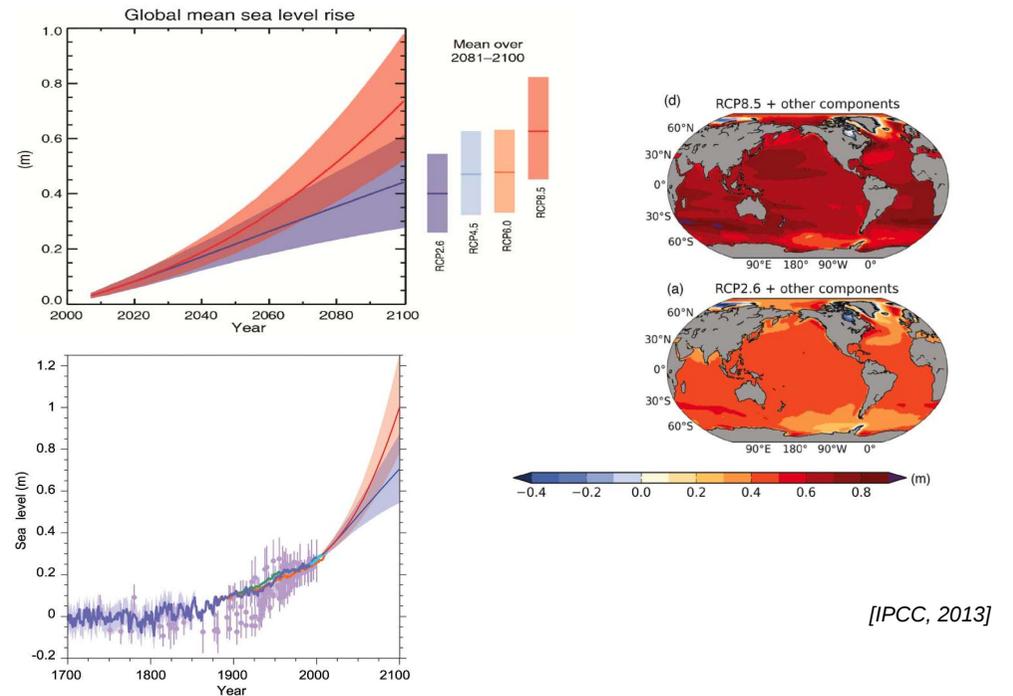


[Frieler et al., 2012]

