



IPSL Climate Modelling Centre



Physical basis of climate and climate change modelling

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Outlook

- I. Emergence of climate and climate change science
- II. Climate modeling
- III. Climate and climate change simulations
- IV. Understanding some climate phenomena
- V. Climate changes and climate variability
- VI. Conclusions

Emergence of the physics of climate

J. Fourier:

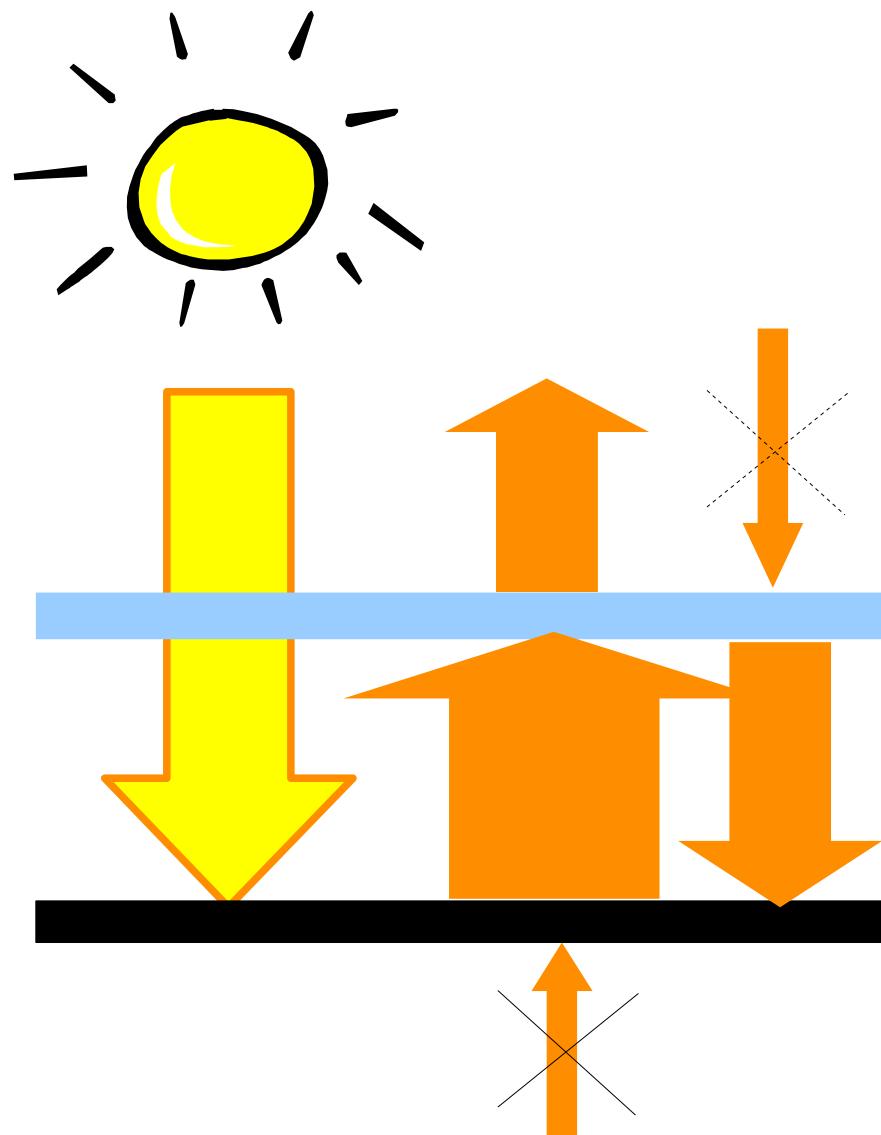
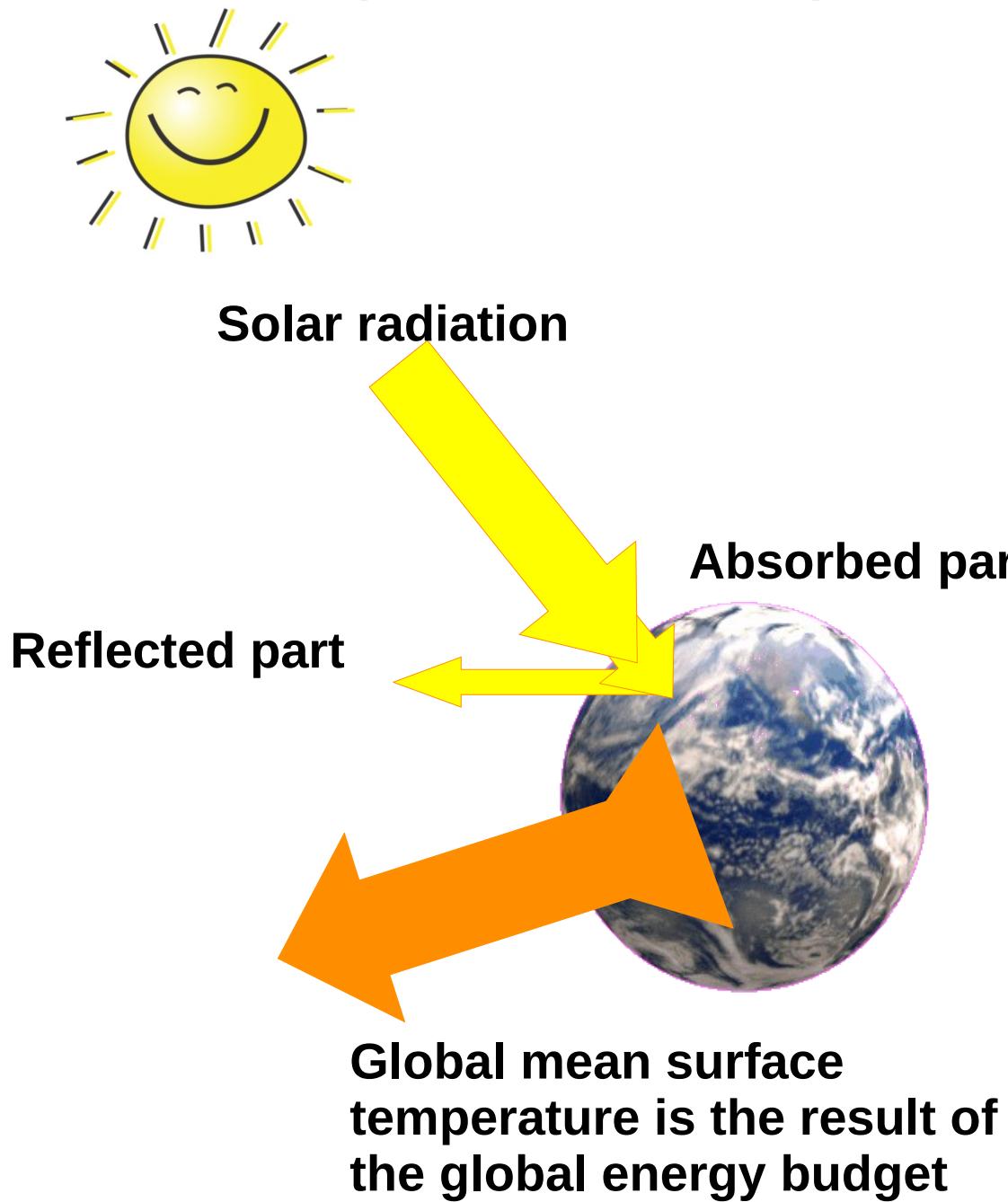
- *Mémoire sur les températures du globe terrestre et des espaces planétaires*, Mémoires de l'Académie des Sciences de l'Institut de France, 1824
- *General remarks on the Temperature of the Terrestrial Globe and the Planetary Spaces*; American Journal of Science, Vol. 32, N°1, 1837.



Joseph Fourier
(1768-1830)

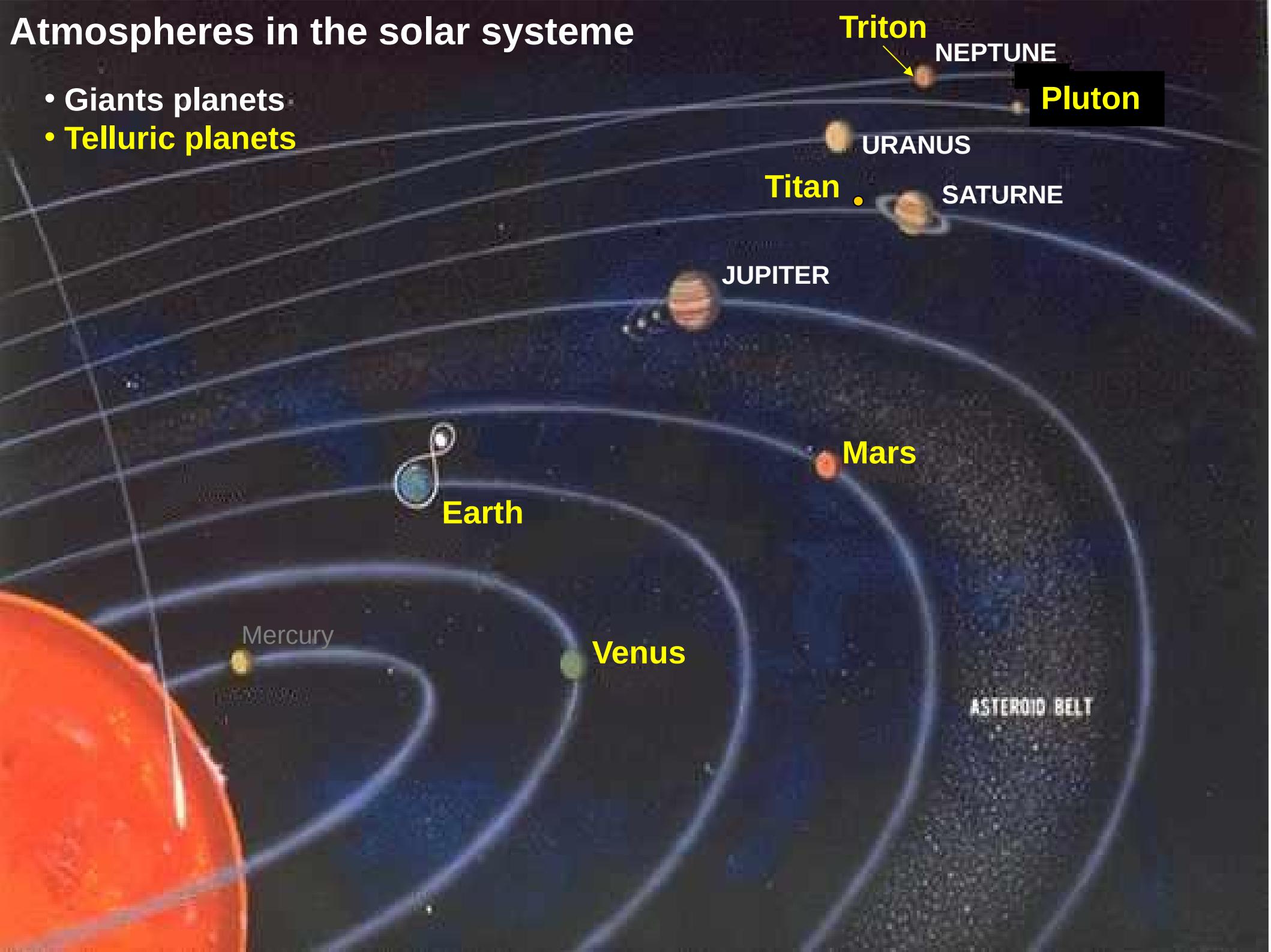
- He consider the Earth like any other planet
- The energy balance equation drives the temperature of all the planets
- The major heat transfers are
 1. Solar radiation
 2. Infra-red radiation
 3. Diffusion with the interior of Earth

Equilibrium temperature of a planet



Atmospheres in the solar system

- Giants planets
- Telluric planets



Emergence of the physics of climate

J. Fourier:

- *Mémoire sur les températures du globe terrestre et des espaces planétaires*, Mémoires de l'Académie des Sciences de l'Institut de France, 1824
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Joseph Fourier

(1768-1830)

- He **envisages the importance of any change of the sun** « *The least variation in the distance of that body[the sun] from the earth would occasion very considerable changes of temperature. »*
- He **envisages that climate may change**: « *The establishment and progress of human society, and the action of natural powers, may, in extensive regions, produce remarkable changes in the state of the surface, the distribution of waters, and the great movements of the air. Such effects, in the course of some centuries, must produce variations in the mean temperature for such places ».*

Equilibrium temperature of a planet



Incoming solar radiation on a **plan**: $F_0 = 1364 \text{ W.m}^{-2}$

Incoming solar radiation on a **sphere**: $F_s = F_0 / 4 = 341 \text{ W.m}^{-2}$

All the incoming solar radiation is absorbed : $F_a = 240 \text{ W.m}^{-2}$



$T_s = 278 \text{ K (5°C)}$

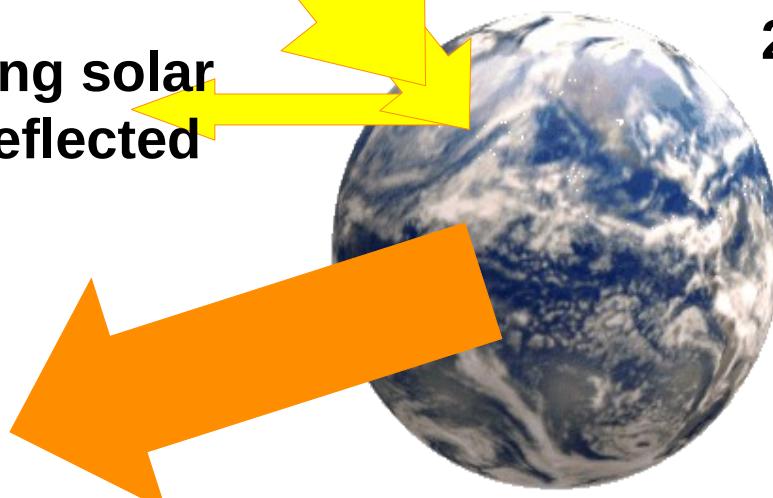
Equilibrium temperature of a planet



Incoming solar radiation on a **plan**: $F_0 = 1364 \text{ W.m}^{-2}$

Incoming solar radiation on a **sphere**: $F_s = F_0 / 4 = 341 \text{ W.m}^{-2}$

1/3 of incoming solar radiation is reflected

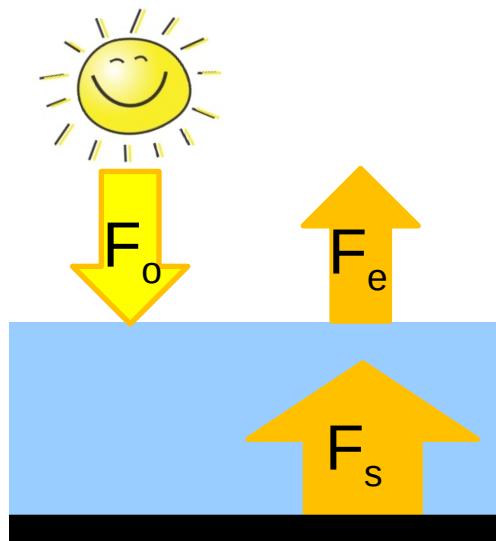


$T_s = 255\text{K} (-18^\circ\text{C})$

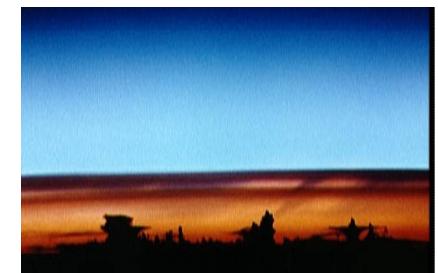
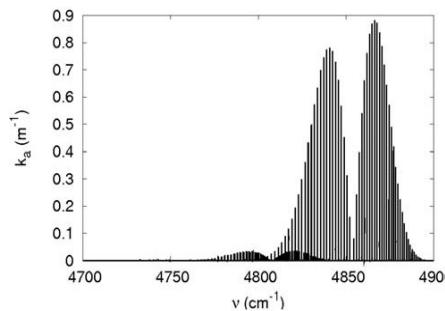
2/3 of incoming solar radiation is absorbed : $F_a = 240\text{W.m}^{-2}$

Global mean surface temperature is 15°C due to greenhouse effect

What radiation heat transfer theory tell us



Greenhouse effect: $G = F_s - F_e$



Gas radiative properties

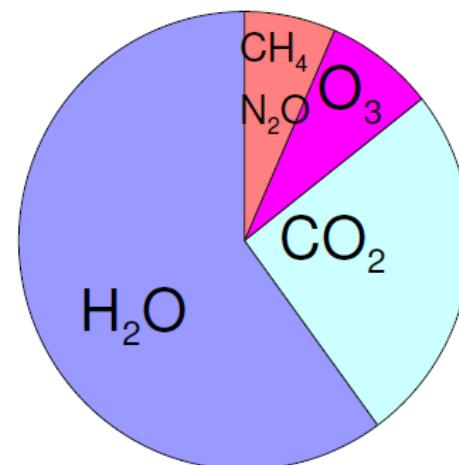
Atmospheric characteristics

Computation of the radiative fluxes and the greenhouse effect

Current greenhouse effect: $G \approx 150 \text{ W.m}^{-2}$

Contribution of atmospheric gases (clear sky)

Water vapour	60%
CO_2	26%
Ozone O_3	8%
$\text{N}_2\text{O} + \text{CH}_4$	6%



For a doubling of CO_2 concentration, green house effect increases by $\approx 3.7 \text{ W.m}^{-2}$

From radiative transfer computation to climate modelling

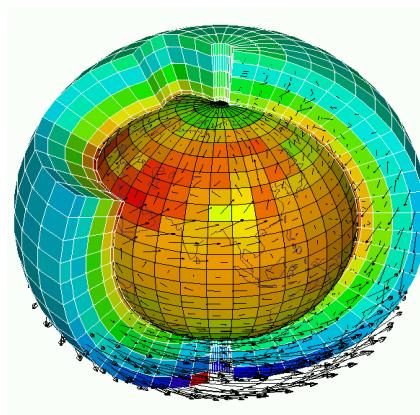
For a doubling of the CO₂ concentration:

- the green house effect increases by 3.7 W.m⁻²
- the temperature increases by ≈ 1.2 K, if nothing change except an uniform increase of temperature that only impact radiation

But feedbacks exist:

- Snow and sea ice reflect solar radiation; if they decrease, more solar energy will be absorbed ⇒ **positive feedback**
- Water vapour is the main greenhouse gas; if it increases, the greenhouse effect will be enhanced ⇒ **positive feedback**
- Clouds reflect solar radiation and contribute to the greenhouse effect; if they change, the energy budget will be modified ⇒ **positive or negative feedback**

Need of 3D numerical climate models



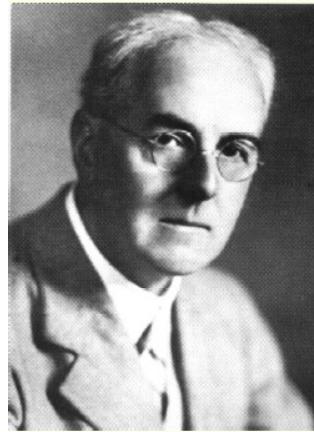
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Numerical climate models (numerical weather simulators)



Wilhelm Bjerknes
(1862–1951)



L. F. Richardson
(1881–1953)



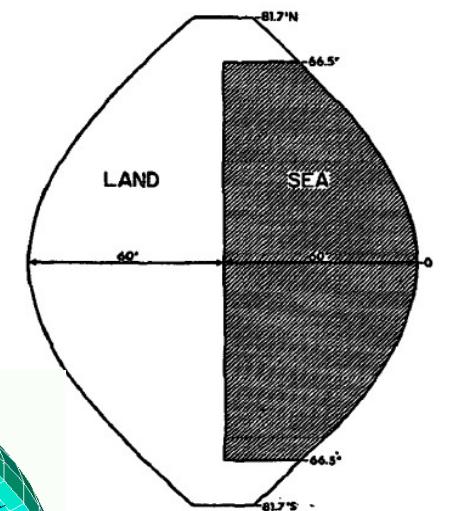
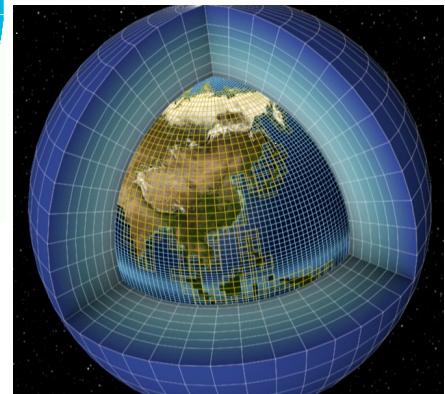
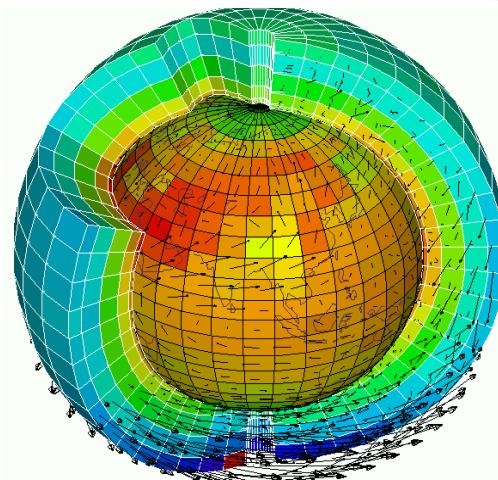
J. von Neumann
(1903–1957)



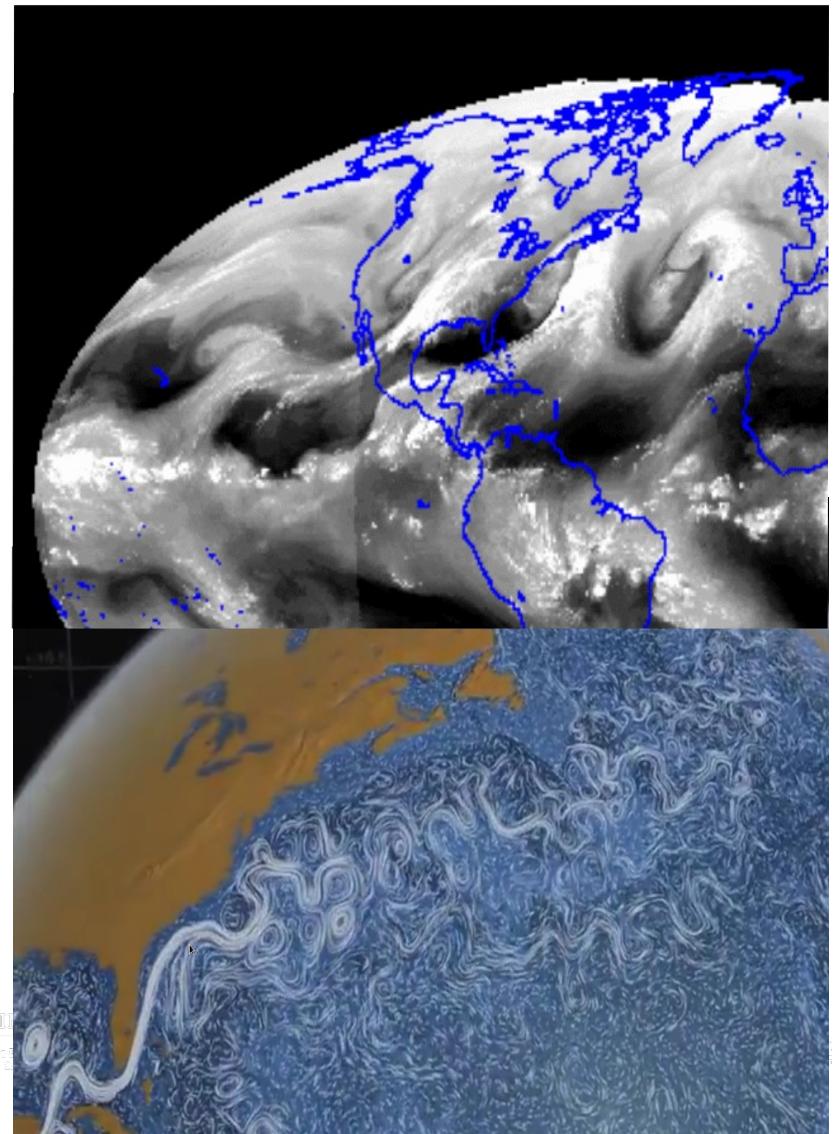
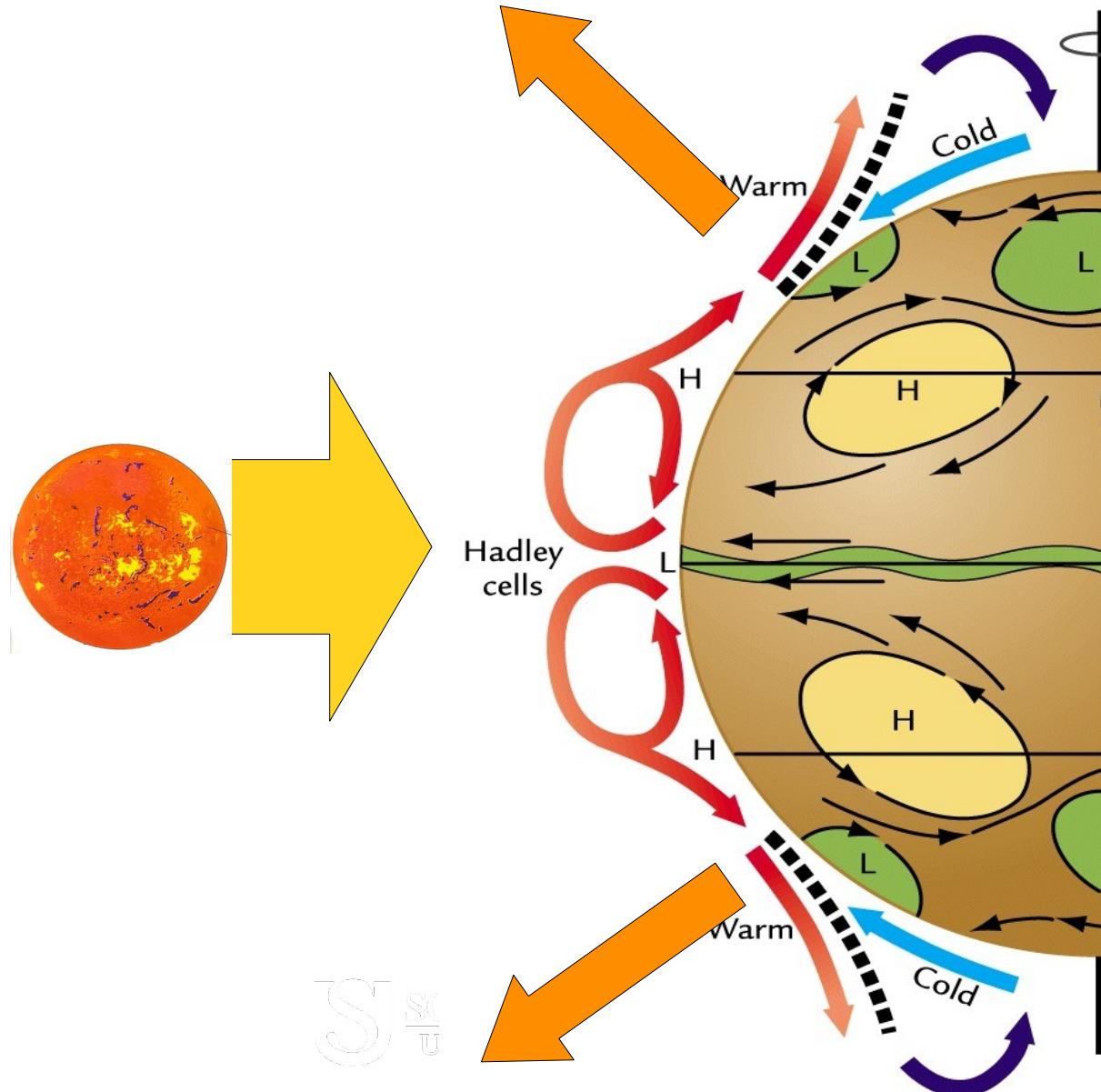
Jule Charney
(1917–1981)