# Arctic sea-ice 1970-2100

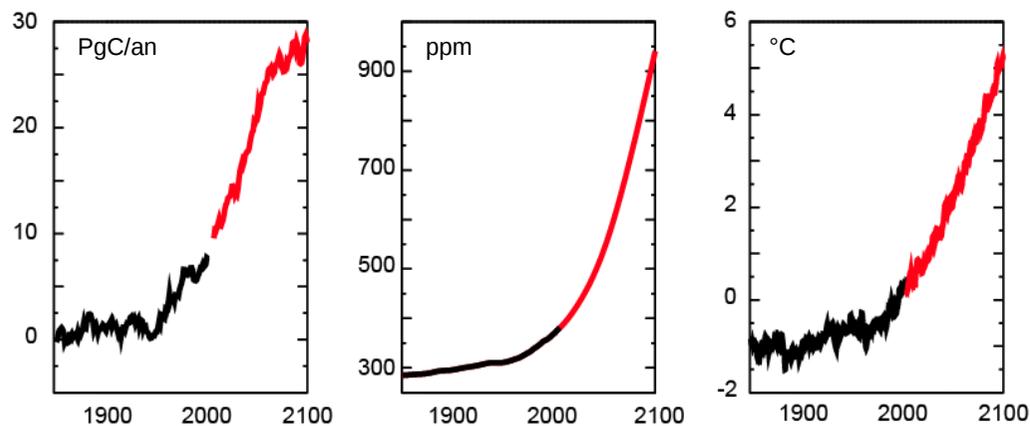


# Sea level rise



# Carbone emission, CO<sub>2</sub> concentrations and global temperature: time constants

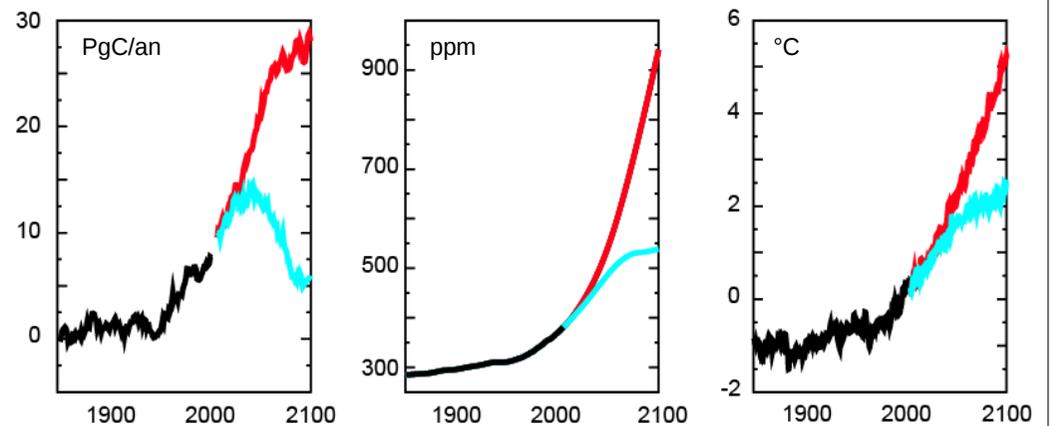
**Higher scenario** : emissions, concentration and temperatures continue to grow



Courtesy L. Bopp

# Carbone emission, CO<sub>2</sub> concentrations and global temperature: time constants

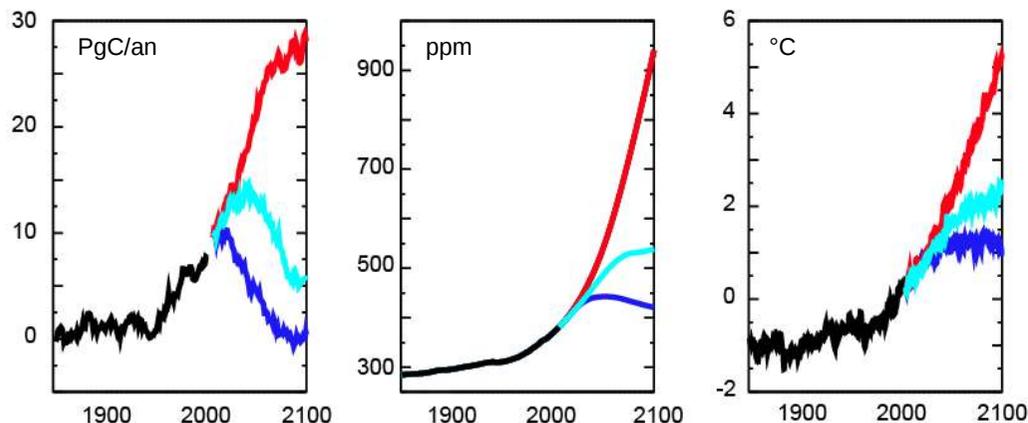
**Higher scenario** : emissions, concentration and temperatures continue to grow  
**Medium scenario** : to stabilize CO<sub>2</sub> concentration 550 ppm, emissions need to be strongly reduced. However, temperature will continue to increase



Courtesy L. Bopp

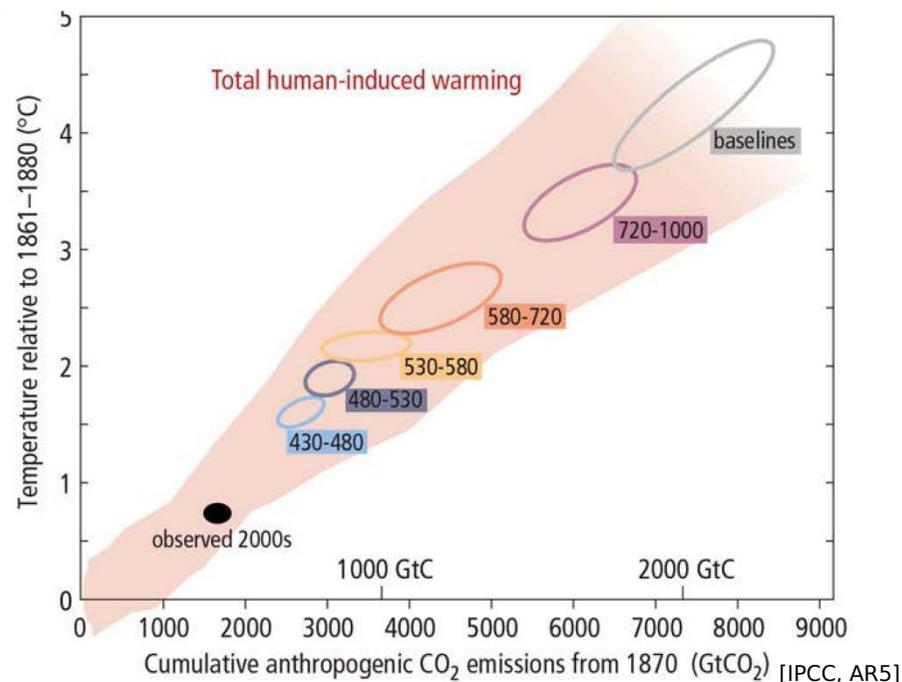
## Carbone emission, CO<sub>2</sub> concentrations and global temperature: time constants

- Higher scenario** : emissions, concentration and temperatures continue to grow
- Medium scenario** : to stabilize CO<sub>2</sub> concentration 550 ppm, emissions need to be strongly reduced. However, temperature will continue to increase
- Lower Scenario** : to limit a 2° global warming, CO<sub>2</sub> concentration has to be limited to less than 450 ppm, and emissions need to be 0 before the end of the century



Courtesy L. Bopp

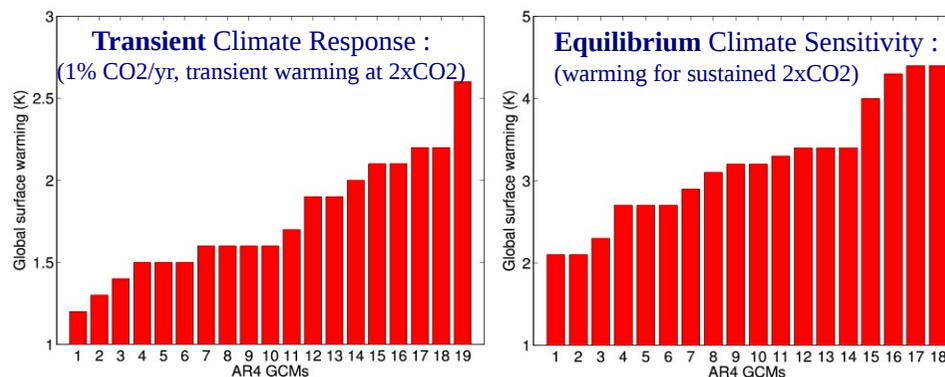
## Temperature increase as a function of cumulative CO<sub>2</sub> emissions.



## Outlook

- Numerical climate models
- Simulating and attributing recent climate changes
- Scenarios and future climate changes
- Climate sensitivity and climate feedbacks
- Climate variability and past climate changes

## Climate sensitivity estimates from CMIP3 GCMs (IPCC-AR4)



Climate sensitivity and TCR estimates depend on :

- radiative forcing
- climate feedbacks
- ocean heat uptake (transient only)

[IPCC, 2007]

*How do these different components contribute to inter-model differences in climate sensitivity ?*

# Climate feedback

$$\Delta R = \Delta Q + \lambda \Delta T_s$$

Global average surface temperature change

flux at TOA change      radiative forcing      "feedback parameter"