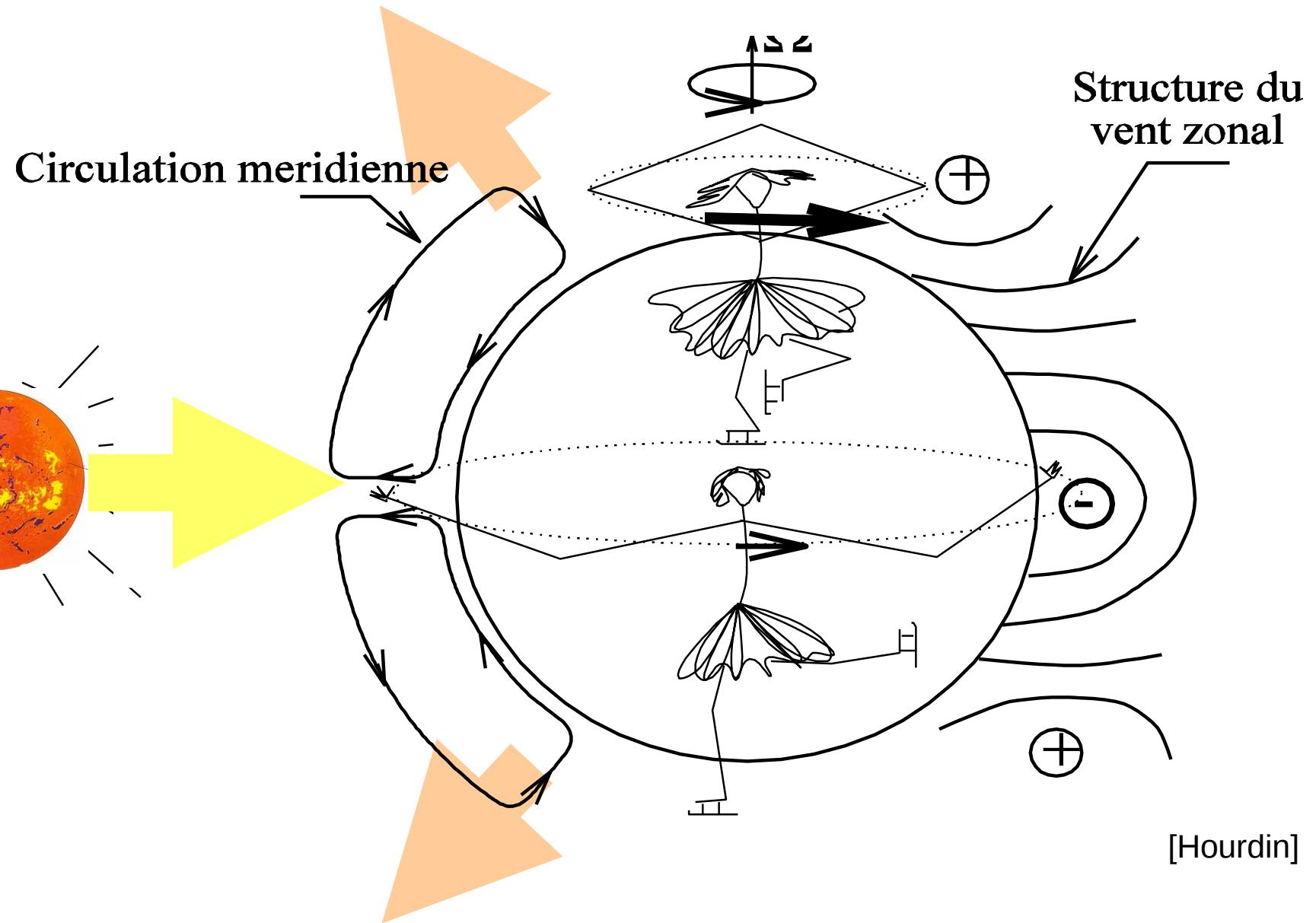
Syukuro Manabe
(1931–)



Large scale circulation: Meridional heat transport and the effects of Earth rotation

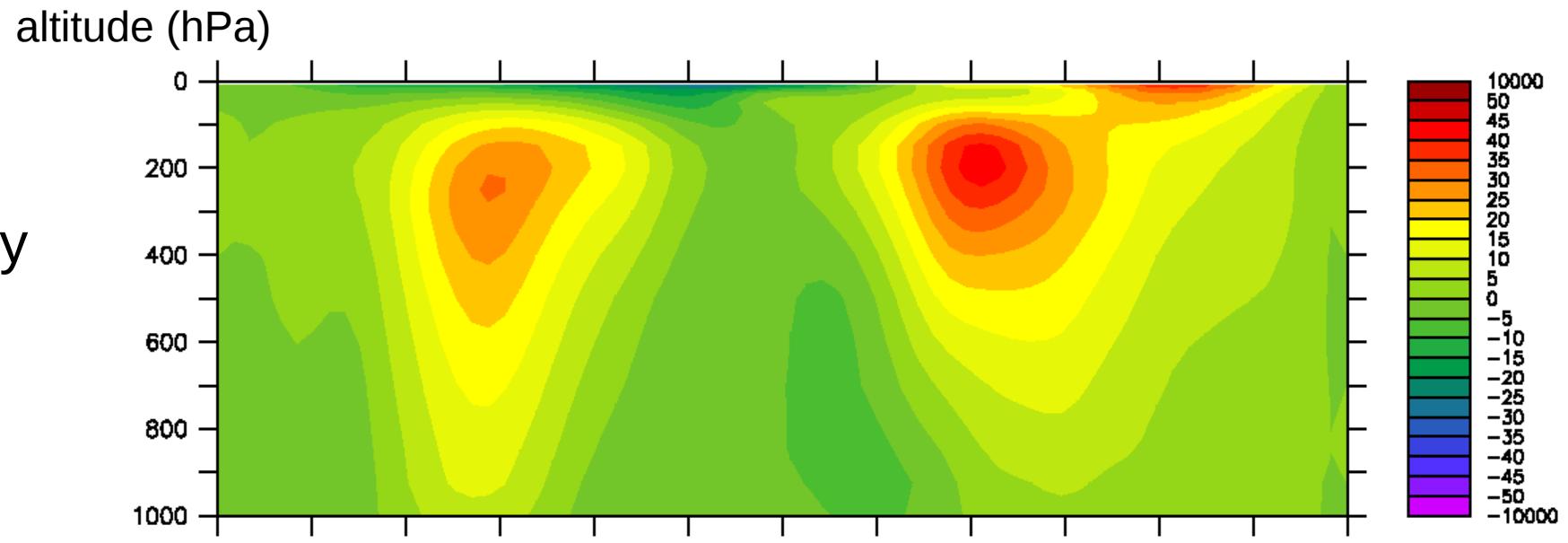


Hadley cell and its extension polarward

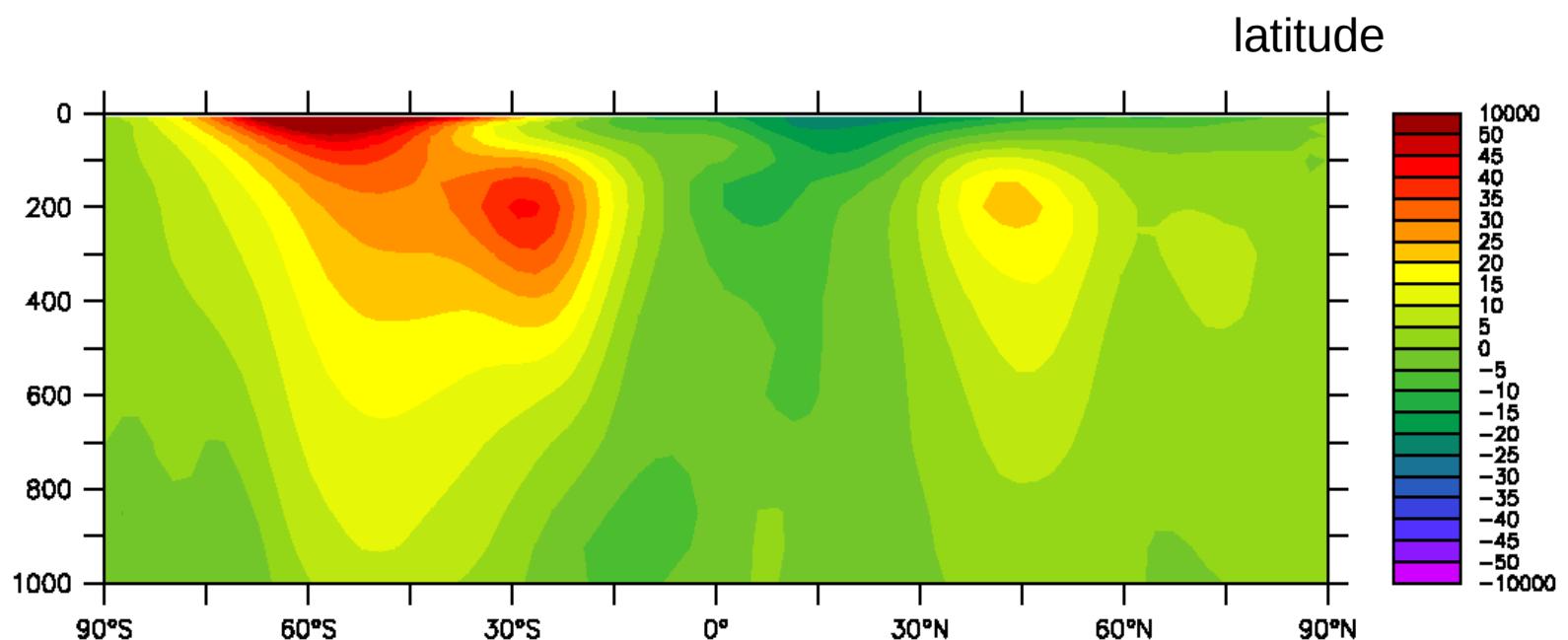


Zonal wind

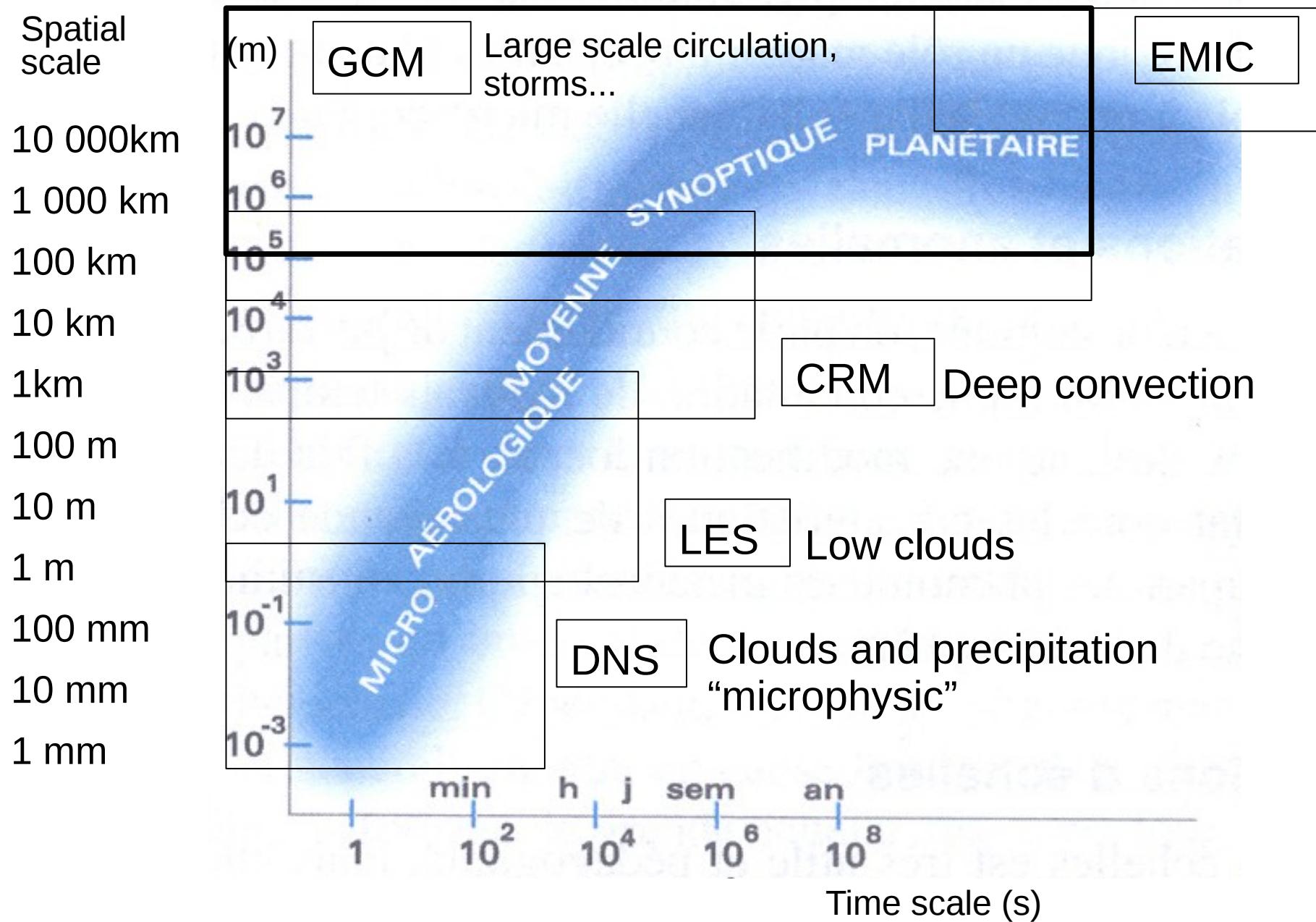
January



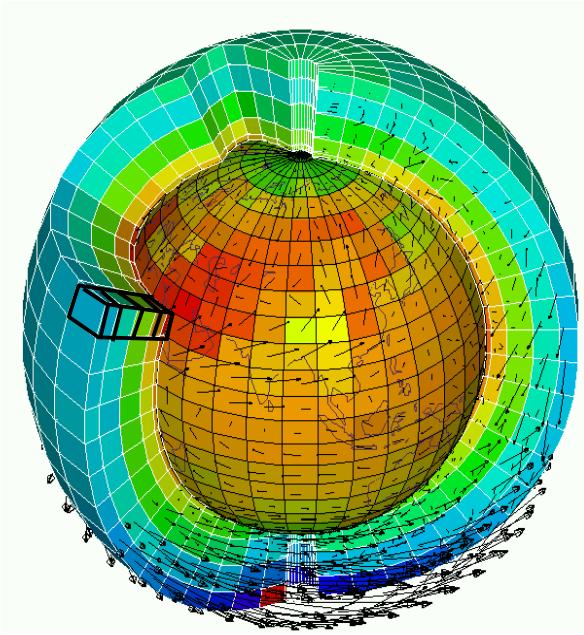
July



Relevant spatial and time scales



General circulation models (GCMs)



Dynamical core : discretized version of the equations of fluid mechanics

- Mass Conservation
 $D\rho/Dt + \rho \operatorname{div} \underline{U} = 0$
- Energy Conservation
 $D\theta / Dt = Q / Cp (p_0/p)^{\kappa}$
- Momentum Conservation
 $D\underline{U}/Dt + (1/\rho) \operatorname{grad} p - g + 2 \underline{\Omega} \wedge \underline{U} = \underline{F}$
- Conservation of Water (and other species)
 $Dq/Dt = S_q$

In red, source terms : other than fluid mechanics and unresolved scales

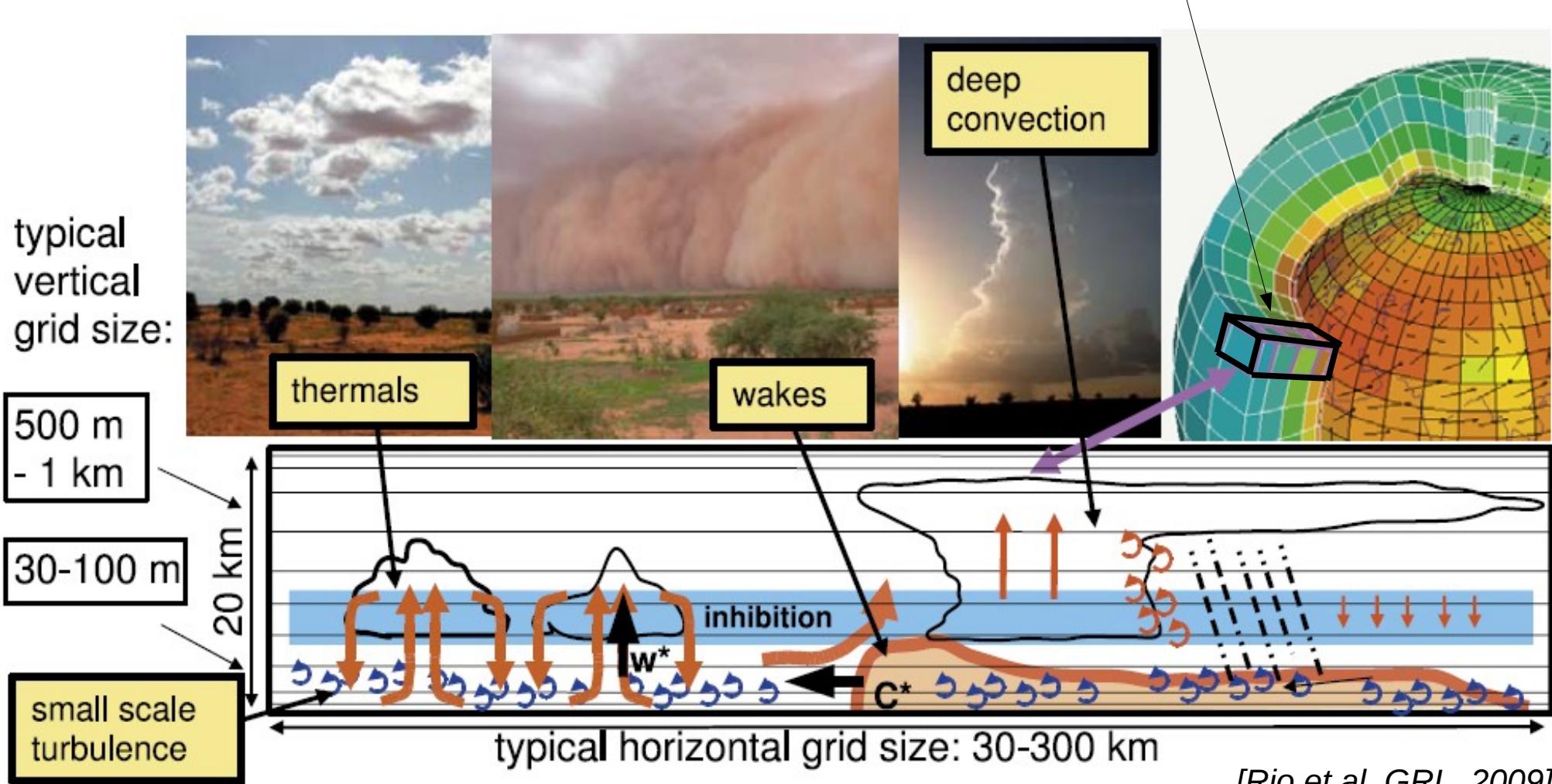
General Circulation Models

- Developed in the 60s for the purpose of weather forecast
- Based on a discretized version of the « primitive equations of meteorology »
- On the Earth but also very rapidly on other planets
- A number of important processes are subgrid scale and must be parameterized

Modeling of unresolved scales

Development of parameterization

A typical vertical atmospheric column



[Rio et al, GRL, 2009]

Typical time step : a few minutes to half an hour

Parameterization development and the use of high resolution explicit models



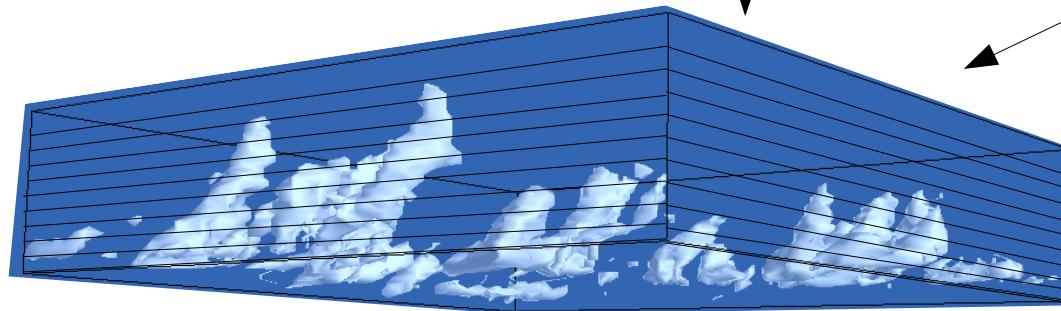
Test case, field campaign experiment



Explicit simulations, Grid cell, 20-100 m

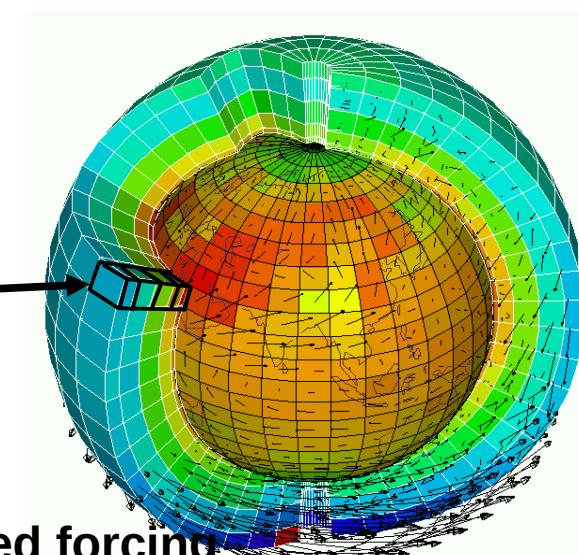
Observation

Evaluation



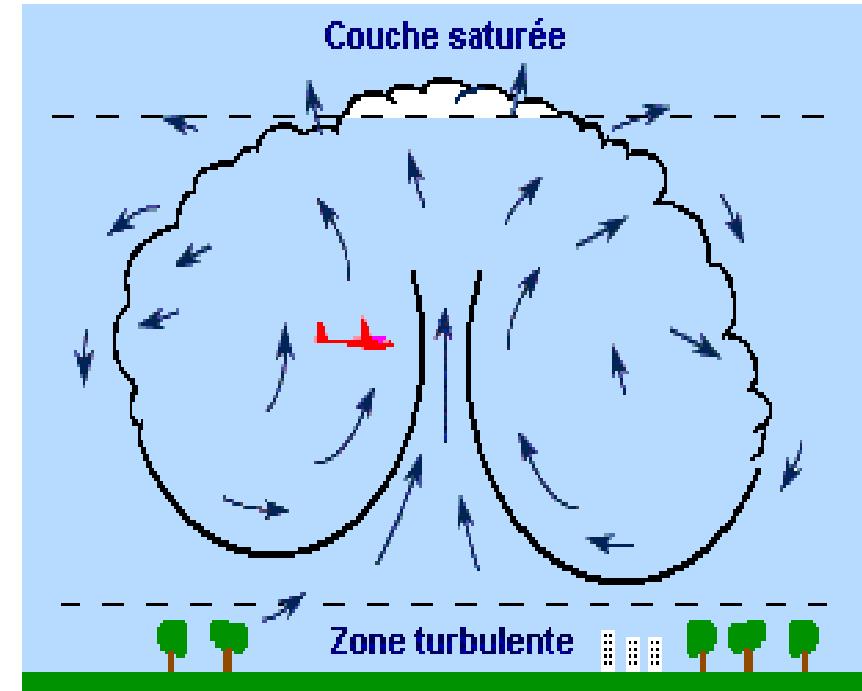
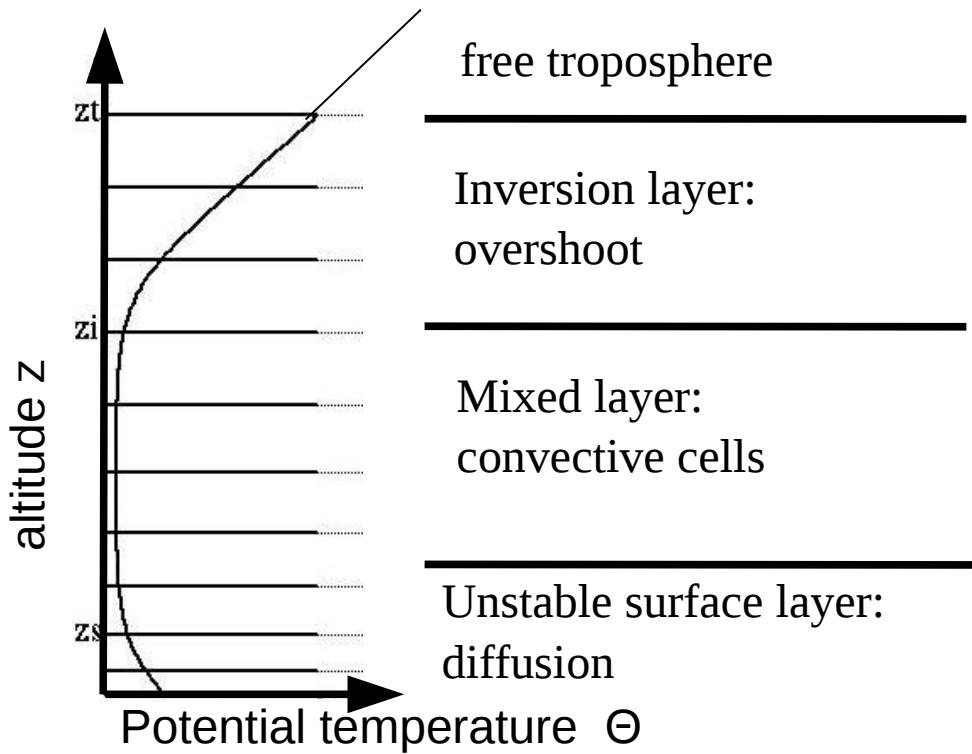
Climate model, parameterizations, « single-column » mode

« Large scale » conditions imposed



- Parameterizations are evaluated against other models
- Can be done for realistic test cases but also with more idealized forcing
(check the response of the parameterization to perturbations)

Convective boundary layer clouds

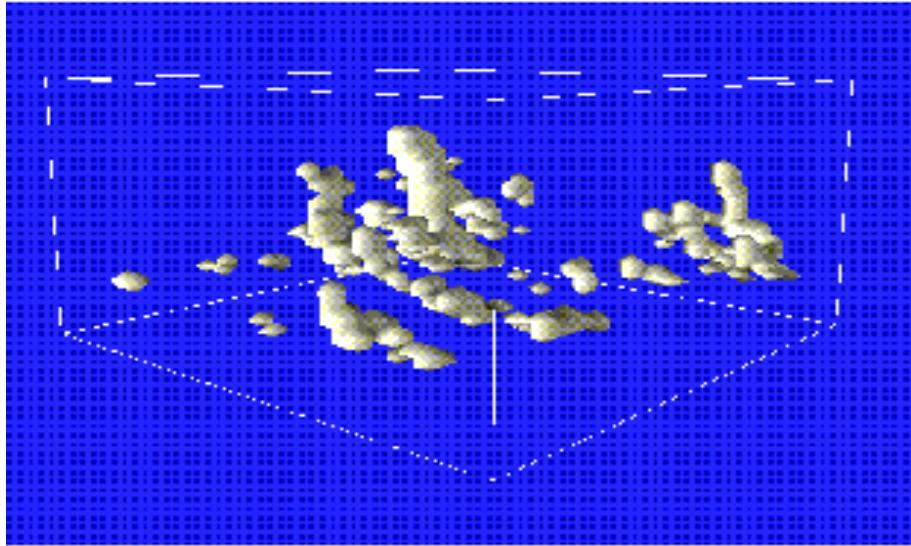


<http://www.astrosurf.org/lombry/meteo-vol-a-voile.htm>

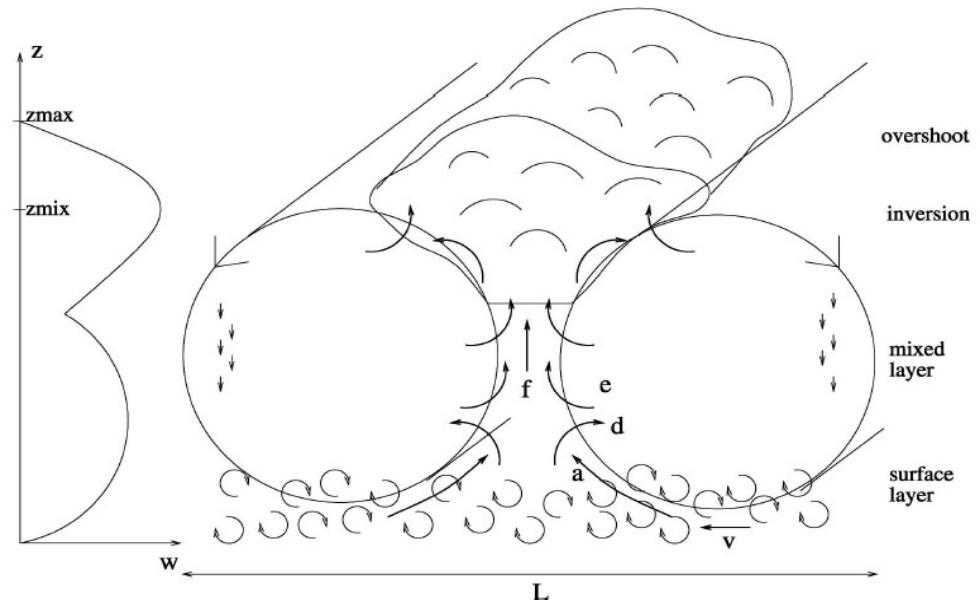


Convective boundary layer clouds

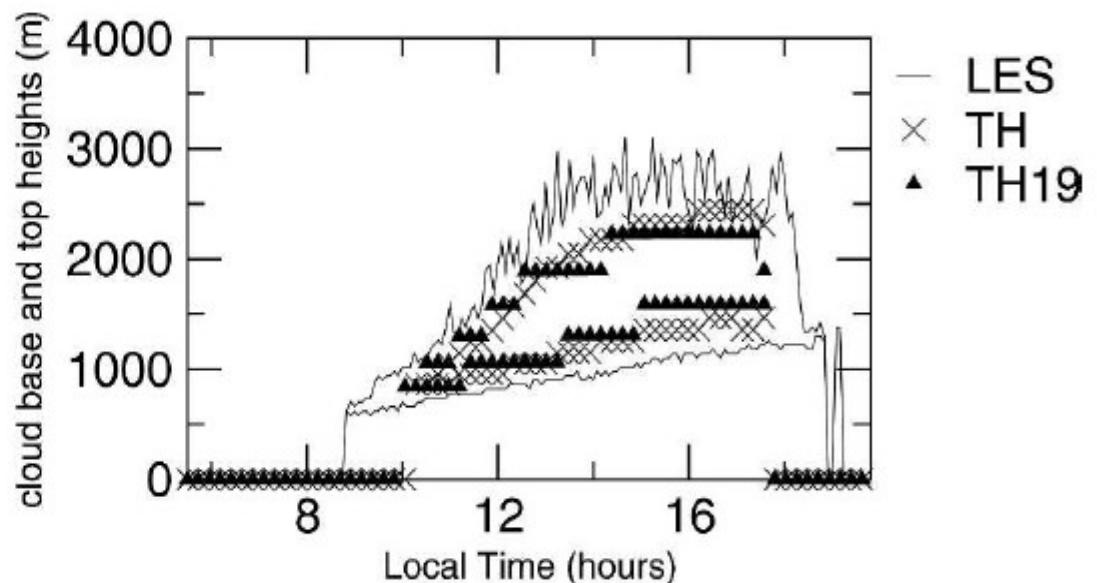
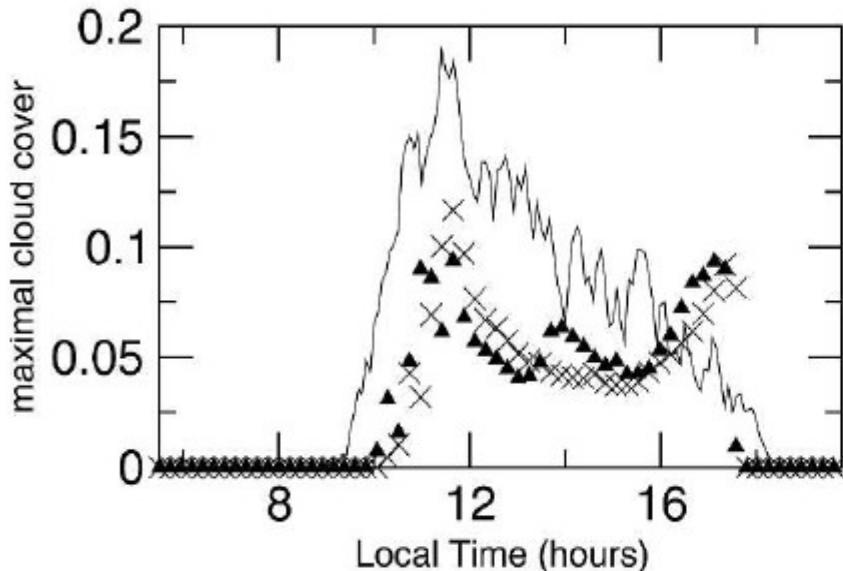
Reference simulation (LES model)



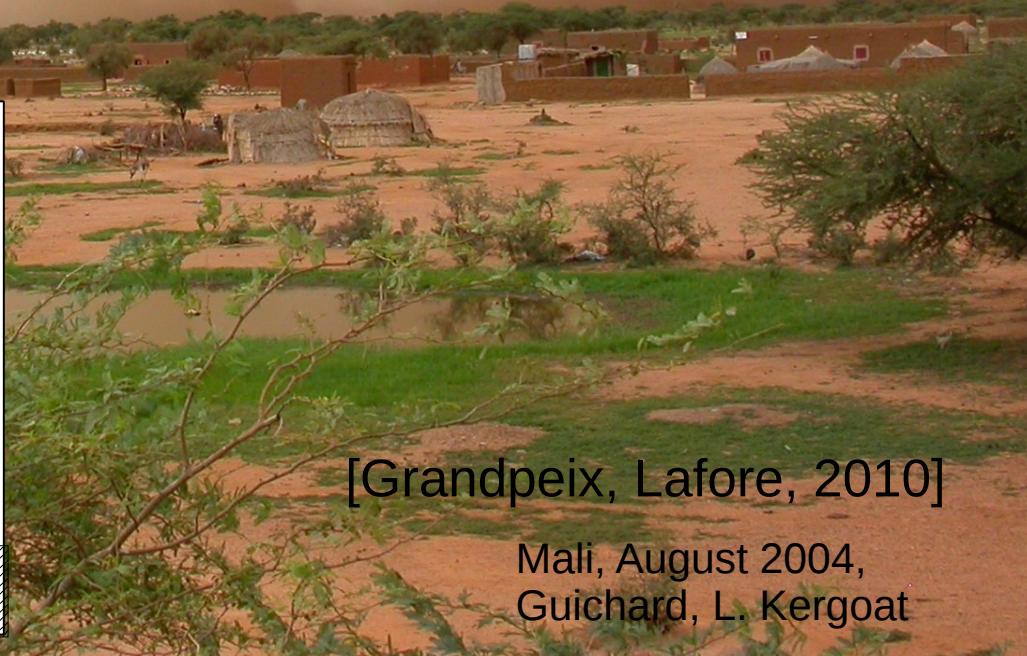
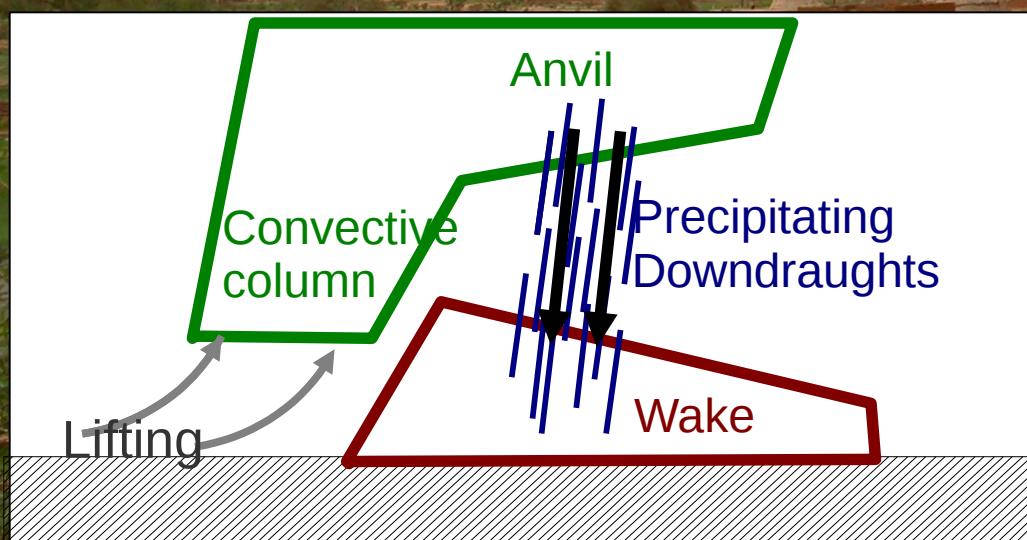
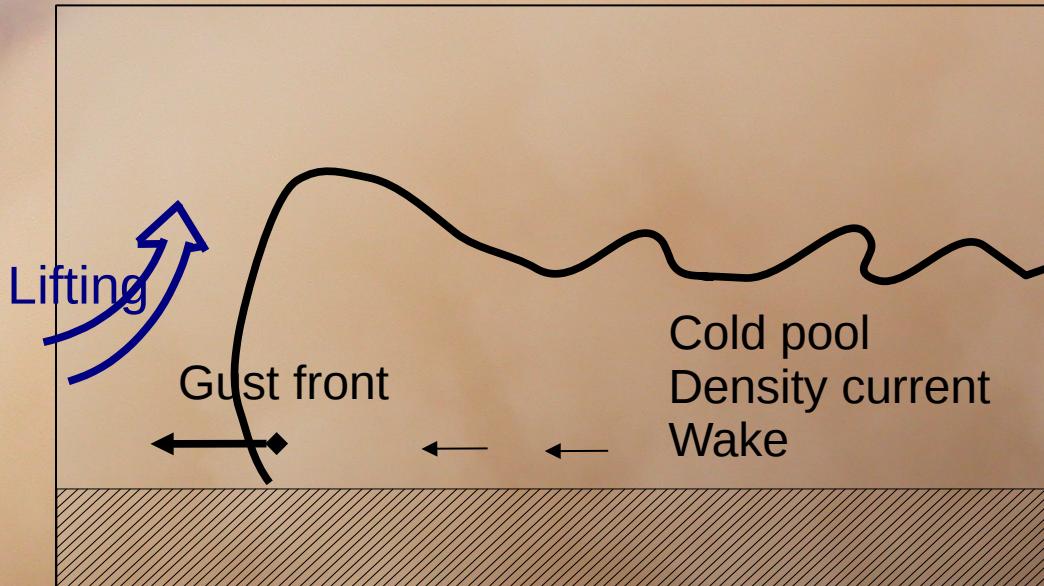
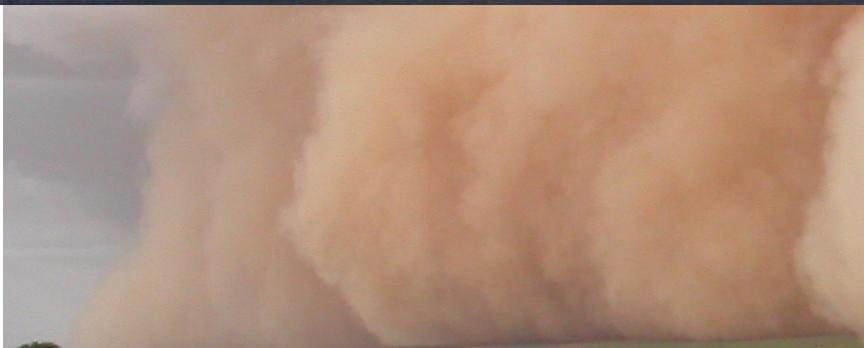
Heuristic model for the parameterization



Comparison between reference model and single column model fir case studies



Deep convection and wakes

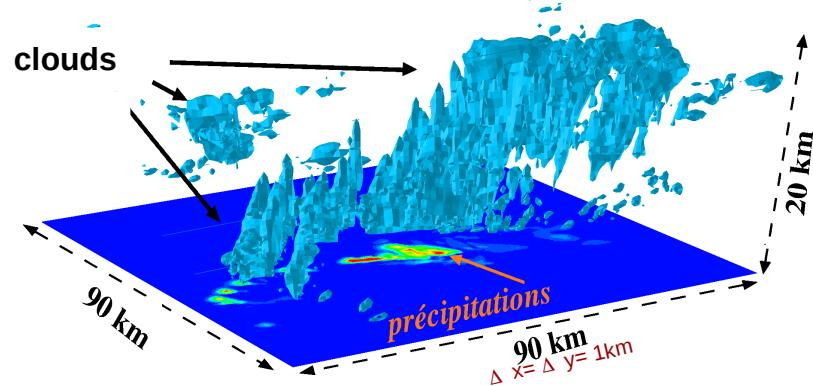


[Grandpeix, Lafore, 2010]

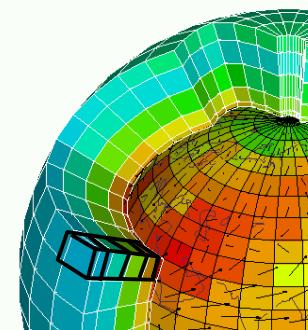
Mali, August 2004,
Guichard, L. Kergoat

A multi-step evaluation, from 1D to 3D, with a progressive relaxation of constraints

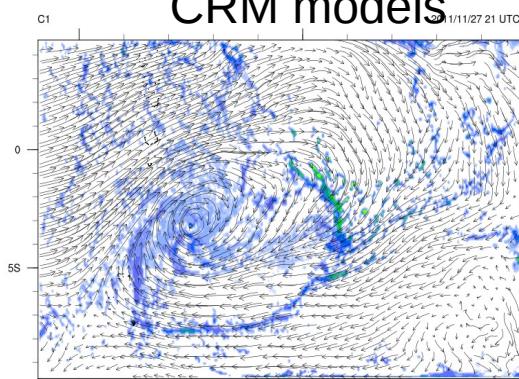
LES or CRM models



Single column model

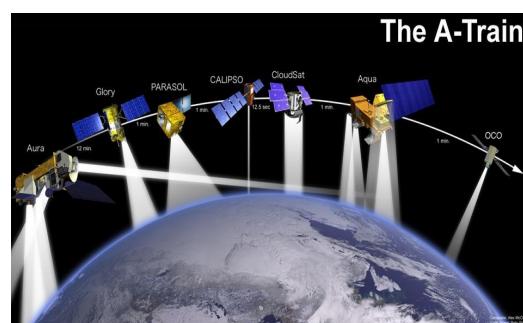
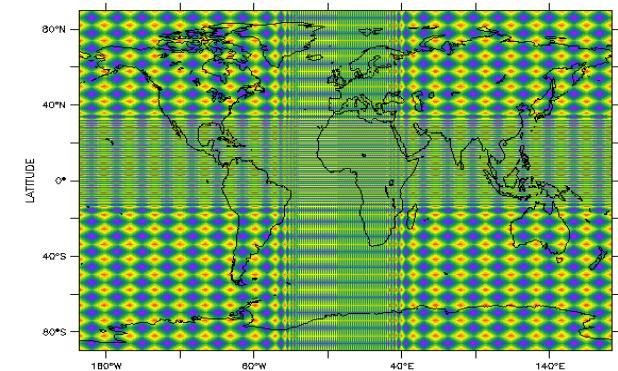


Large domain
CRM models



Same large scale atmospheric profile

Same large scale circulation



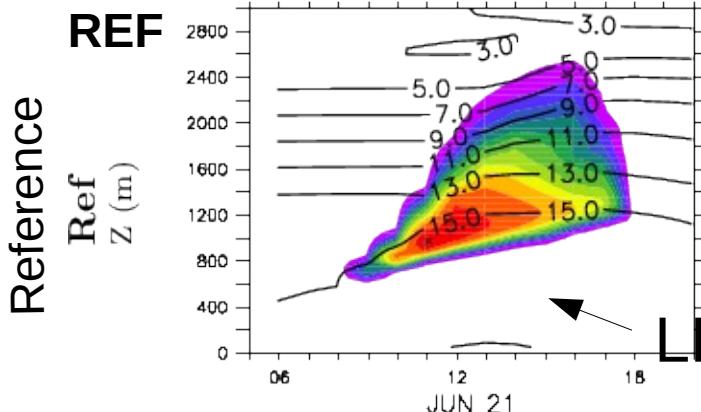
The A-Train
Free or prescribed circulation

Le modèle du thermique et l'amélioration (robuste 1D&3D) des nuages bas

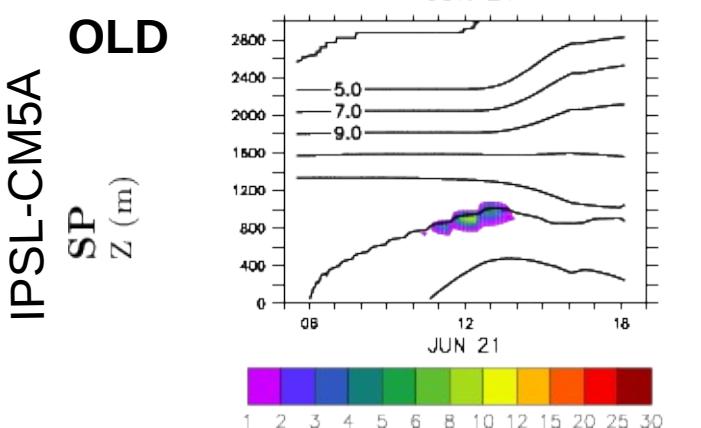
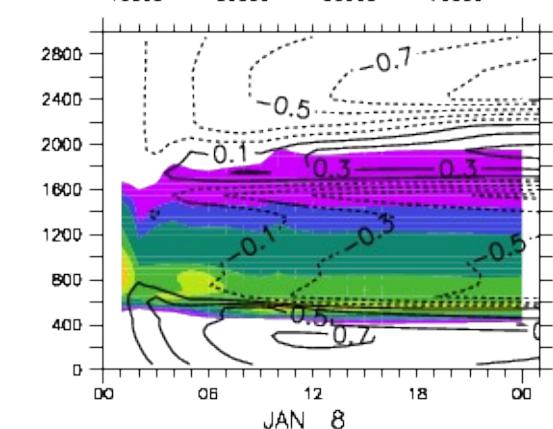
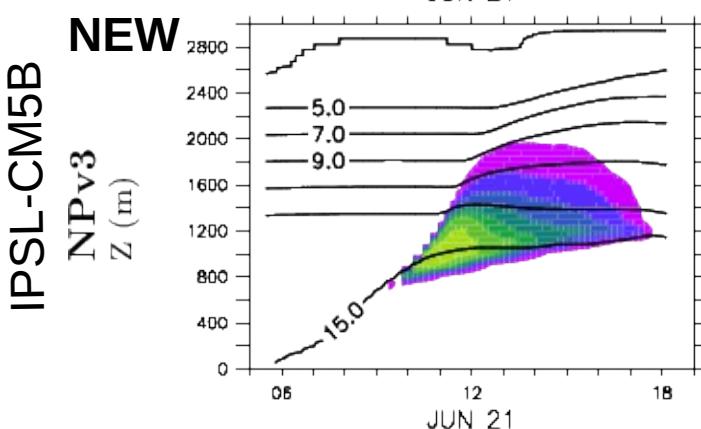
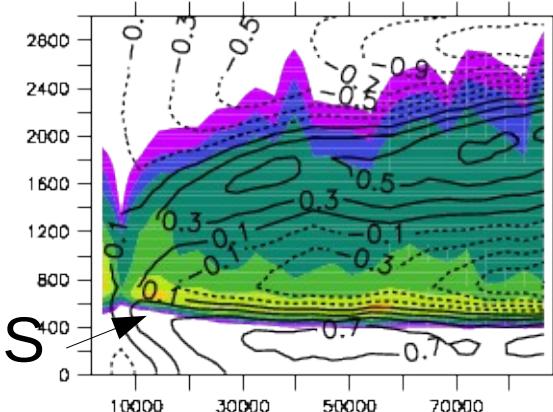
Cas test 1D

nébulosité (%) et vapeur d'eau (g/kg)

Eurocs Cumulus



Rico



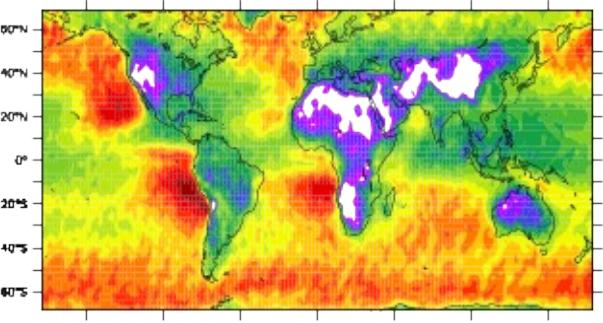
0.2 0.5 1 3 5 7 10 15 20

Simulations 3D

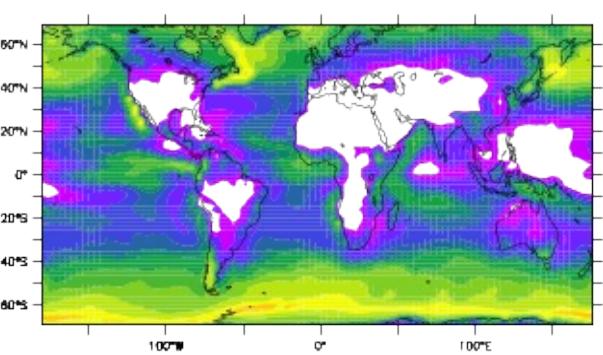
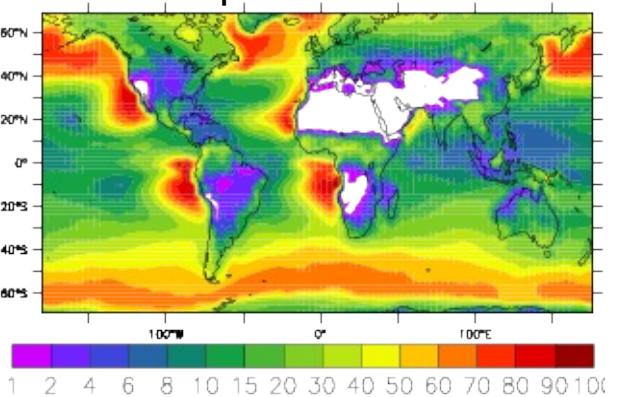
Couverture nuages bas (%)