When a new equilibrium is reached,  $\Delta R=0$

$$\Delta T_s^e = \frac{-\Delta Q_t}{\lambda}$$

$$\lambda = \lambda_P + \lambda_w + \lambda_L + \lambda_c + \lambda_\alpha$$

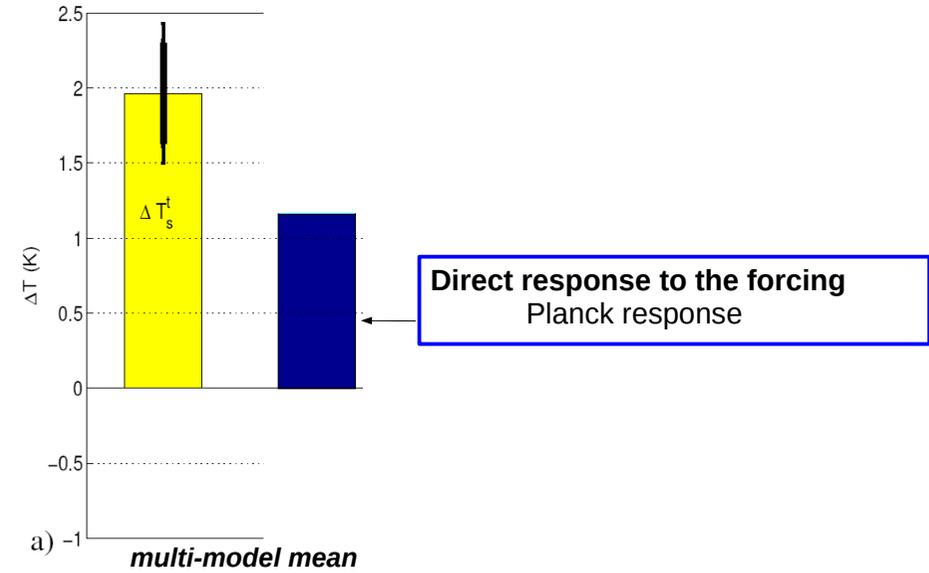
Planck      water vapor      lapse rate      clouds      surface albedo

$$\lambda = \frac{\partial R}{\partial T_s} = \sum_x \frac{\partial R}{\partial x} \frac{\partial x}{\partial T_s}$$

$\lambda$  can be diagnosed from model results with different technics

# Transient temperature response to a CO<sub>2</sub> doubling

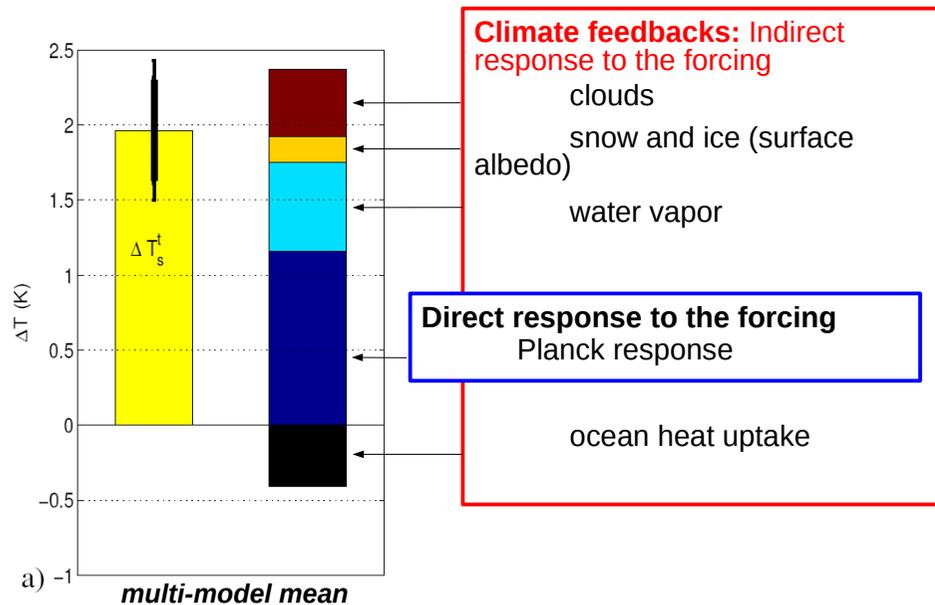
(CO<sub>2</sub> increase 1%/year, 70 years)



(Dufresne & Bony, 2008)

# Transient temperature response to a CO<sub>2</sub> doubling

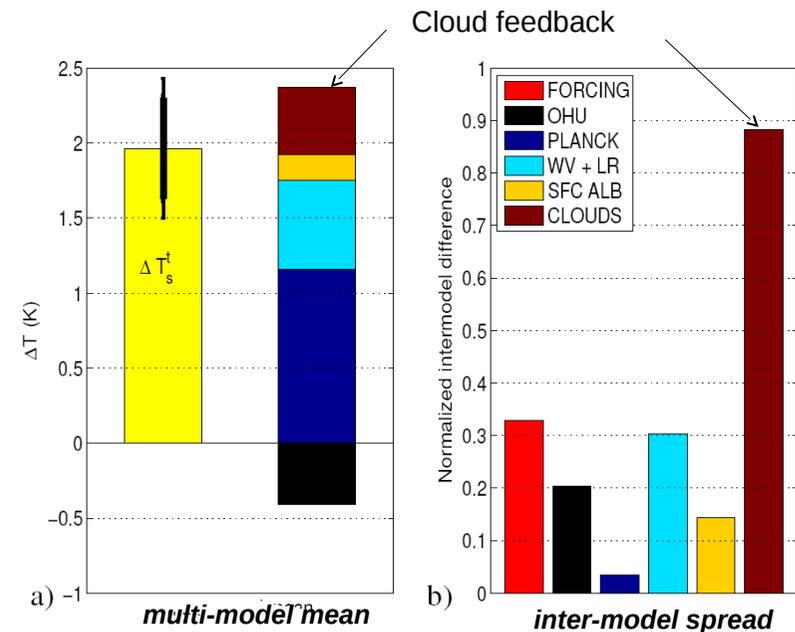
(CO<sub>2</sub> increase 1%/year, 70 years)



(Dufresne & Bony, 2008)

# Transient temperature response to a CO<sub>2</sub> doubling

(CO<sub>2</sub> increase 1%/year, 70 years)

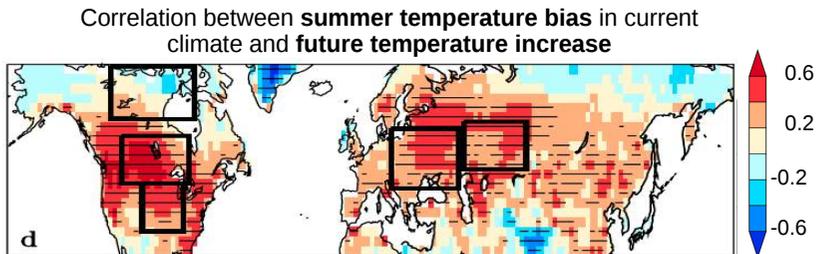


(Dufresne & Bony, 2008)

## How model defaults influence model projections?

As a whole, there is **no simple relationship** between **model performances** with current conditions and **climate change projections simulated by models**.

However, some relation between model characteristics and simulated climate change under global warming exists for some variables/processes.



Warm bias is mainly due to a erroneous representation of evaporation and cloud cover over continents in models.

[Cheruy et al., 2014]

## Outlook

- I. Numerical climate models
- II. Simulating and attributing recent climate changes
- III. Scenarios and future climate changes
- IV. Climate sensitivity and climate feedbacks
- V. Climate variability and past climate changes

## Internal variability and variations due to forcings

Climate variations have different origins:

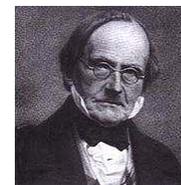
$$\underbrace{\Delta T}_{\text{variation}} \approx \underbrace{\Delta T_{int}}_{\text{Internal variability}} + \underbrace{\frac{\partial T}{\partial Q} \Delta Q_{nat}}_{\text{Response to natural forcings}} + \underbrace{\frac{\partial T}{\partial Q} \Delta Q_{ant}}_{\text{Response to anthropogenic forcings}}$$