Moyenne annuelle

Calipso

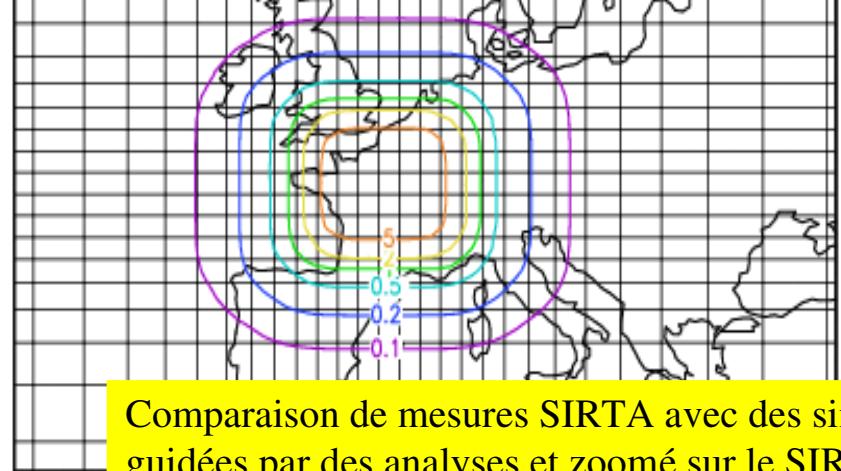


Utilisant le simulateur Cosp
Pour comparer modèle et satellite



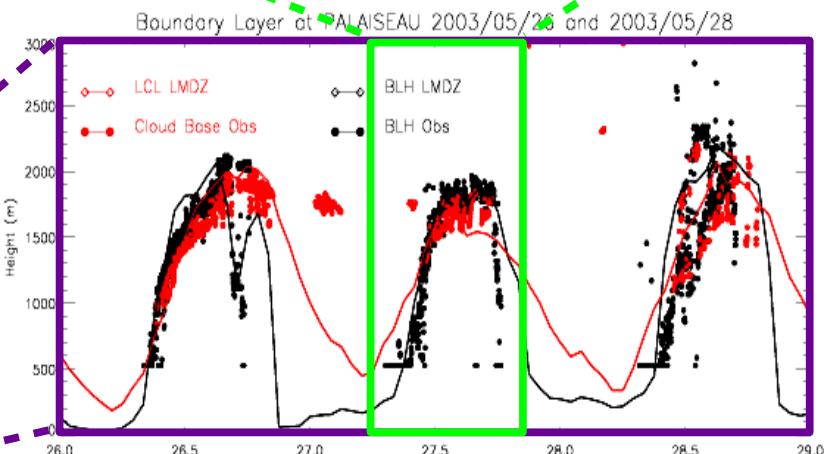
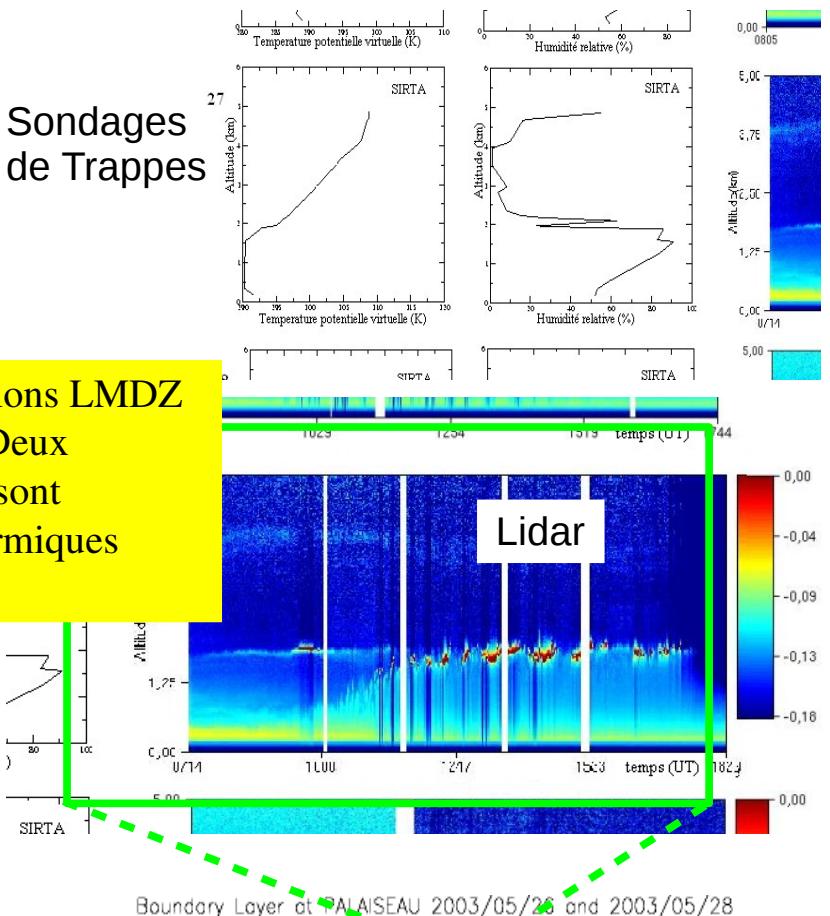
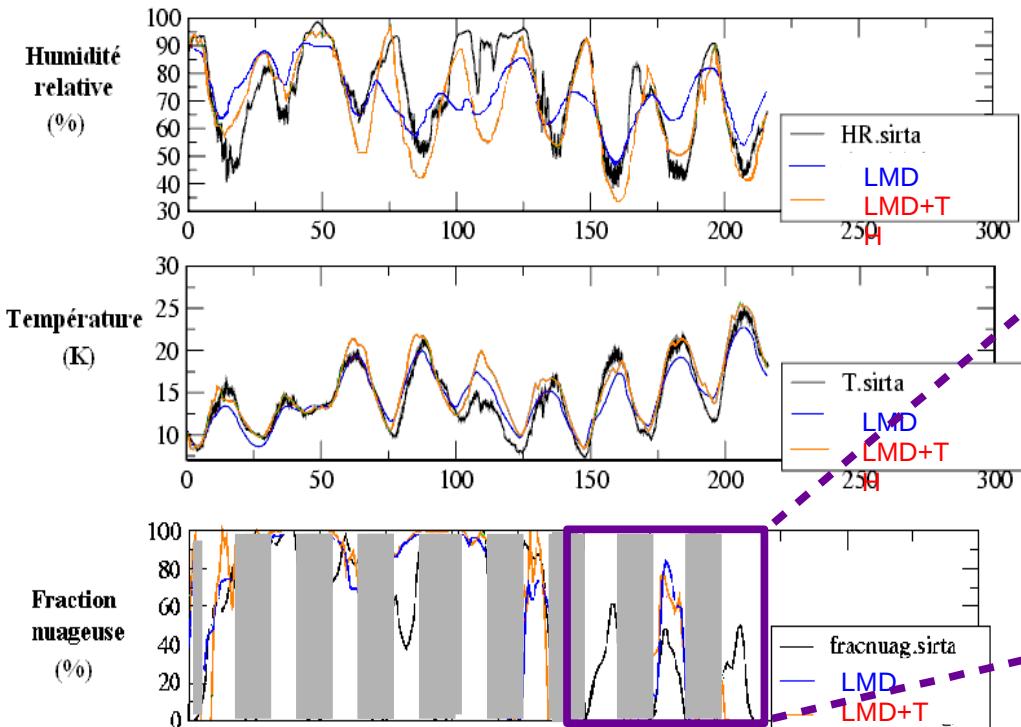
SIRTA observatory a node for national and international networks



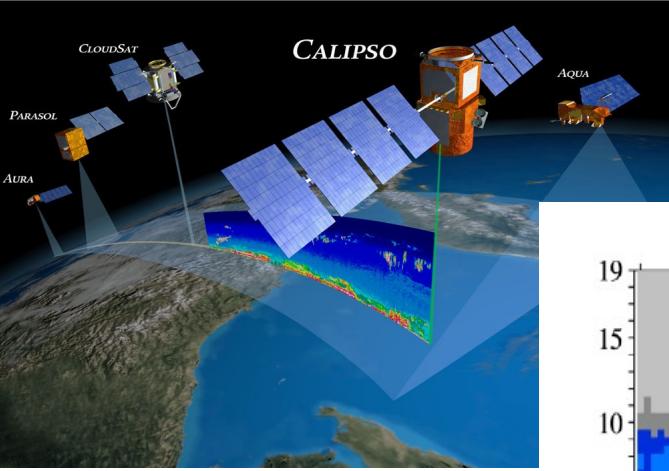


Sondages de Trappes

Comparaison de mesures SIRTA avec des simulations LMDZ guidées par des analyses et zoomé sur le SIRTA. Deux versions de la paramétrisation de la couche limite sont utilisées : couche limite du LMD avec ou sans thermiques (TH).

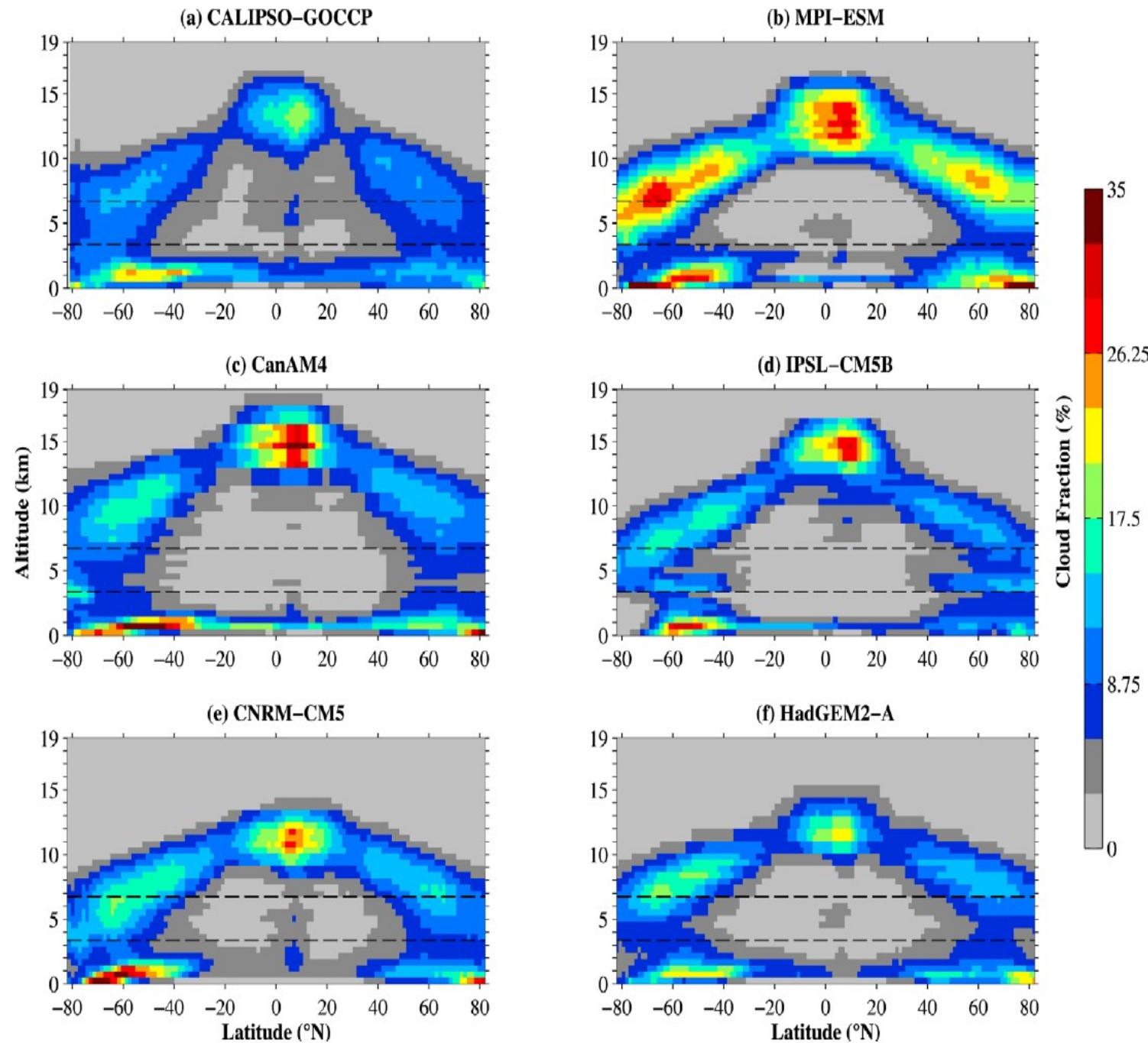


Hauteurs de couche limite et niveau de condensation
Diagnostics effectués par Anne Mathieu



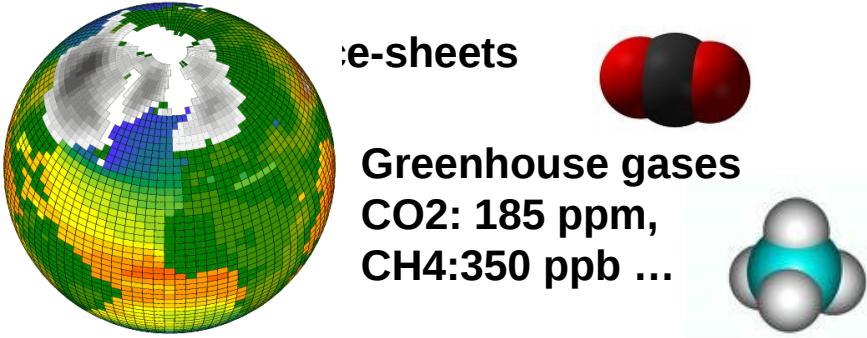
Global scale, satellite observation

Zonal mean of the cloud vertical fraction observed by Calipso et simulated by models + obs.
Simulator COSP



Land-sea contrasts and polar amplification in past and future climates

Last Glacial Maximum main forcings



LGM climate reconstructions

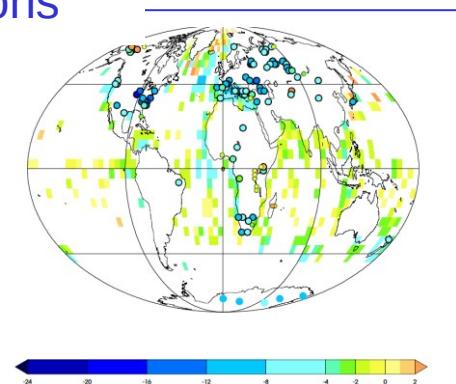
Land data

(pollen and plant macrofossils):
Bartlein et al, Clim Dynam 2011

Ocean data (multi proxy):
MARGO, NGS 2009

Ice-core data:

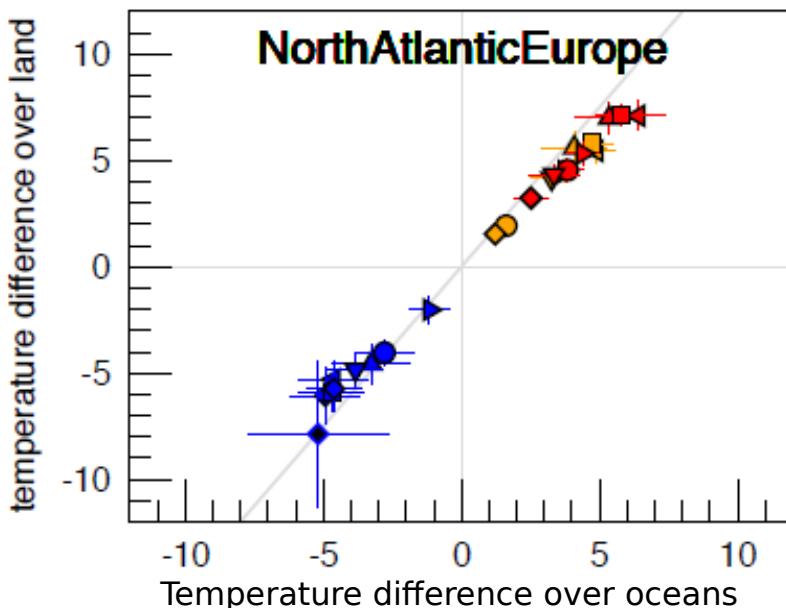
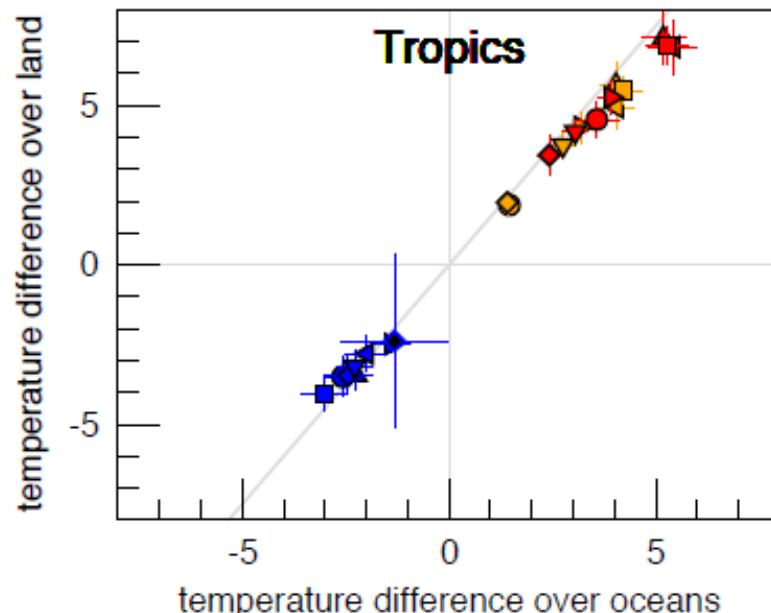
Masson-Delmotte et al pers. comm



Relationships between LGM vs higher CO₂ climates?
Are the large scale relationships stable? Can we evaluate them from paleodata ?

Example: Land sea contrasts

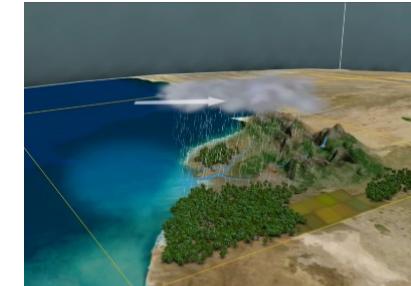
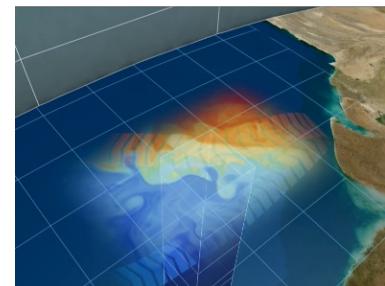
Note: all model averages calculated from grid points where LGM data is available



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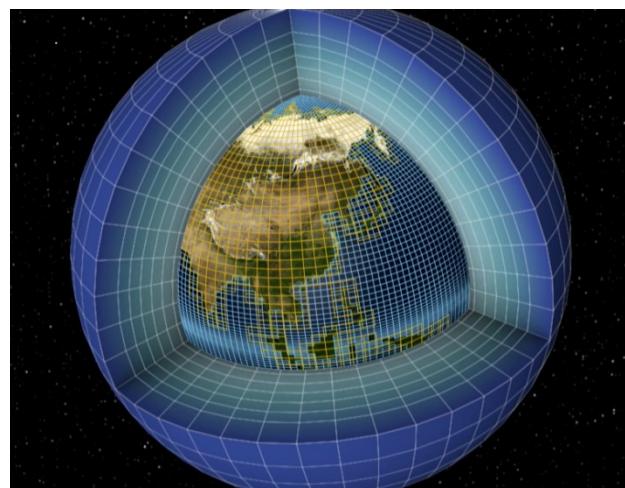
The IPSL Earth System Model



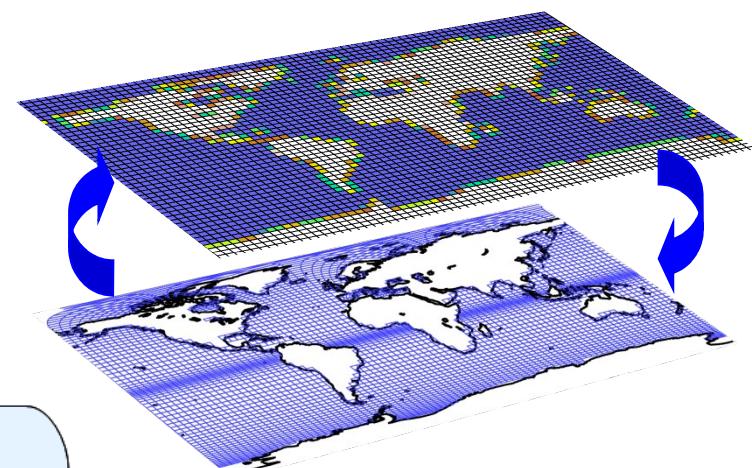
INCA / REPROBUS
(chimie atmosphérique)
(aérosol)

LMDZ
(atmosphère)

ORCHIDEE
(surfaces continentales)
(végétation)



OASIS
(coupleur)



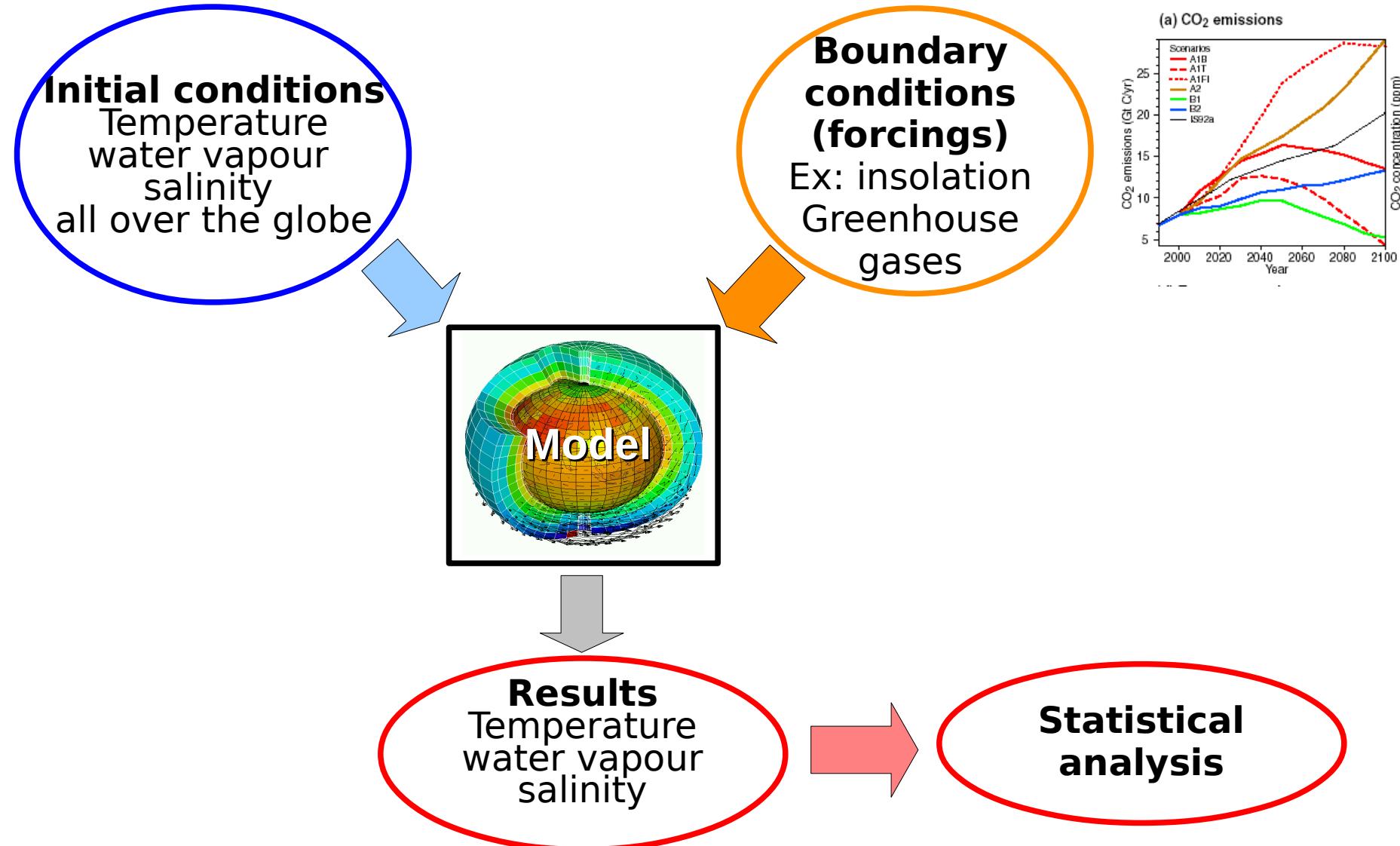
OPA
(océan)

LIM
(glace de mer)

NEMO

PISCES
(biogéochimie marine)

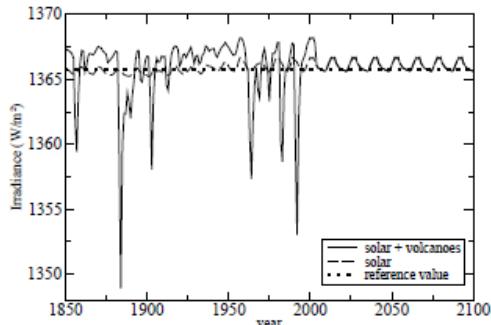
Climate simulations



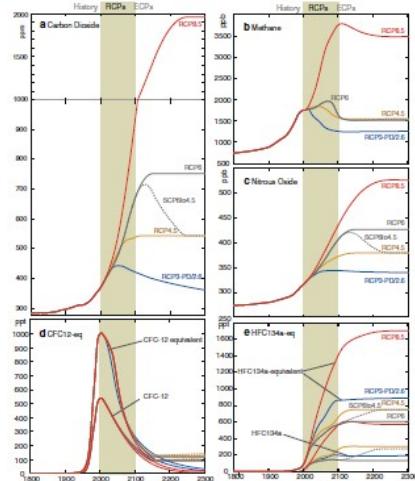
The IPSL Earth System Model

Natural and anthropogenic forcings

Solar and volcanoes

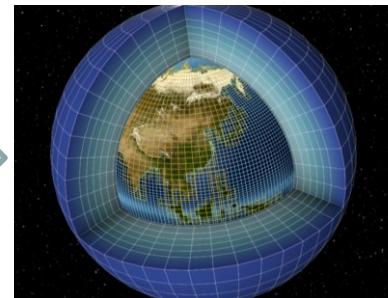


Green house gases and active gases

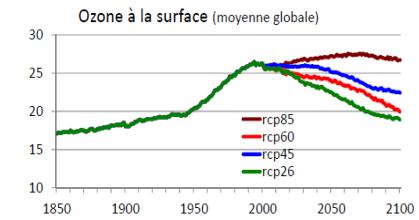


CO₂ concentration

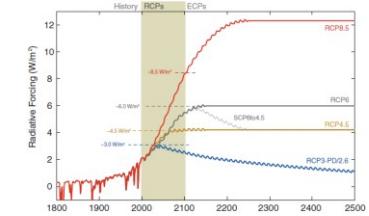
IPSL-CM5A-LR



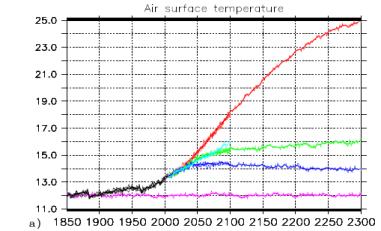
Atmospheric composition



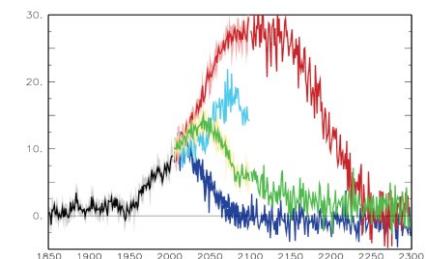
Radiative forcings



Climate changes

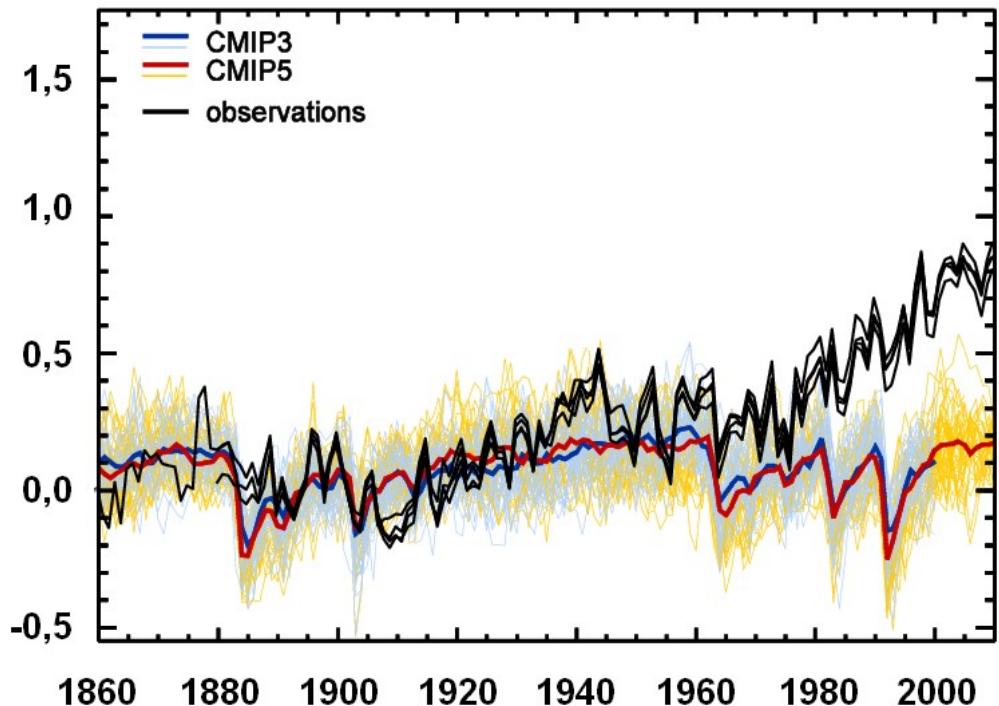


Authorized CO₂ emissions

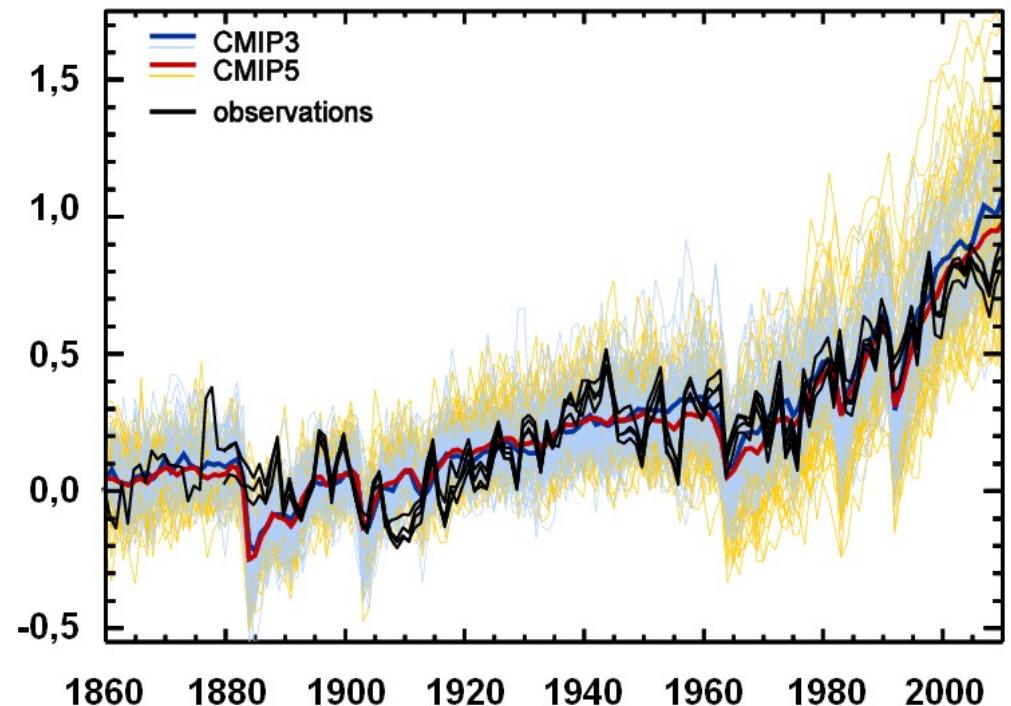


Recent Earth surface temperature trend

Simulations with natural forcings only

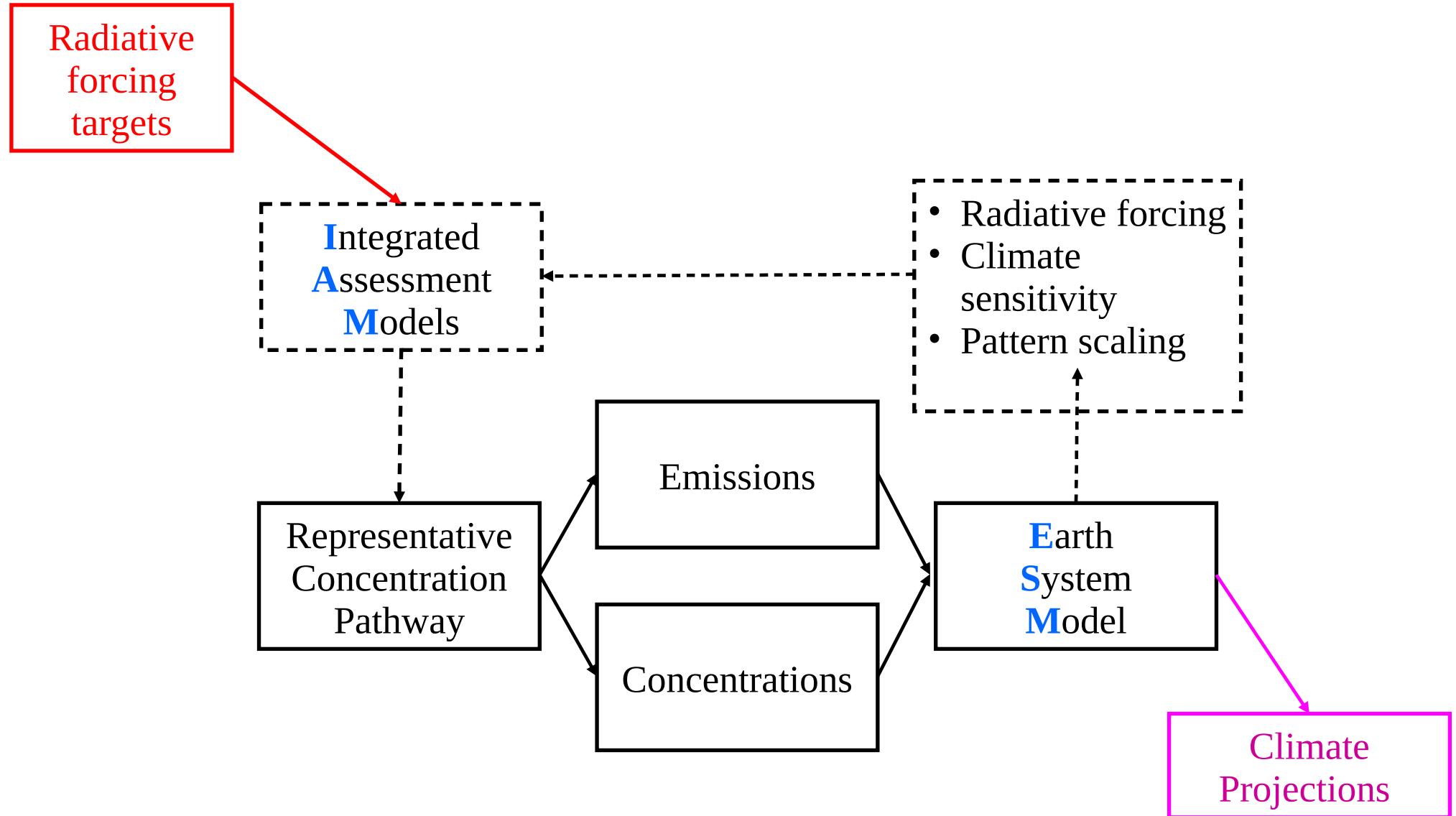


Simulations with natural and anthropological forcings



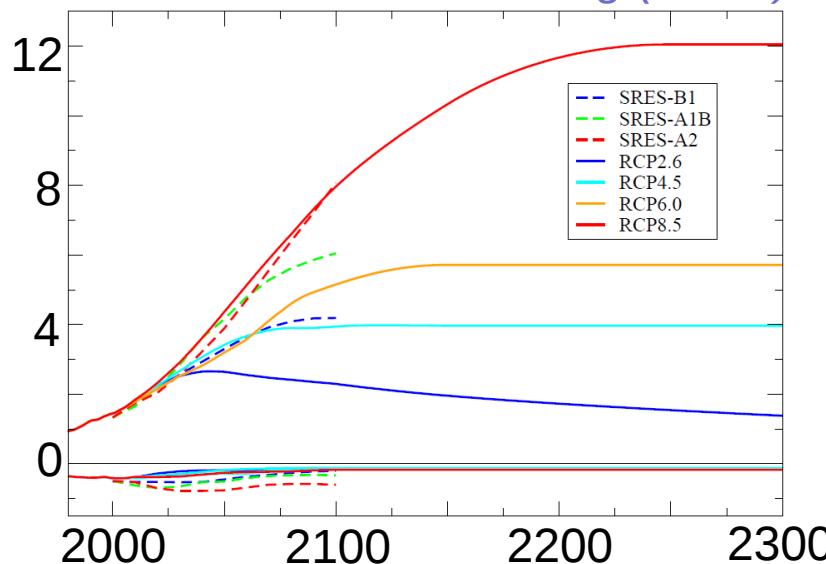
[IPCC, 2013]

Scenario for future climate change projections

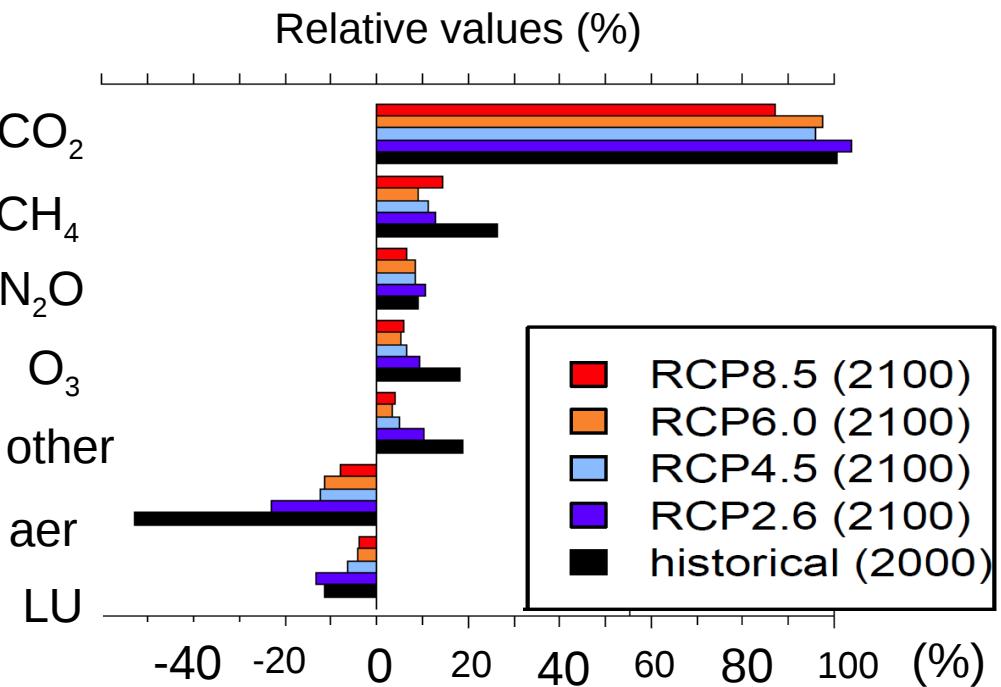
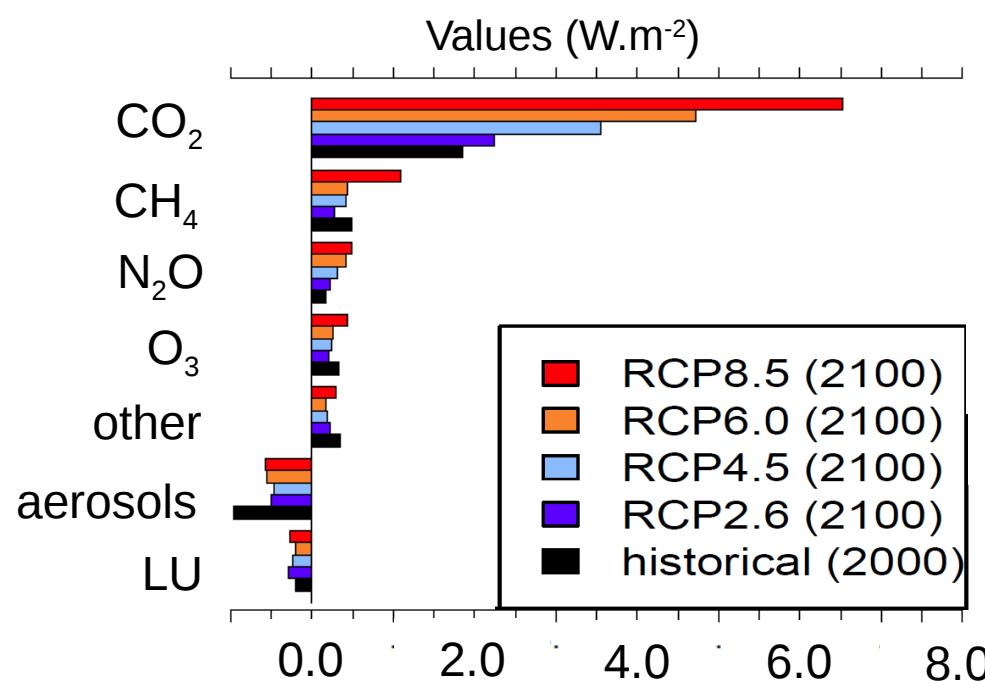


Radiative forcing of future scenarios

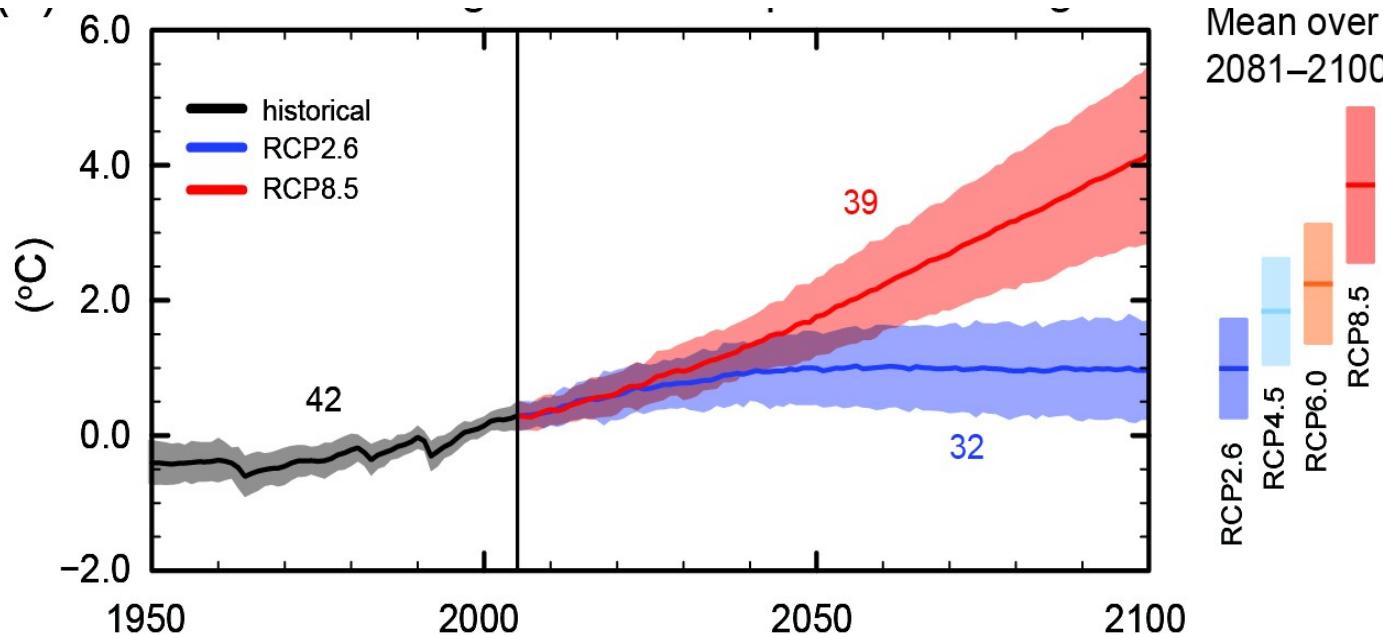
Total radiative forcing (W.m⁻²)



Contribution of individual forcings to total forcing relative to 1850



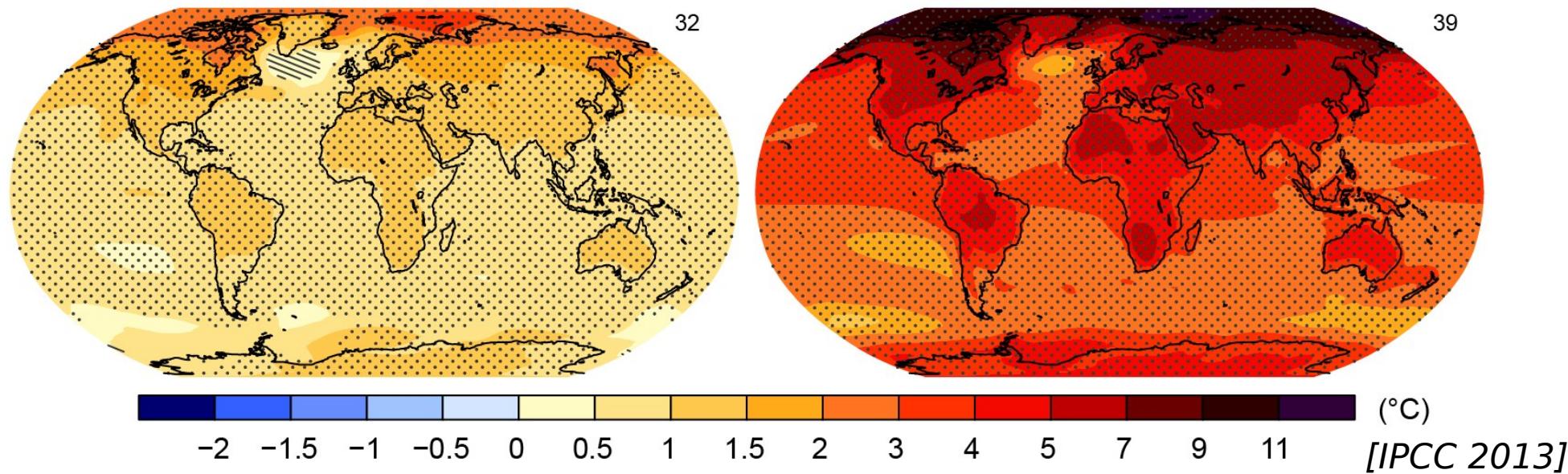
Global mean surface temperature change



RCP 2.6

RCP 8.5

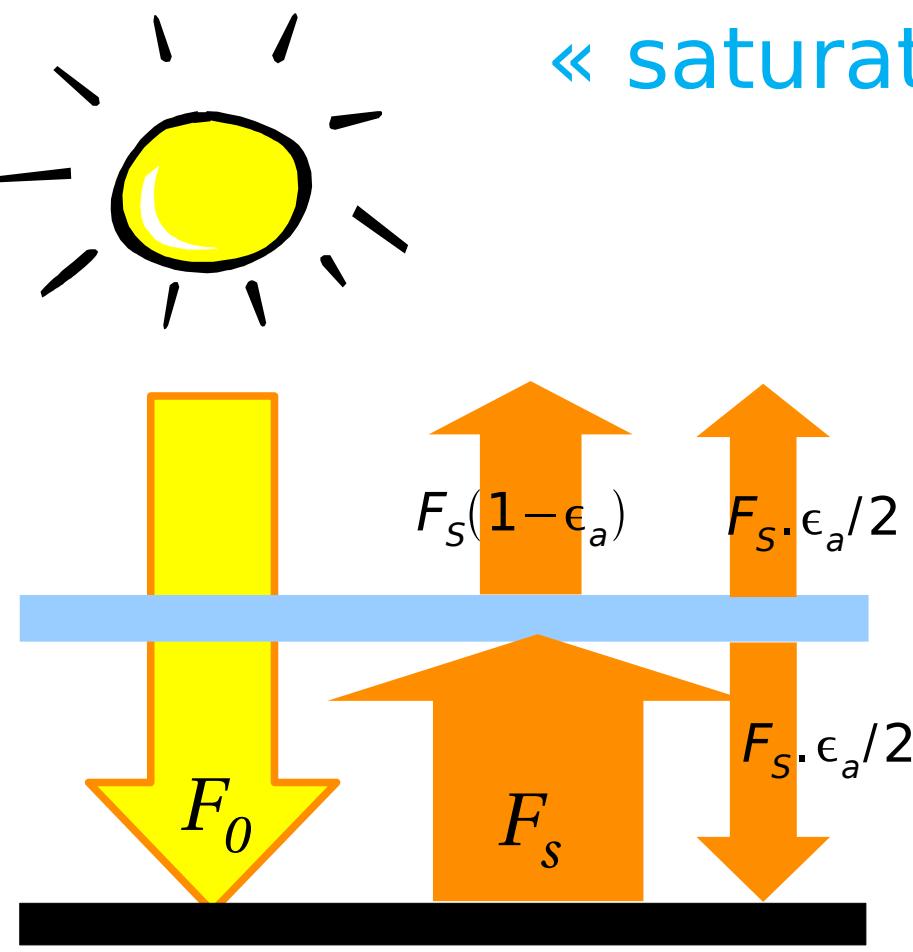
Change in average surface temperature (1986–2005 to 2081–2100)



Outlook

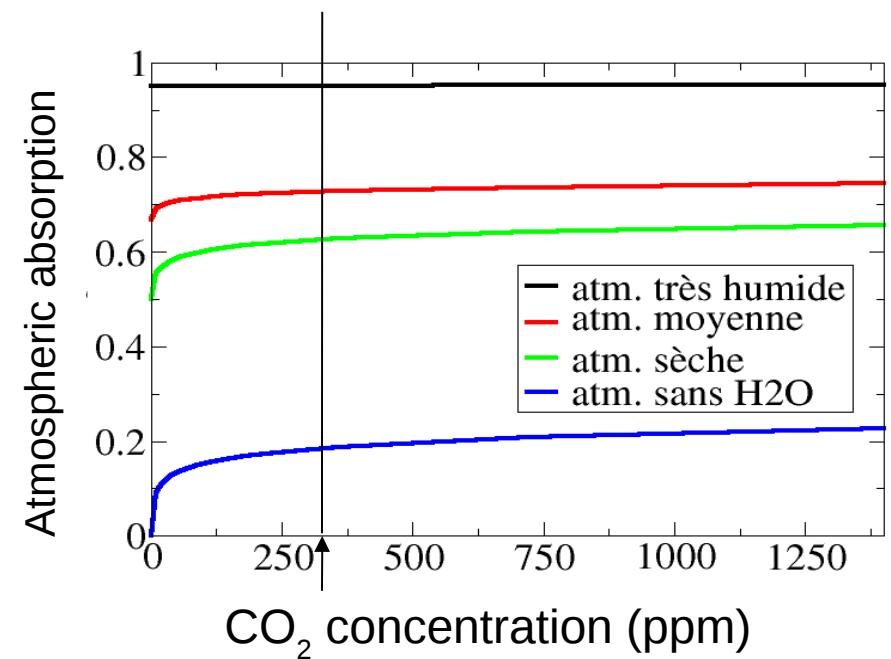
- I. Emergence of climate and climate change science
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The CO₂ greenhouse effect and the « saturation » paradox



$$\sigma T_s^4 = \frac{F_0}{1 - \epsilon_a / 2}$$

Mean atmospheric absorption in the infrared

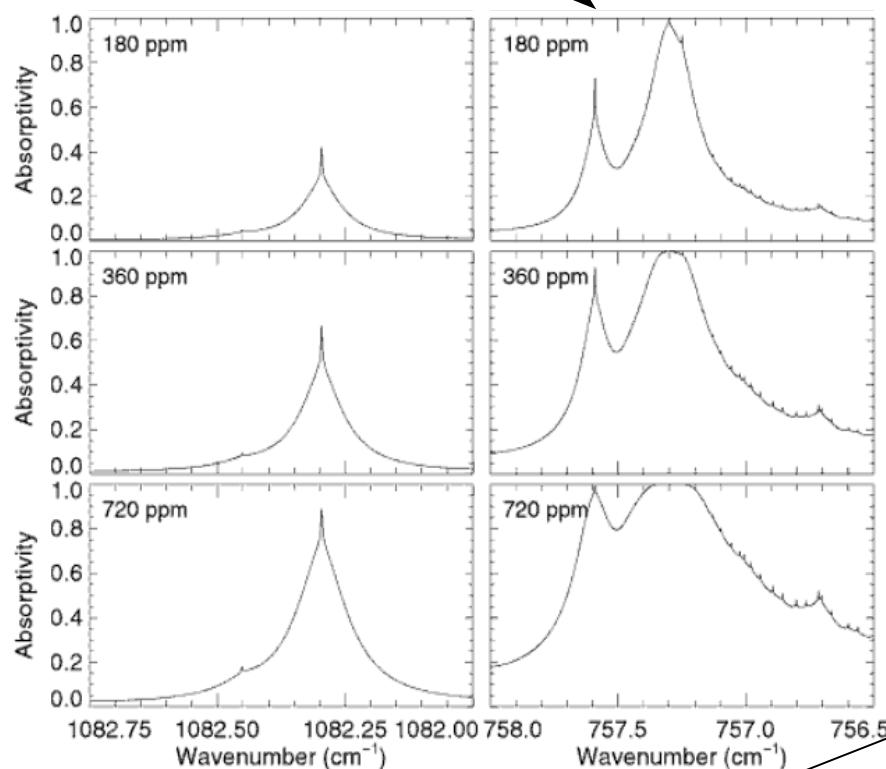


How can the greenhouse effect increase if the atmospheric absorption don't?