Natural variability

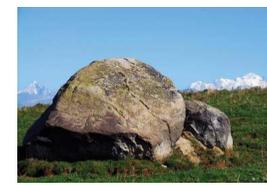
- The relative importance of these various termes depends on the spatial and time average considered, and on the amplitude of the forcings
- The differences between observations and models or between model results can include part or all of these terms, depending on the experimental setup

## The discovery of past climate changes

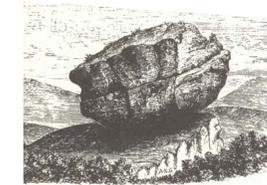
Hypothesis of glacial periods (1840-1860)



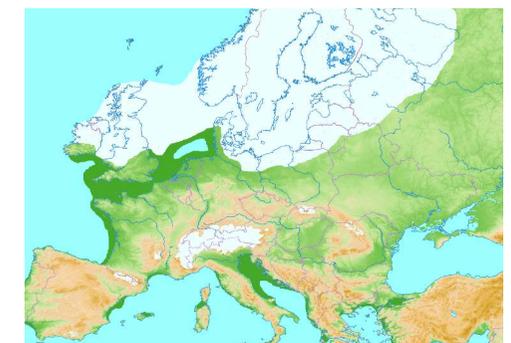
Jean de Charpentier



Erratic rocks



Louis Agassiz



# The discovery of past climate changes

A period described by cave paintings



Cosquer



Lascaux



Chauvet

Origin of these variations : sun or CO<sub>2</sub> (1860-1900) ?