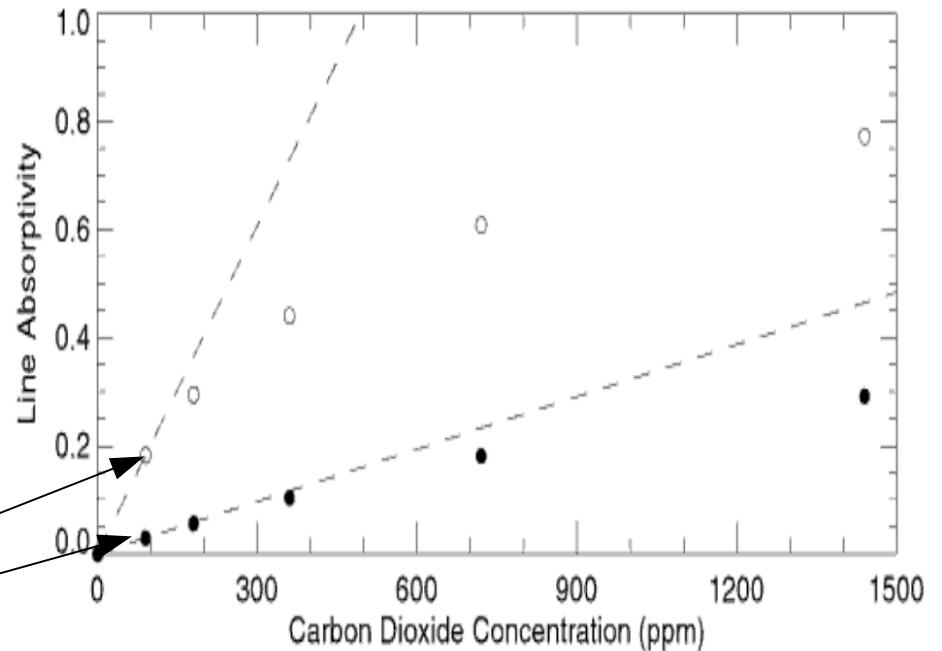
Saturation of absorption bands

Absorption by CO₂, for a vertical column of atmosphere

Absorption spectra,
for 3 CO₂ concentrations and two
narrow bands

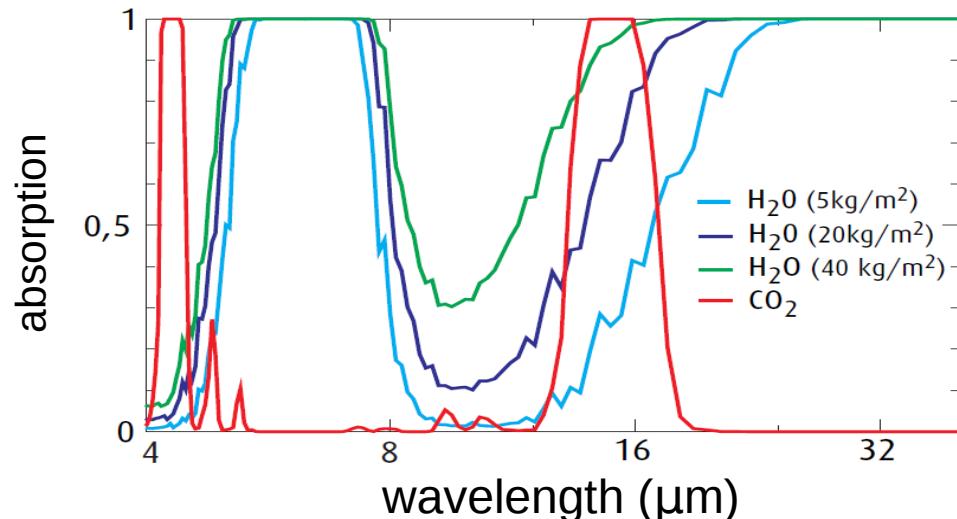


Total absorption of the two narrow bands

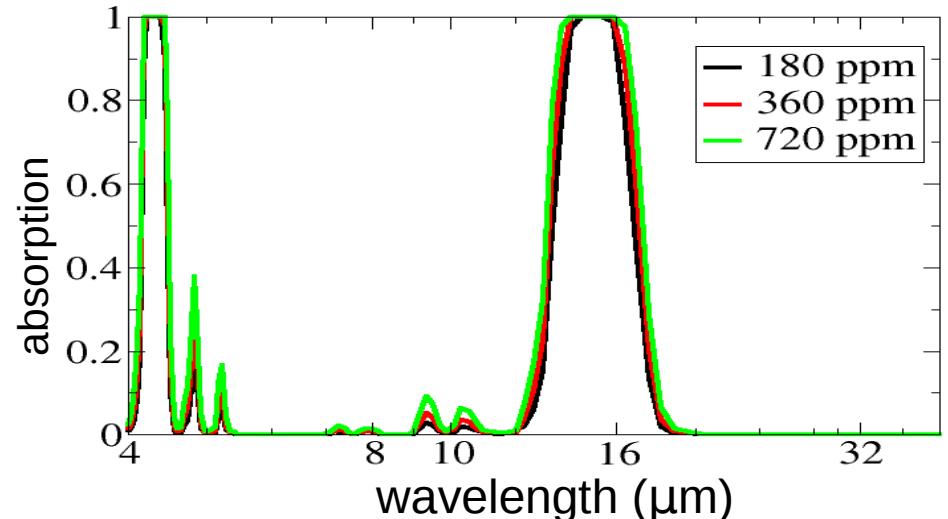


Infrared absorption of the atmosphere

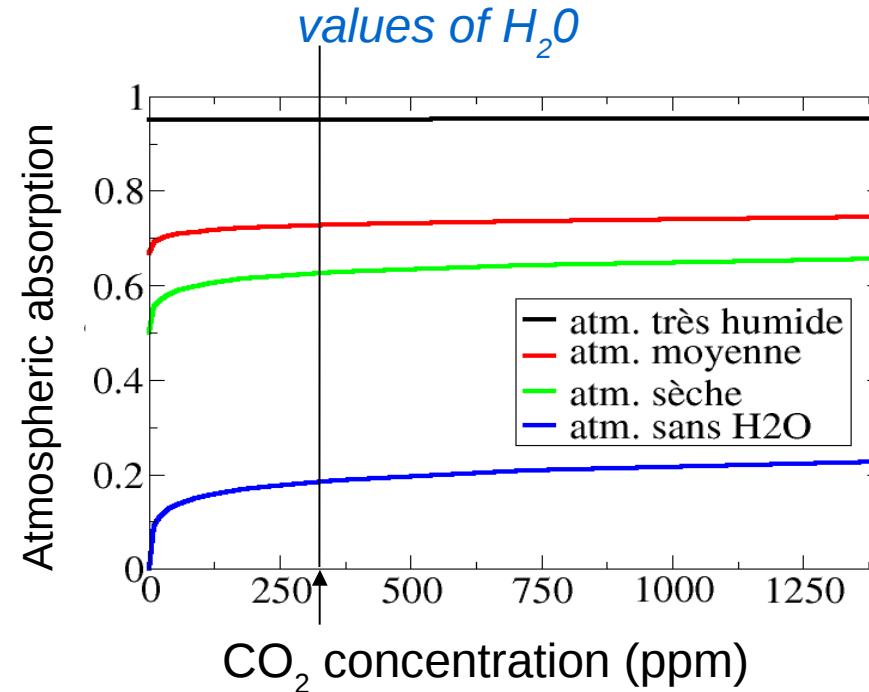
Different H_2O concentration



Different CO_2 concentration

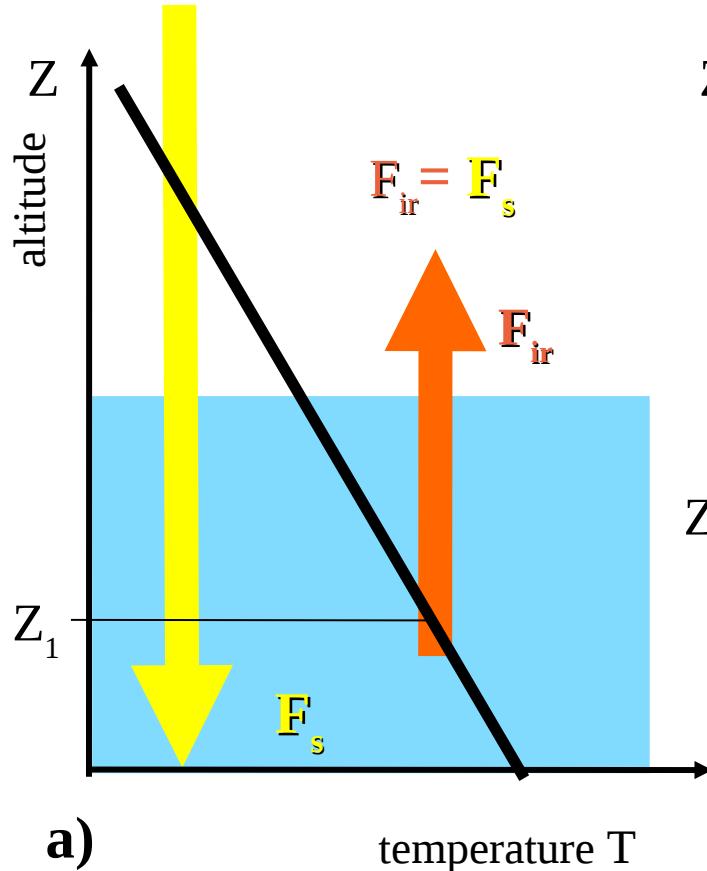


Infrared absorption of the atmosphere as a function of CO_2 , for different values of H_2O

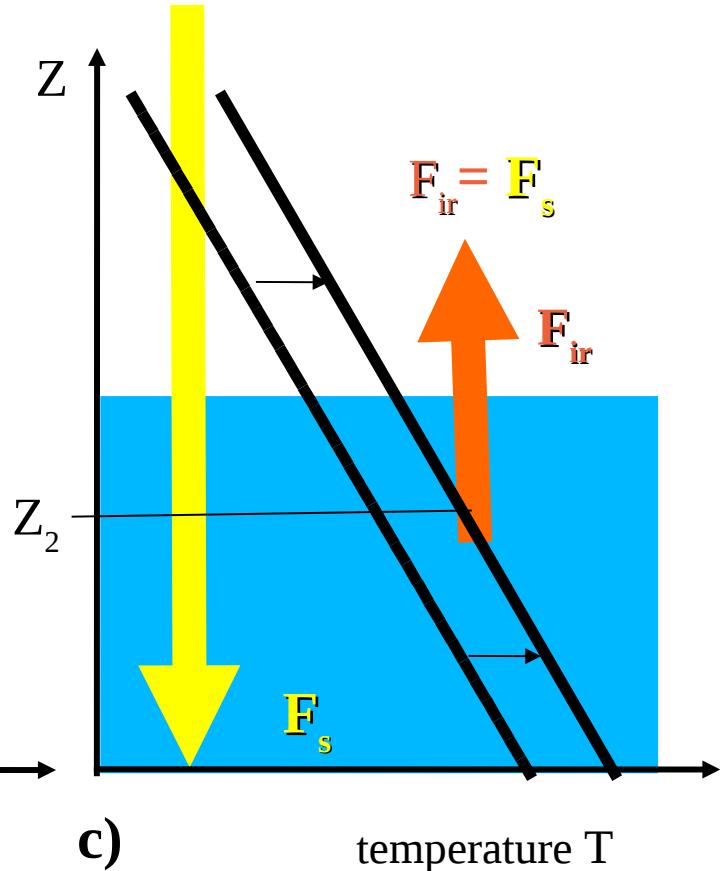
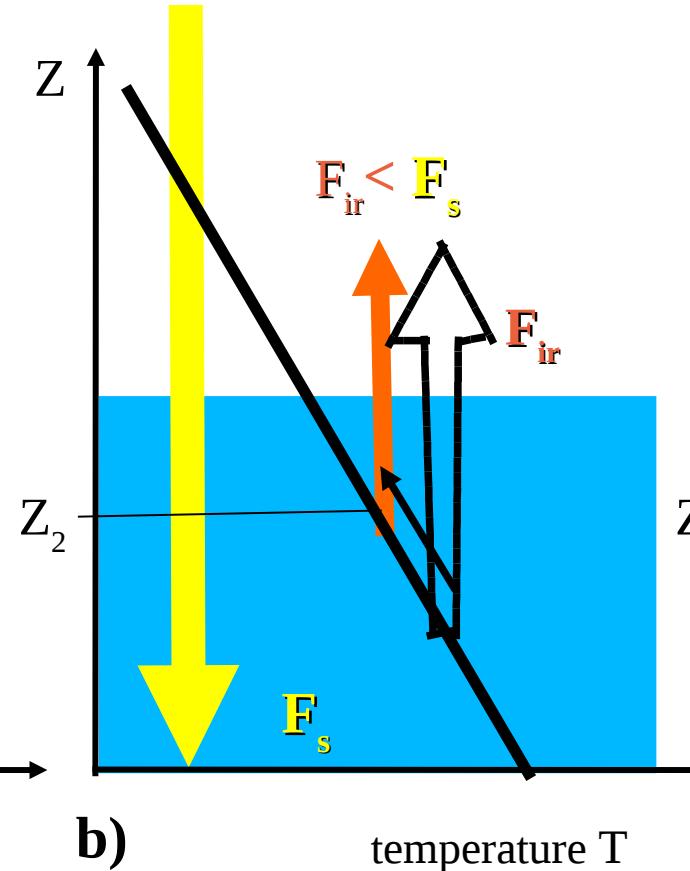


CO_2 increase and greenhouse effect

Net solar radiation F_s

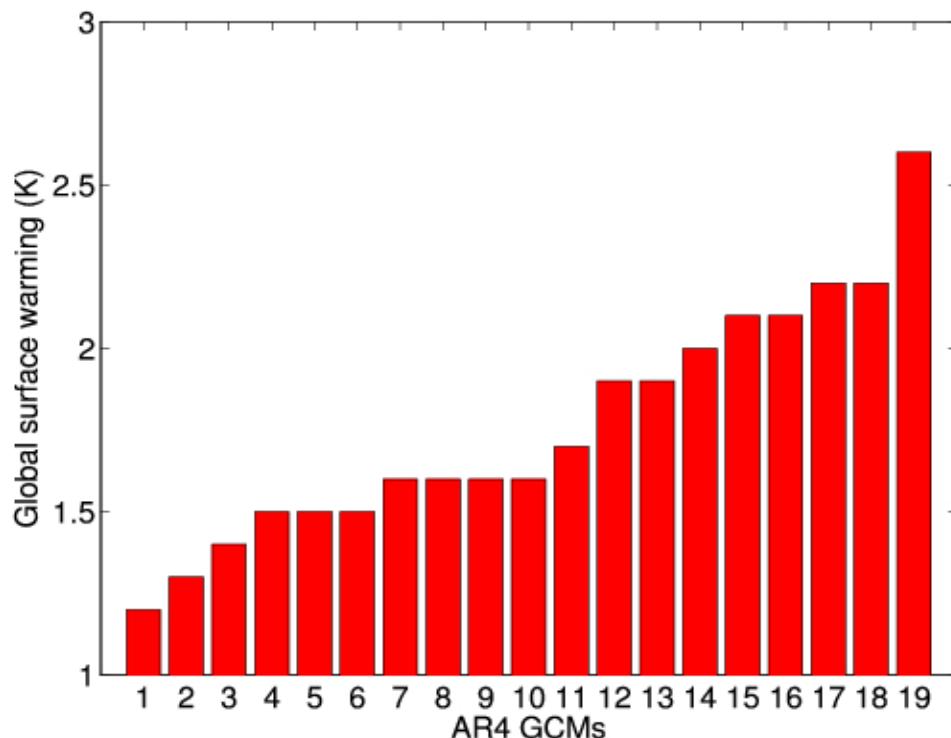


Outgoing longwave radiation F_{ir}

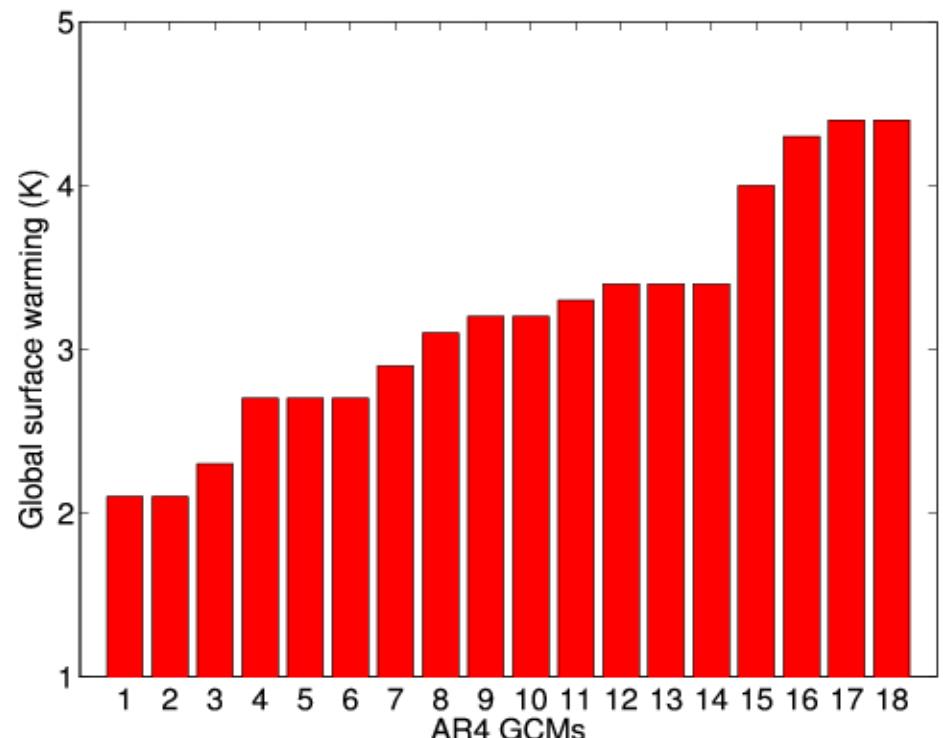


Why climate sensitivity estimate differs among models?

Transient Climate Response :
(1% CO₂/yr, transient warming at 2xCO₂)



Equilibrium Climate Sensitivity :
(warming for sustained 2xCO₂)

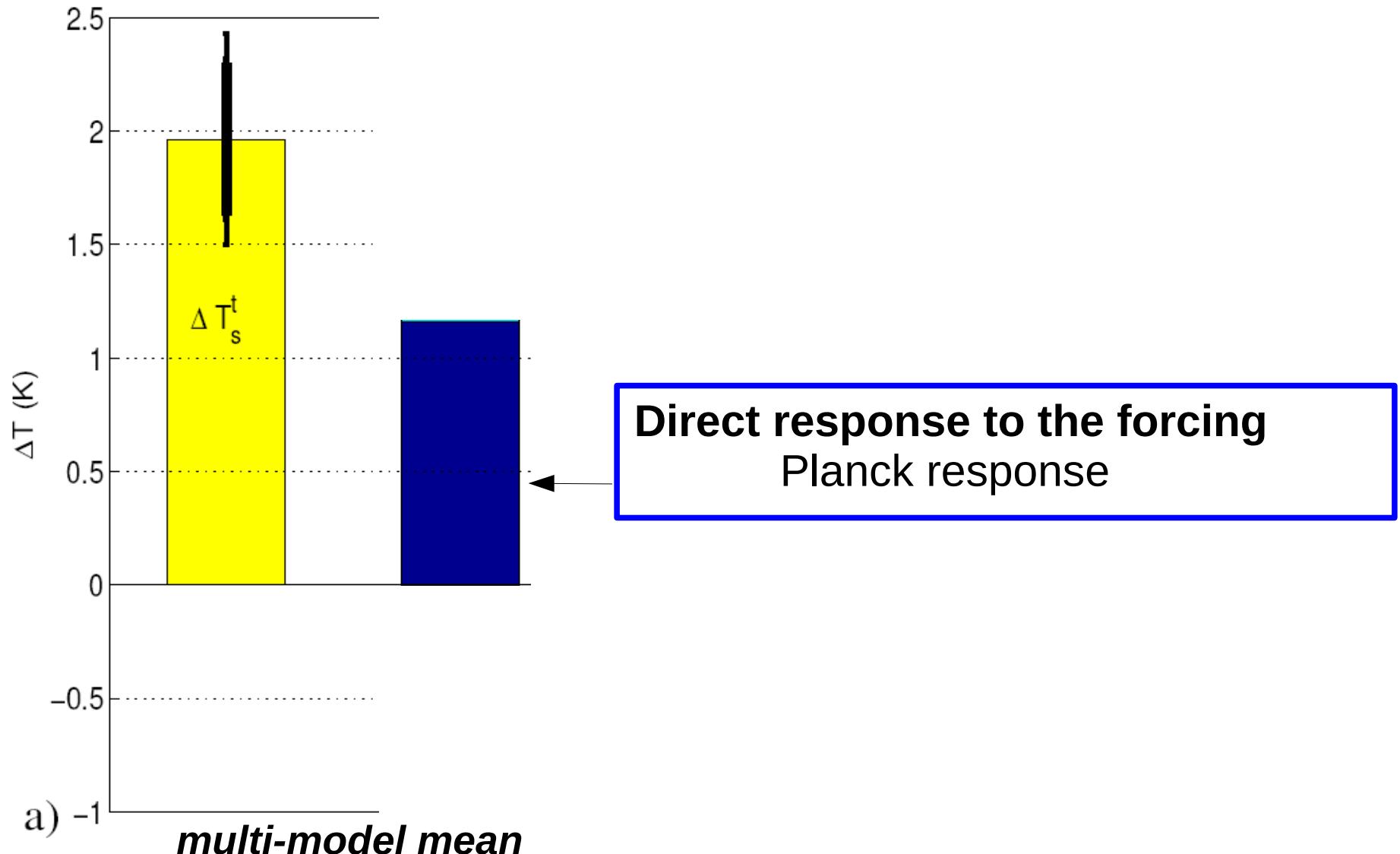


Climate sensitivity and TCR estimates depend on :

- radiative forcing
- climate feedbacks
- ocean heat uptake (transient only)

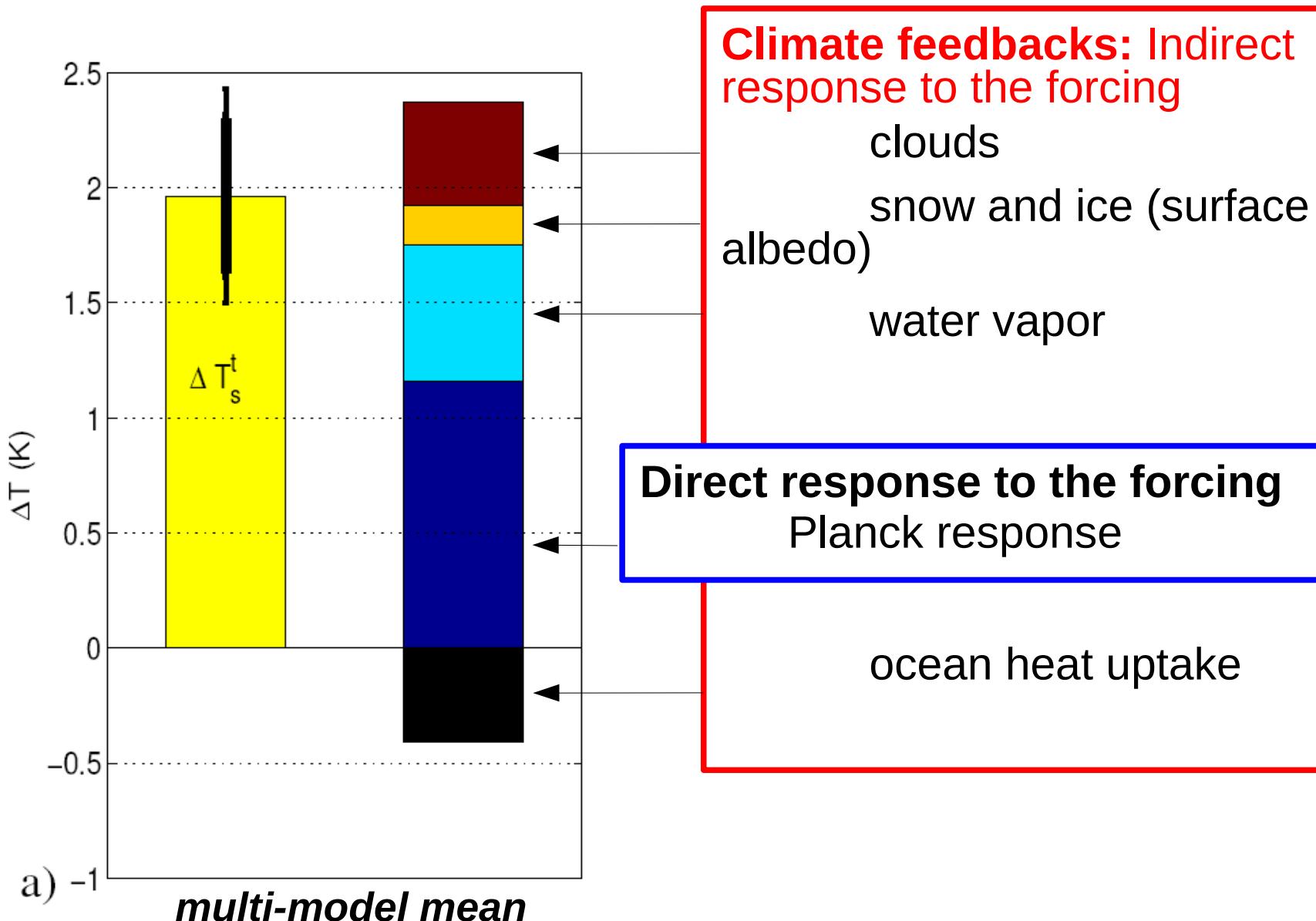
[IPCC, 2007]

Transient temperature response to a CO₂ doubling (CO₂ increase 1%/year, 70 years)