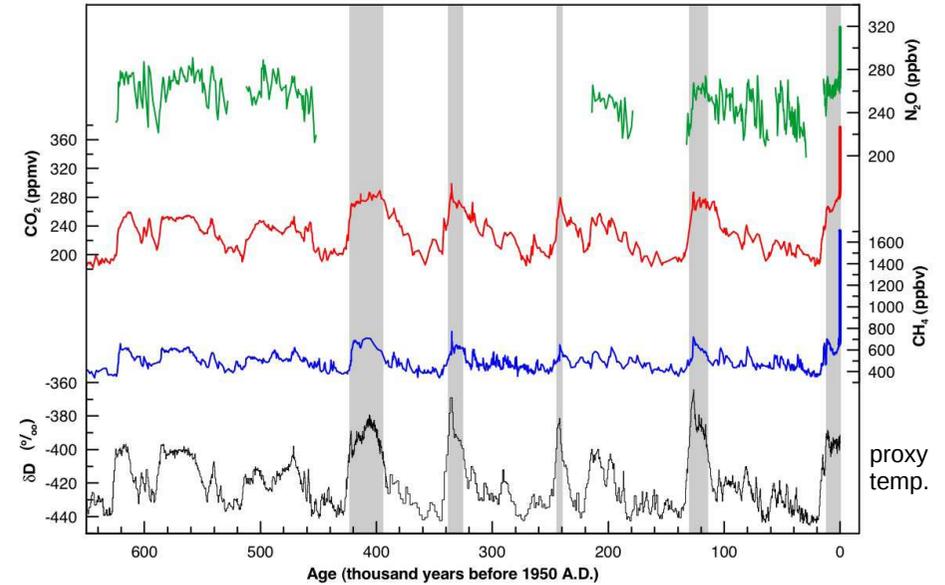


James Croll



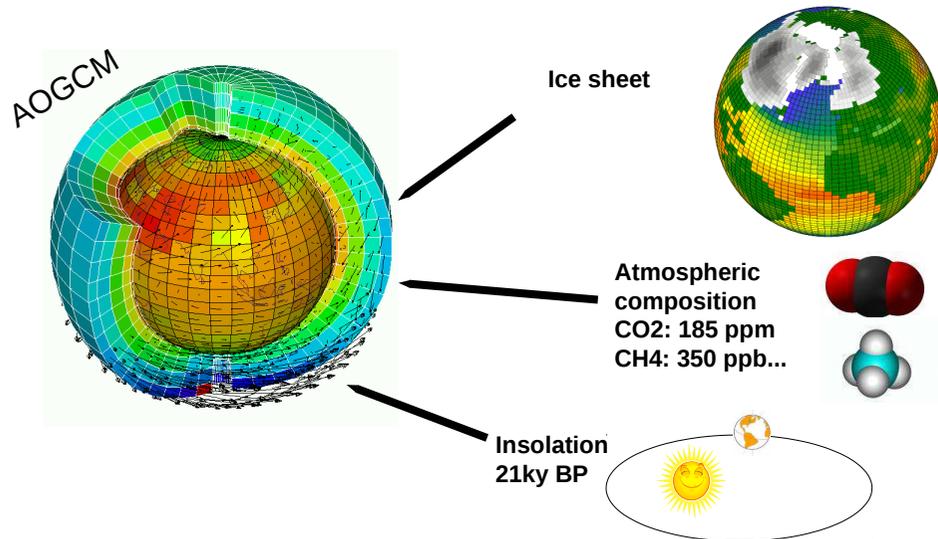
Svante Arrhenius

# Paleoclimate changes



[IPCC, 2013]

# Simulation of Last Glacial Maximum (LGM)



Greenhouse gas forcing ~ future climate  
Other main forcings: ice sheet

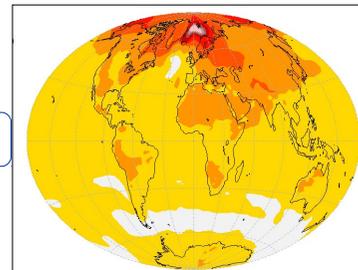
cf. <http://pmip3.lsce.ipsl.fr>

# Change in surface temperature

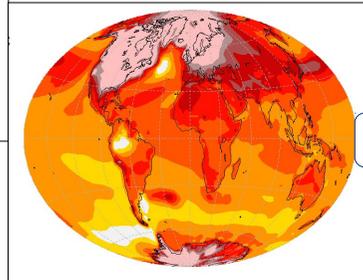
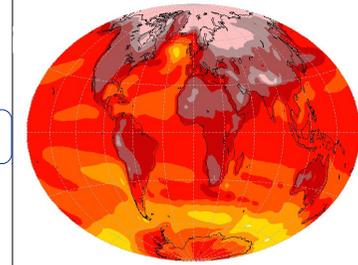
Difference between 2100 and 1990

IPSL-CM5A-LR

RCP2.6



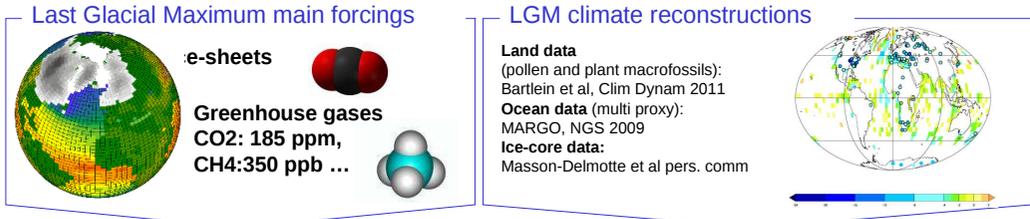
RCP8.5



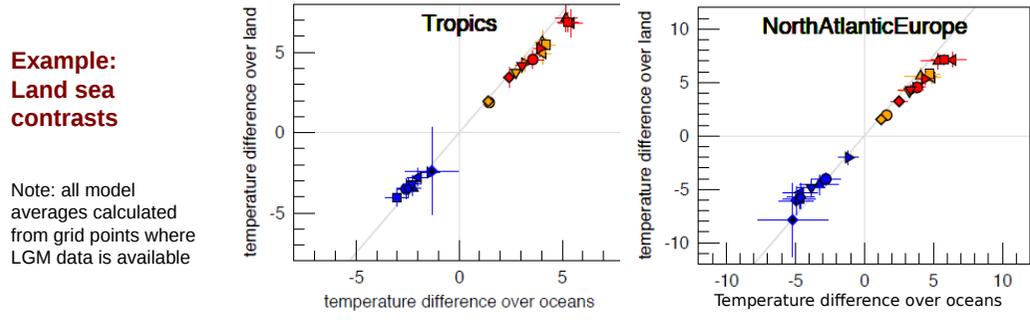
Glacial



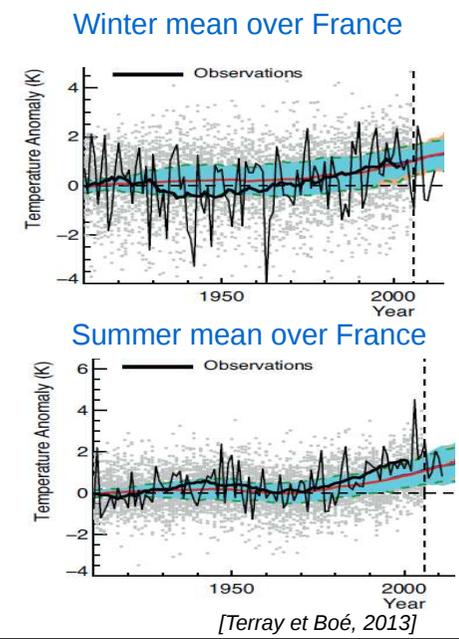
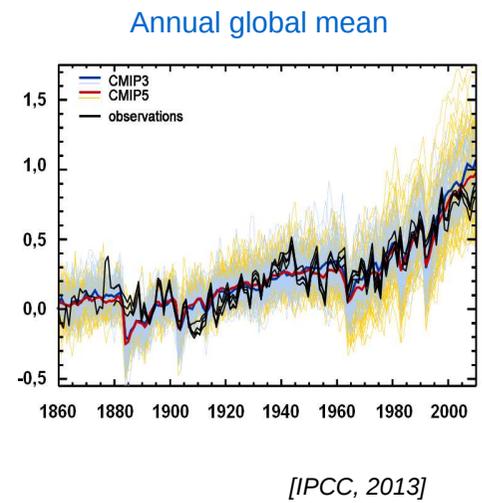
# Land-sea contrasts and polar amplification in past and future climates



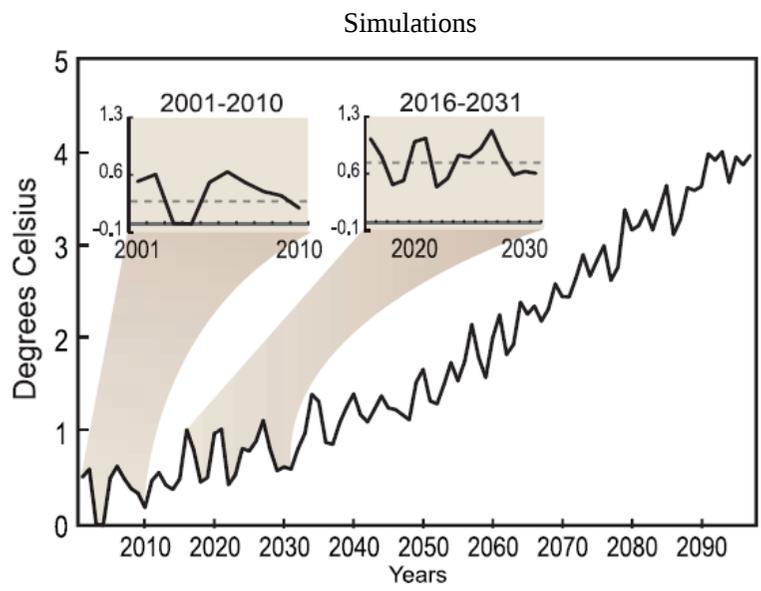
**Relationships between LGM vs higher CO<sub>2</sub> climates?**  
**Are the large scale relationships stable? Can we evaluate them from paleodata ?**



# Surface temperature evolution: observation and models



# Climate change and climate variability



# Climate change and climate variability

50 years trend of the winter surface temperature (°C/50 years)