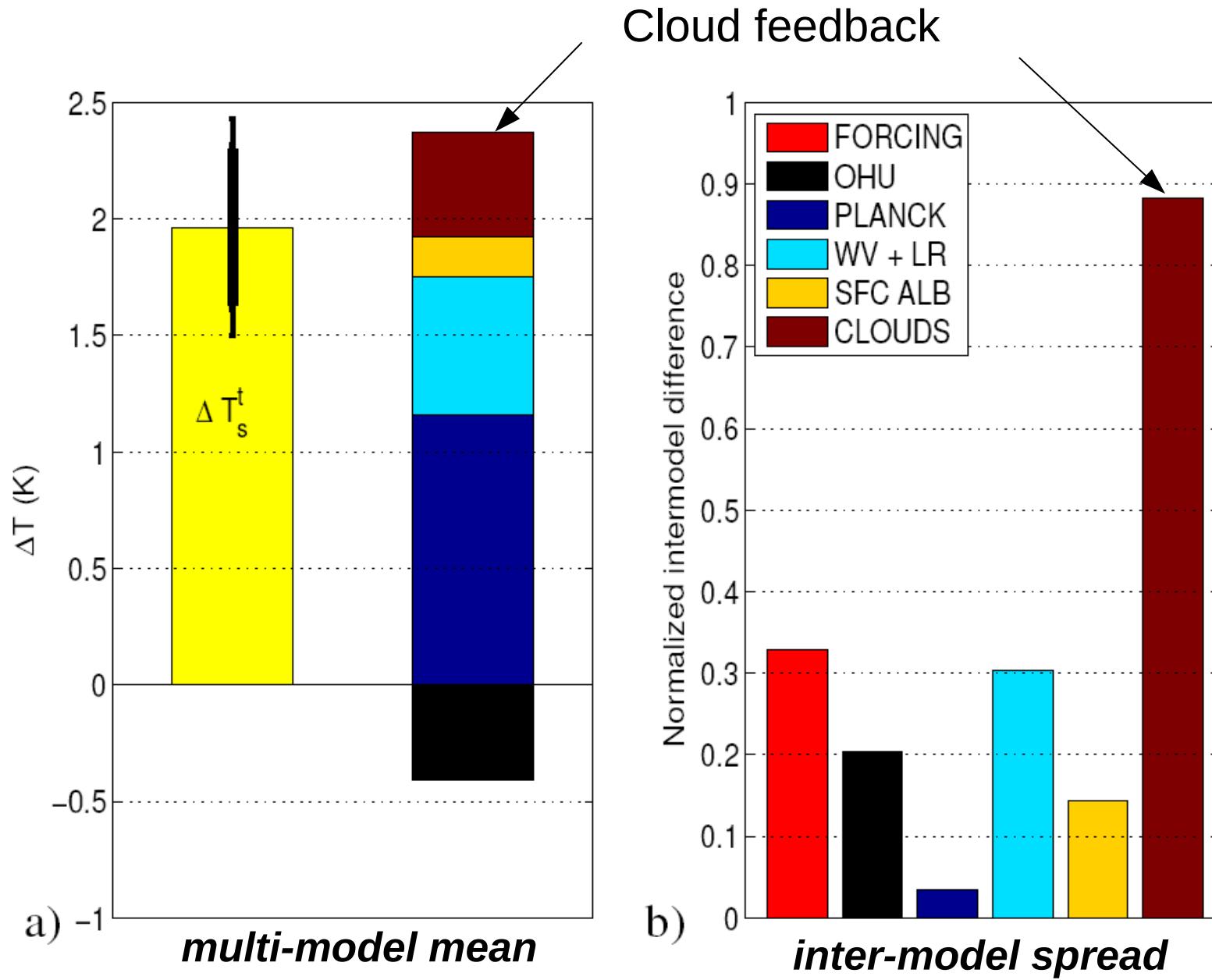
(Dufresne & Bony, 2008)

Transient temperature response to a CO₂ doubling (CO₂ increase 1%/year, 70 years)



(Dufresne & Bony, 2008)

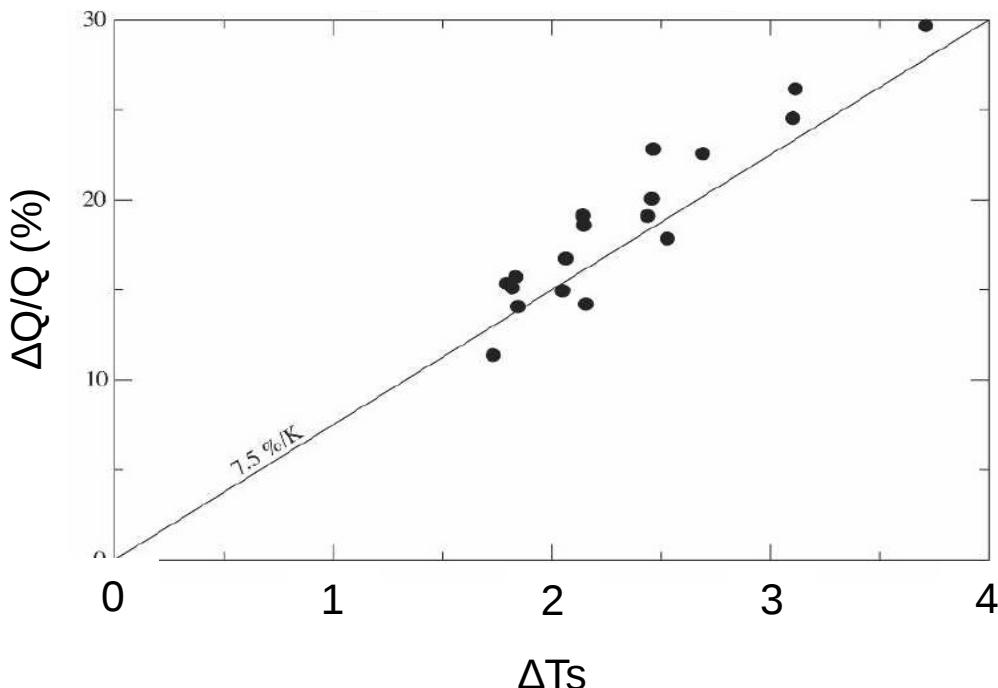
Transient temperature response to a CO₂ doubling (CO₂ increase 1%/year, 70 years)



(Dufresne & Bony, 2008)

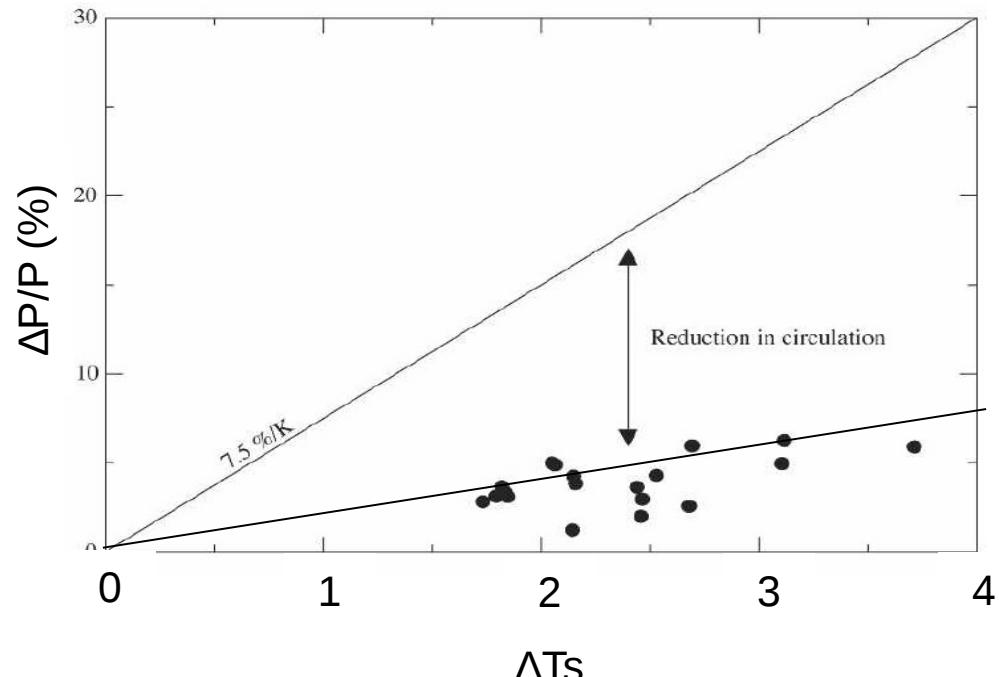
Precipitation changes

Change of the amount of **water vapor H₂O**
vs change of the average surface
temperature



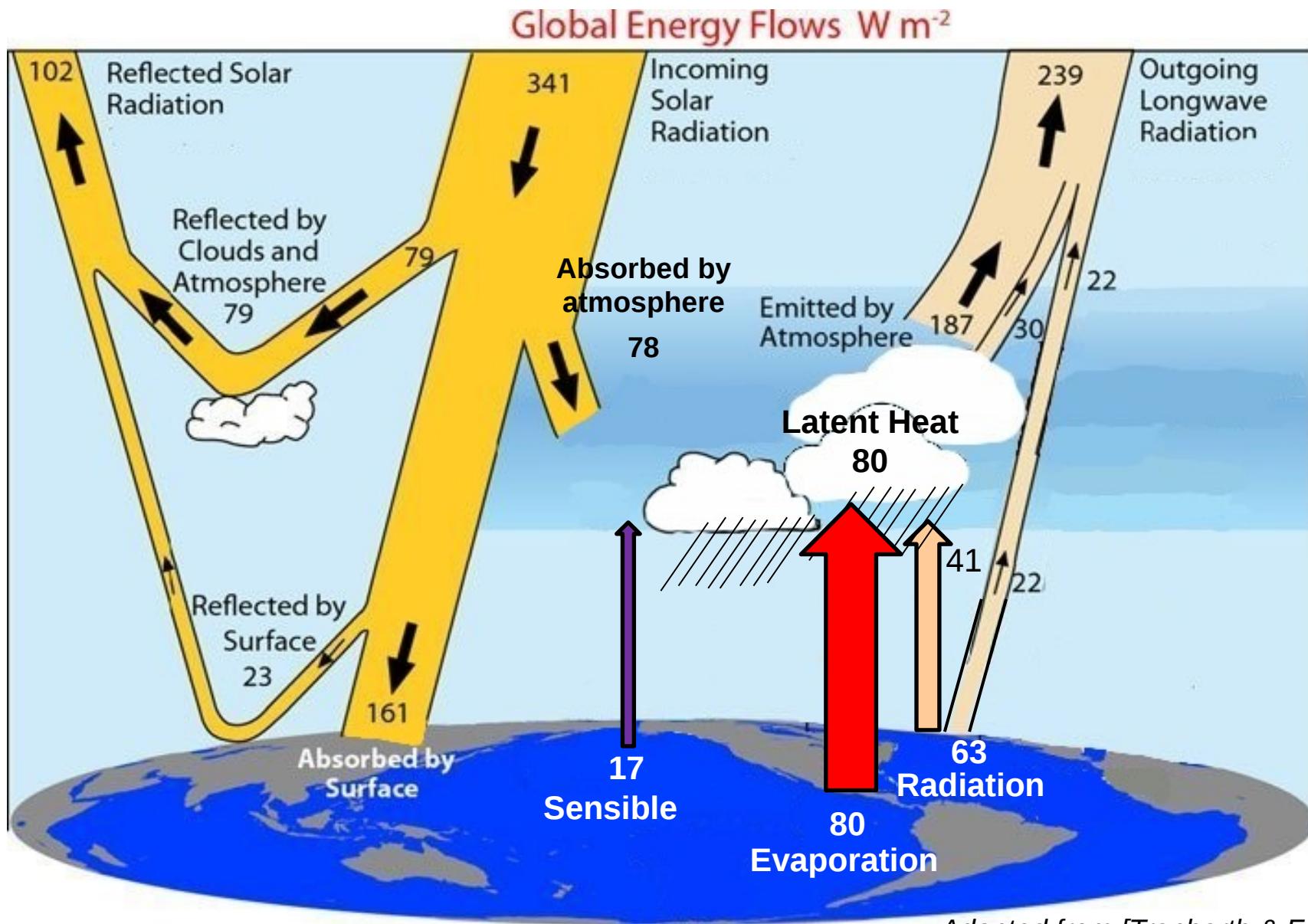
$$\Delta Q/Q (\%) \approx 7.5 \Delta T_s$$

Change of **precipitation** vs change of
the average surface **temperature**



$$\Delta P/P (\%) \approx 1.5 \Delta T_s$$

The change of the global average precipitation does not depend directly from the change of global average water vapor



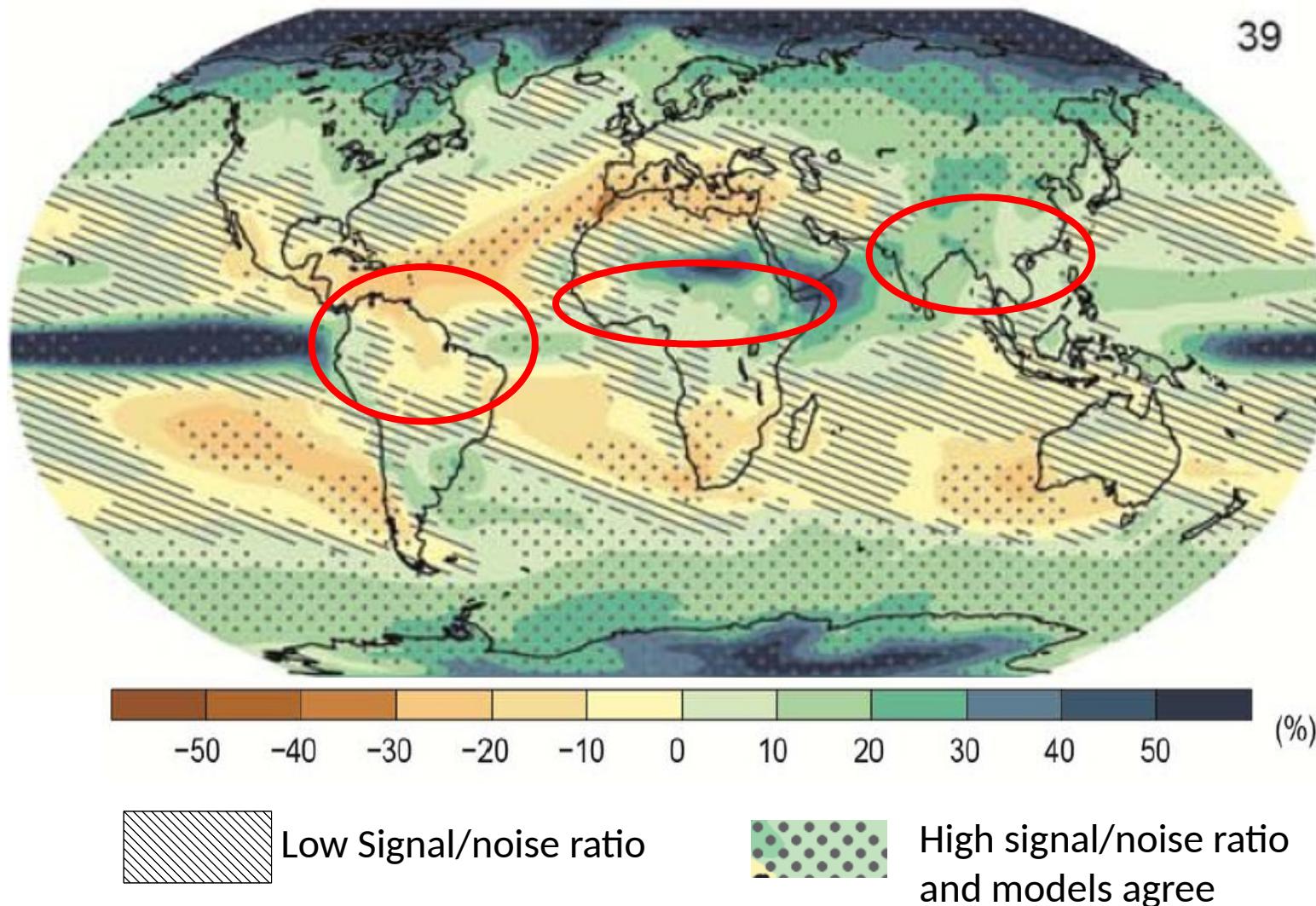
Adapted from [Trenberth & Fasullo, 2012]



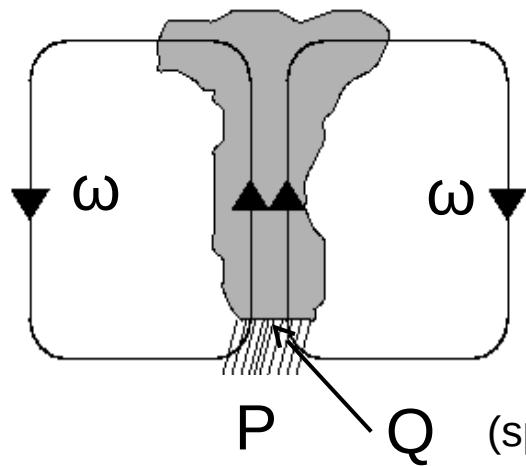
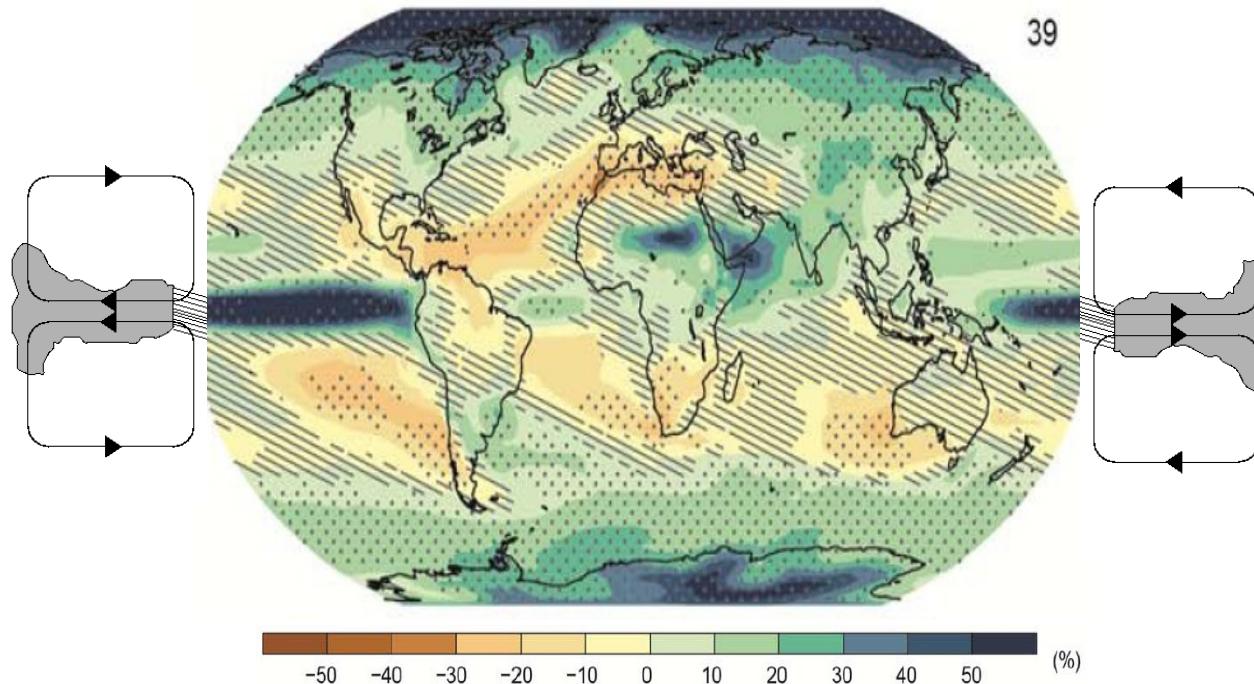
The change of the global average precipitation is constrained by the radiative cooling of the atmosphere

Precipitation changes: Geographical distribution

Relative change in average precipitation, RCP8.5 scenario (2081-2100)



Precipitation changes



Precipitations
changes

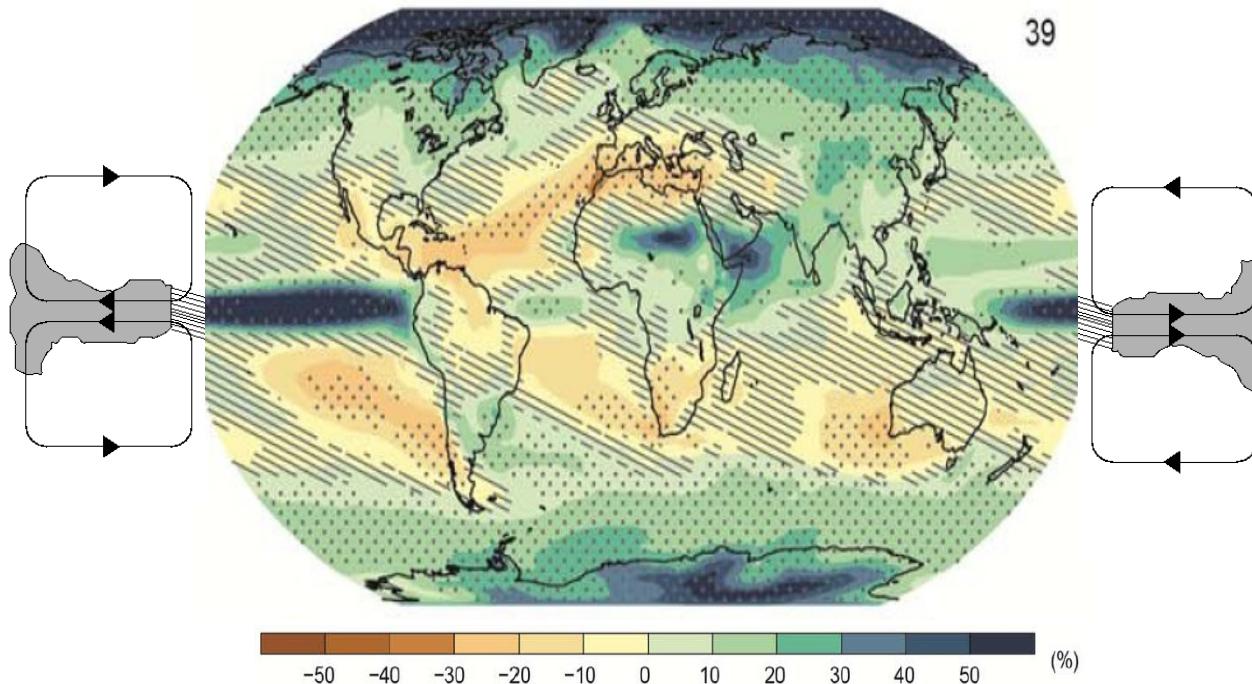
(specific humidity)

Thermodynamic
response

$$\Delta P \approx \omega \Delta Q + Q \Delta \omega$$

Dynamic
response

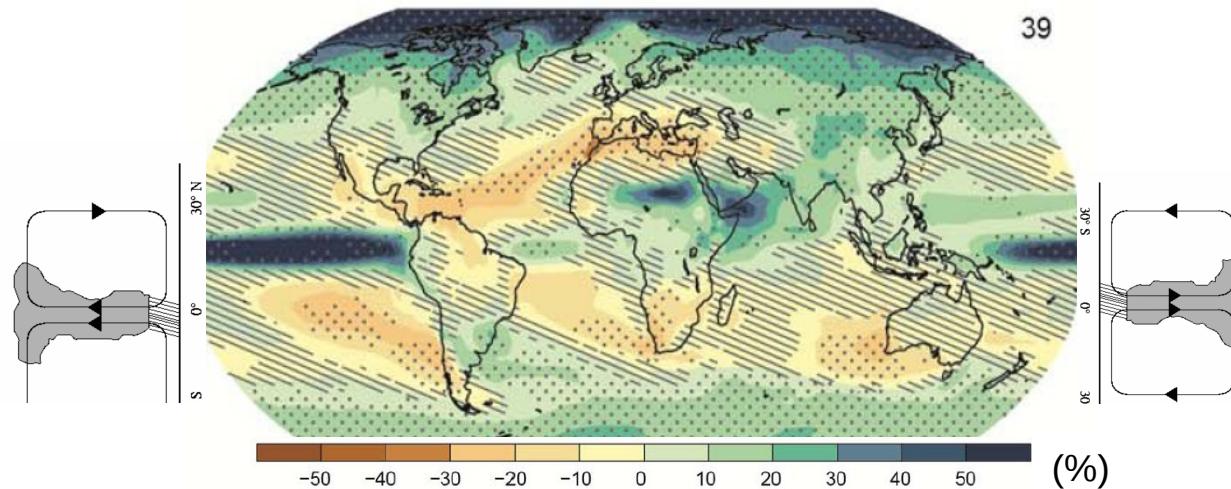
Precipitation changes



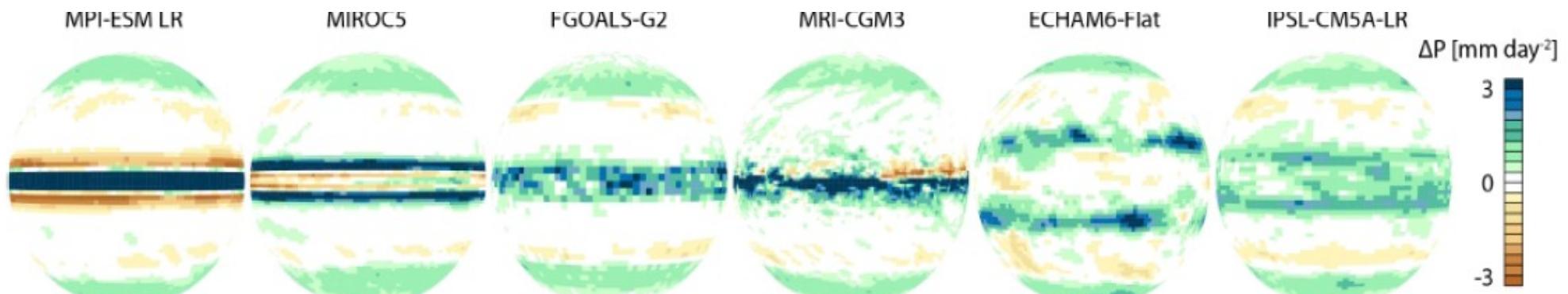
At the global scale:

- Precipitation increases in some regions while decreasing in others
- the **contrast between wet and dry regions** is expected to **increase**
- same with the contrast between wet and dry seasons

Precipitation changes



And in a simpler world? Precipitation changes in response to a uniform increase of temperature of 4K for aqua-planets



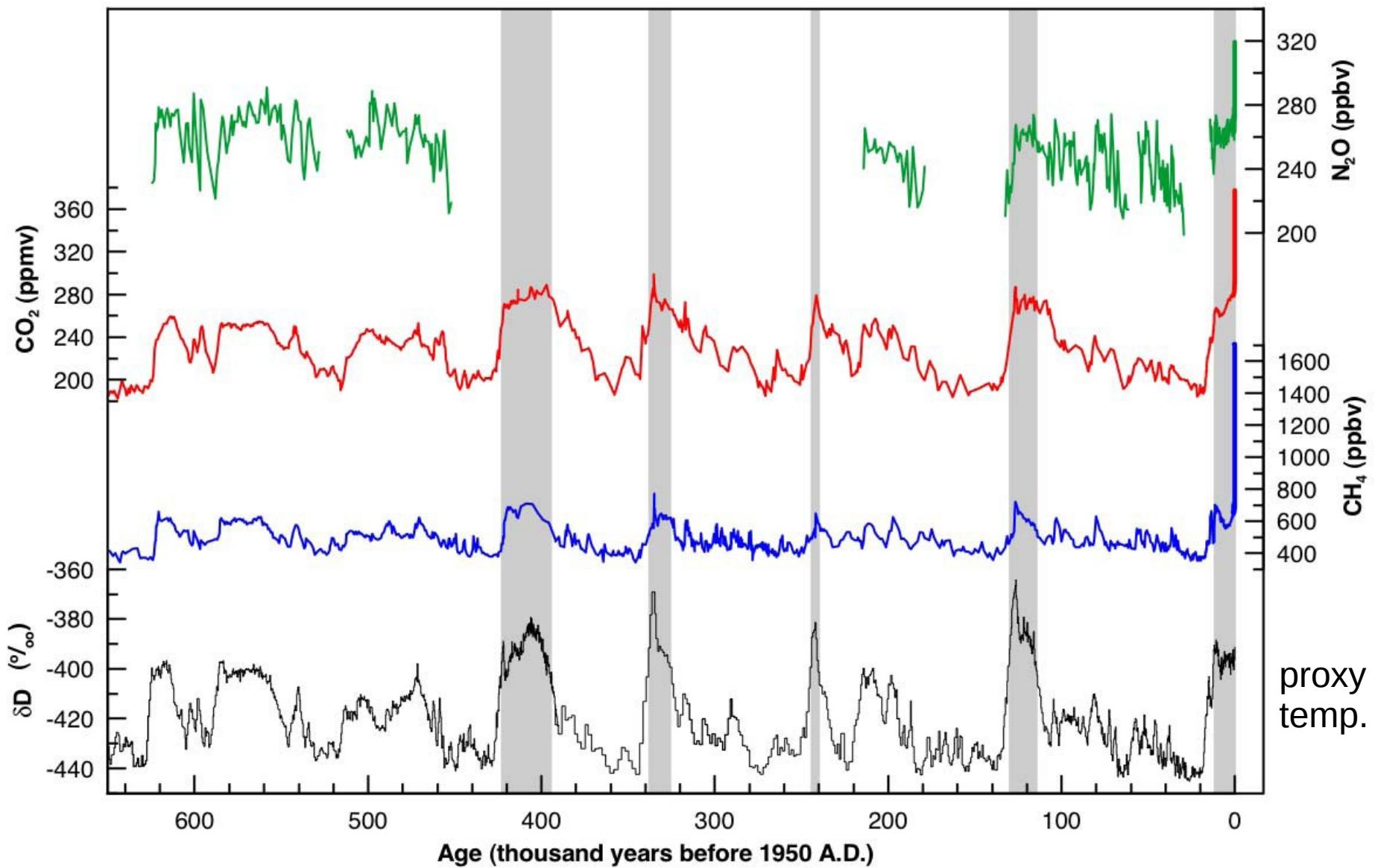
[Stevens & Bony, 2013]

A large fraction of the spread in precipitation changes originates from fundamental problems in water-vapor-temperature-circulation interactions

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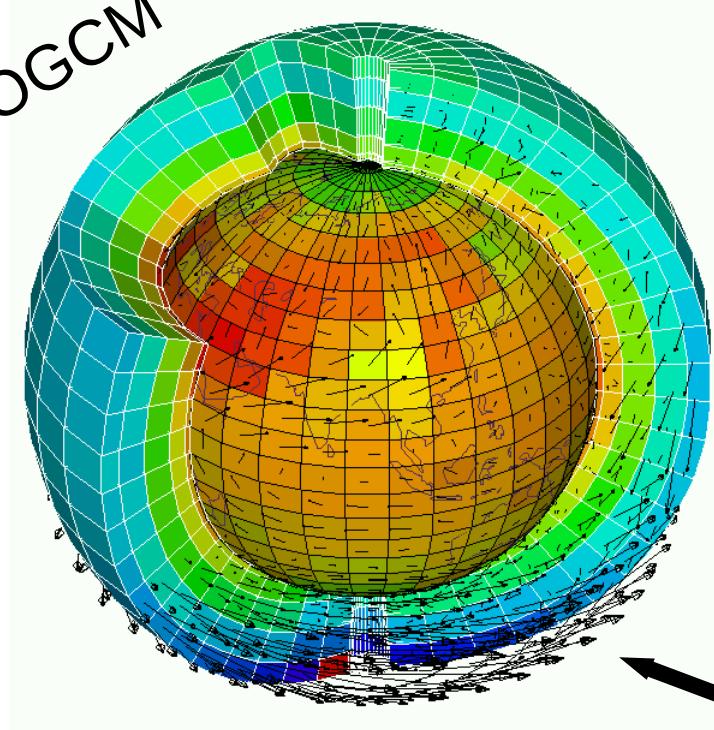
Paleoclimate changes



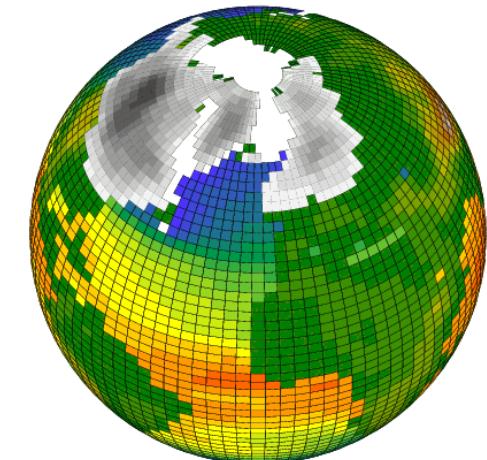
[IPCC, 2013]

Simulation of Last Glacial Maximum (LGM)

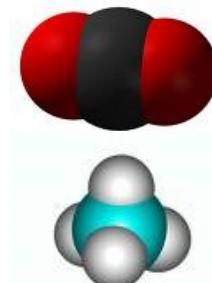
AOGCM



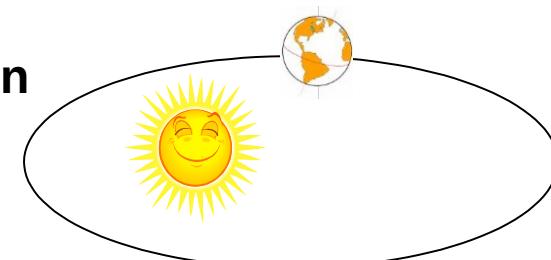
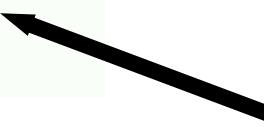
Ice sheet



Atmospheric composition
CO₂: 185 ppm
CH₄: 350 ppb...



Insolation
21ky BP



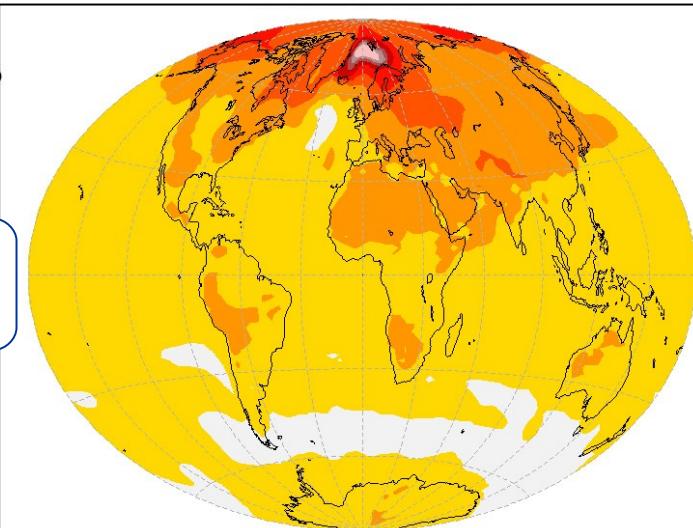
Greenhouse gas forcing ~ future climate
Other main forcings: ice sheet

Change in surface temperature

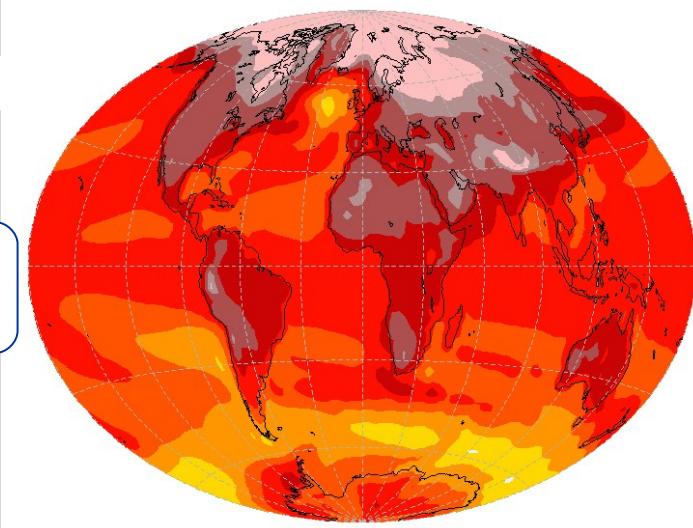
Difference between **2100** and **1990**

IPSL-CM5A-LR

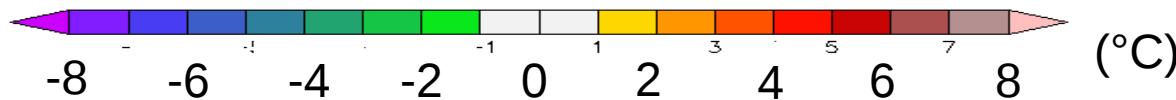
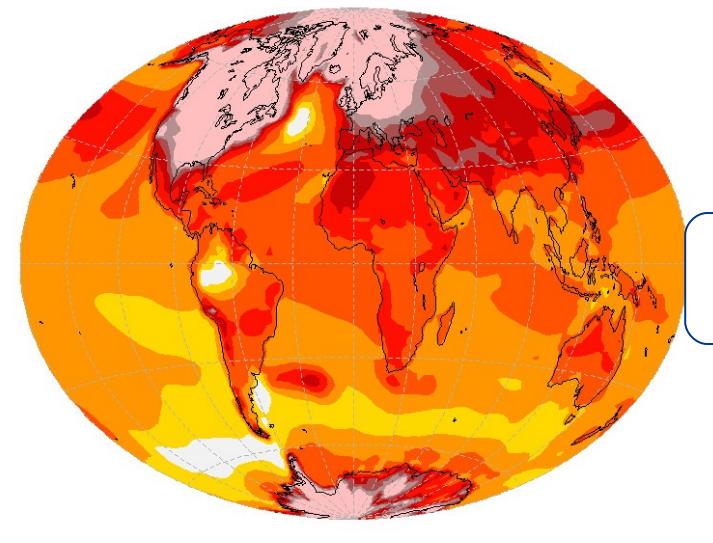
RCP2.6



RCP8.5

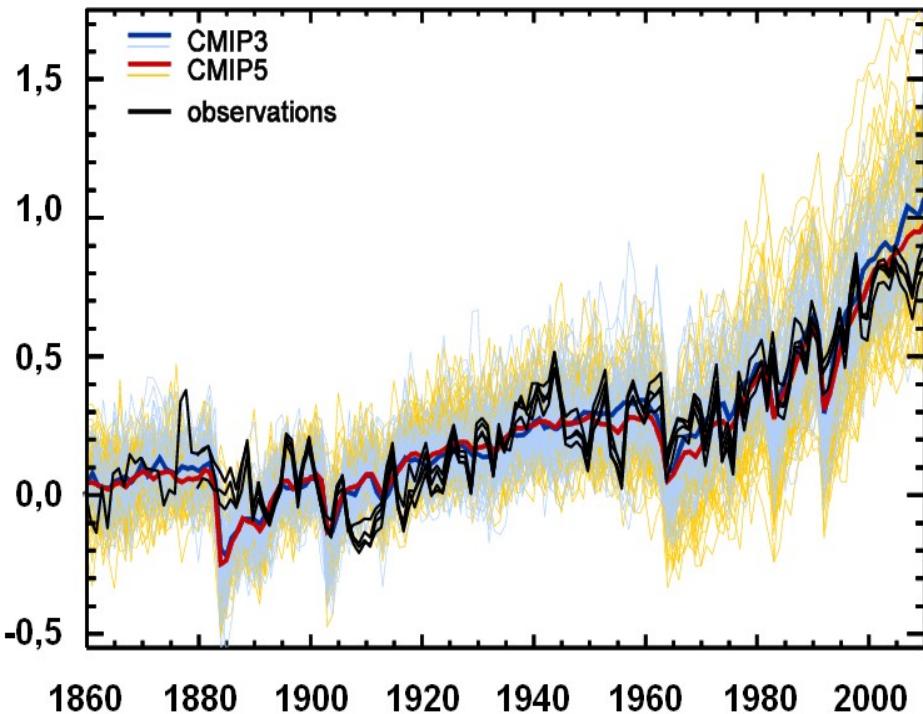


Glacial



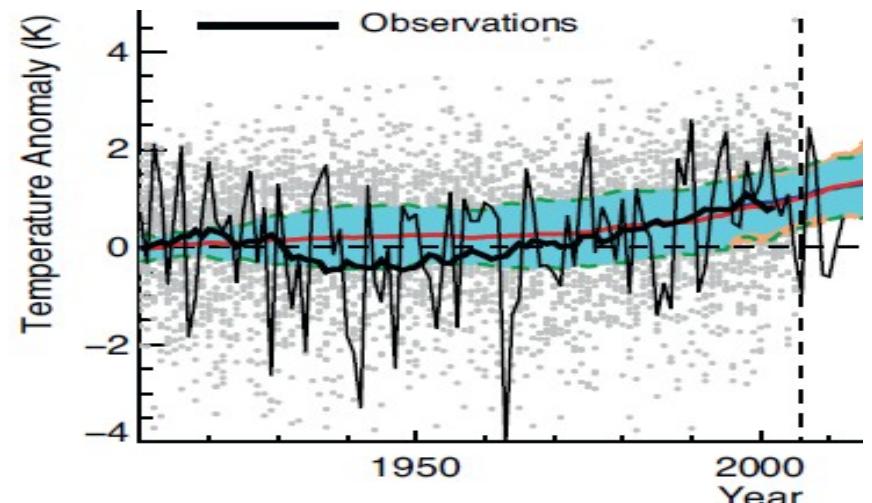
Surface temperature evolution: observation and models

Annual global mean

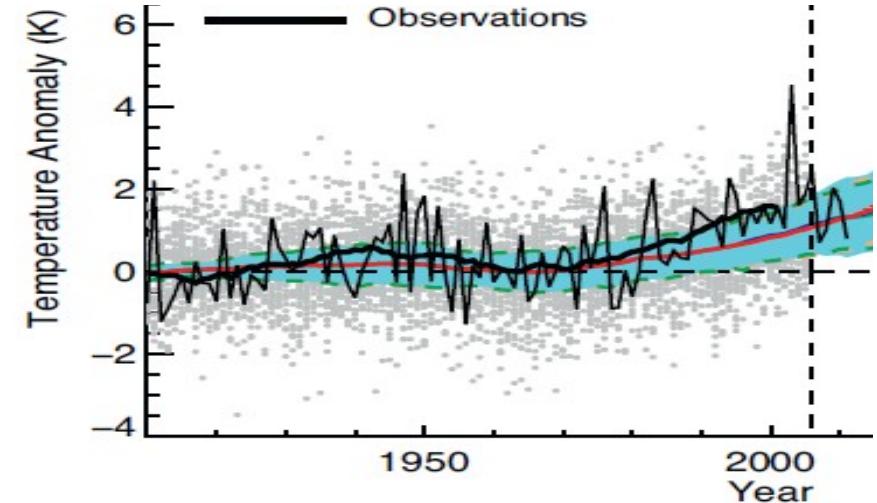


[IPCC, 2013]

Winter mean over France



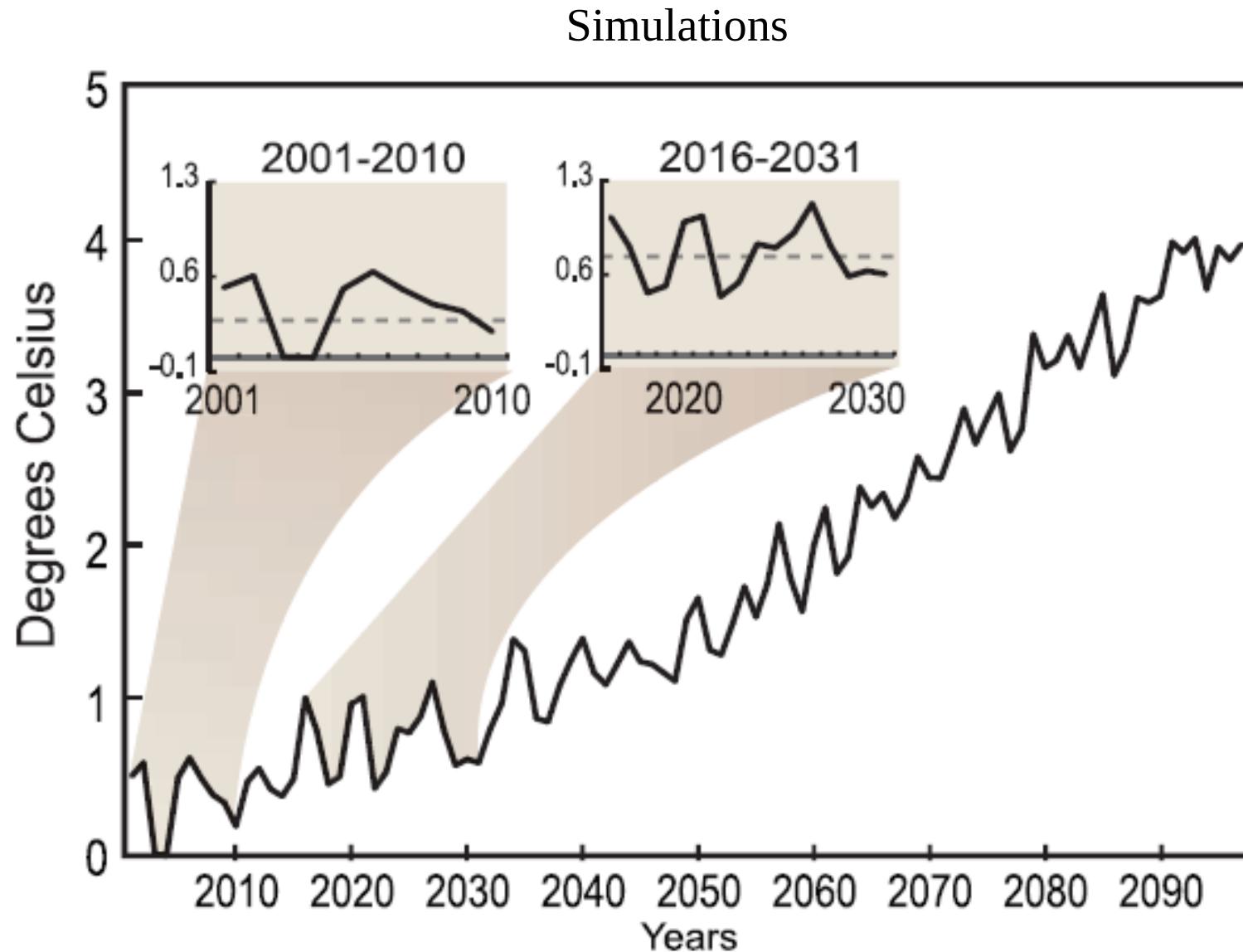
Summer mean over France



[Terray et Boé, 2013]

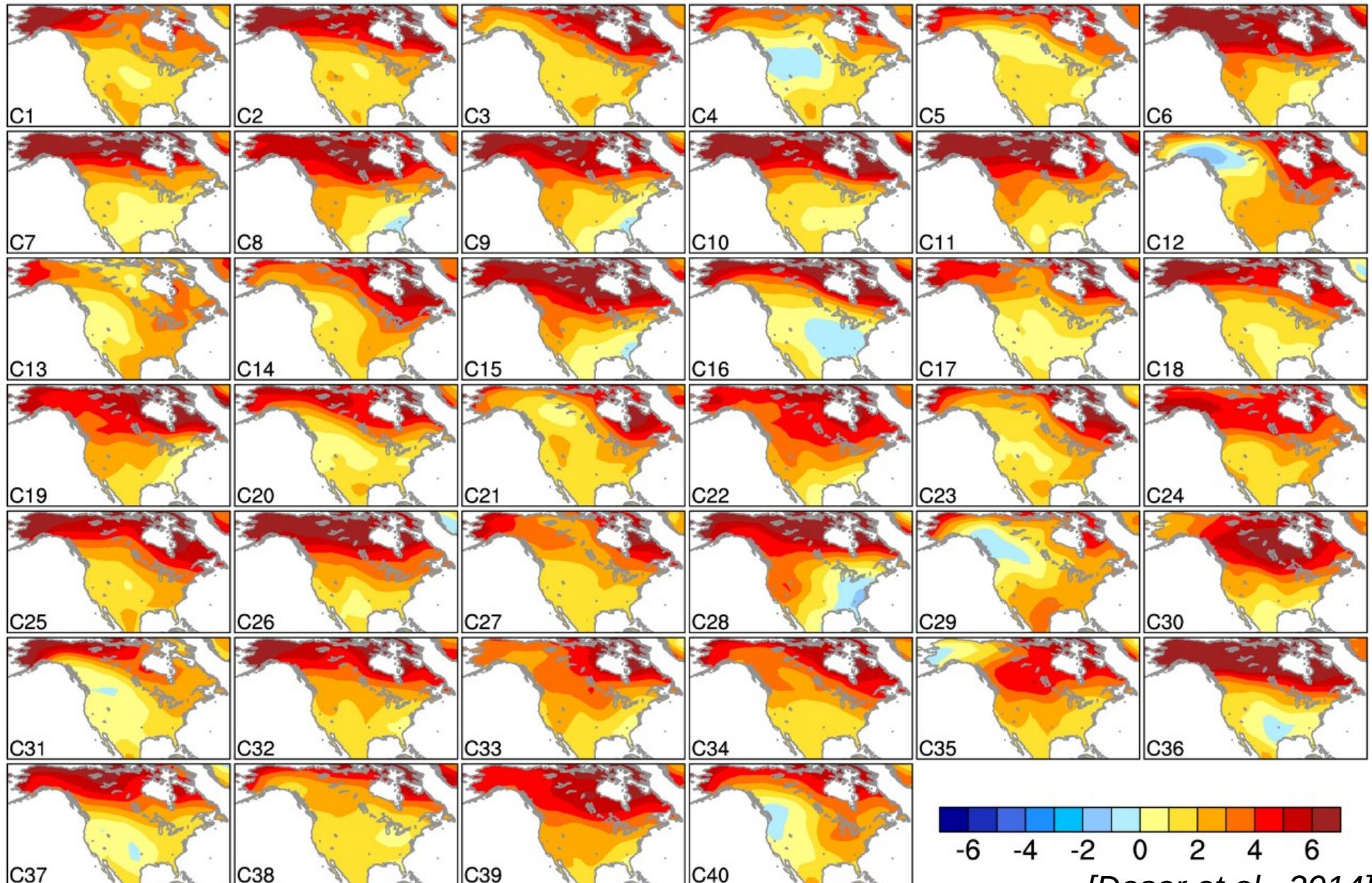
Les variations du climat sont elles régulières?

Variations et variabilité du climat



Changement climatique et variabilité naturelle

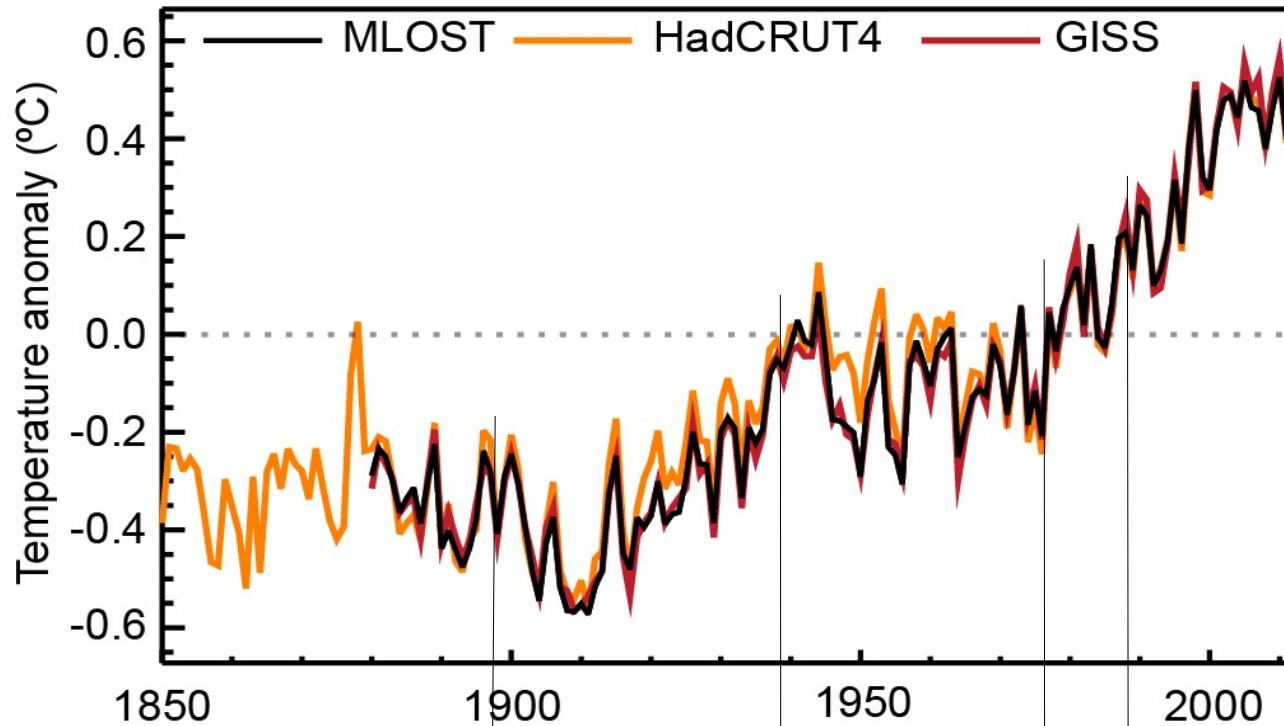
Tendance sur 50 ans de la température hivernale ($^{\circ}\text{C}/50$ ans)



Outlook

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Premières projections climatiques alors que la température a peu augmenté



[GIEC 2013]

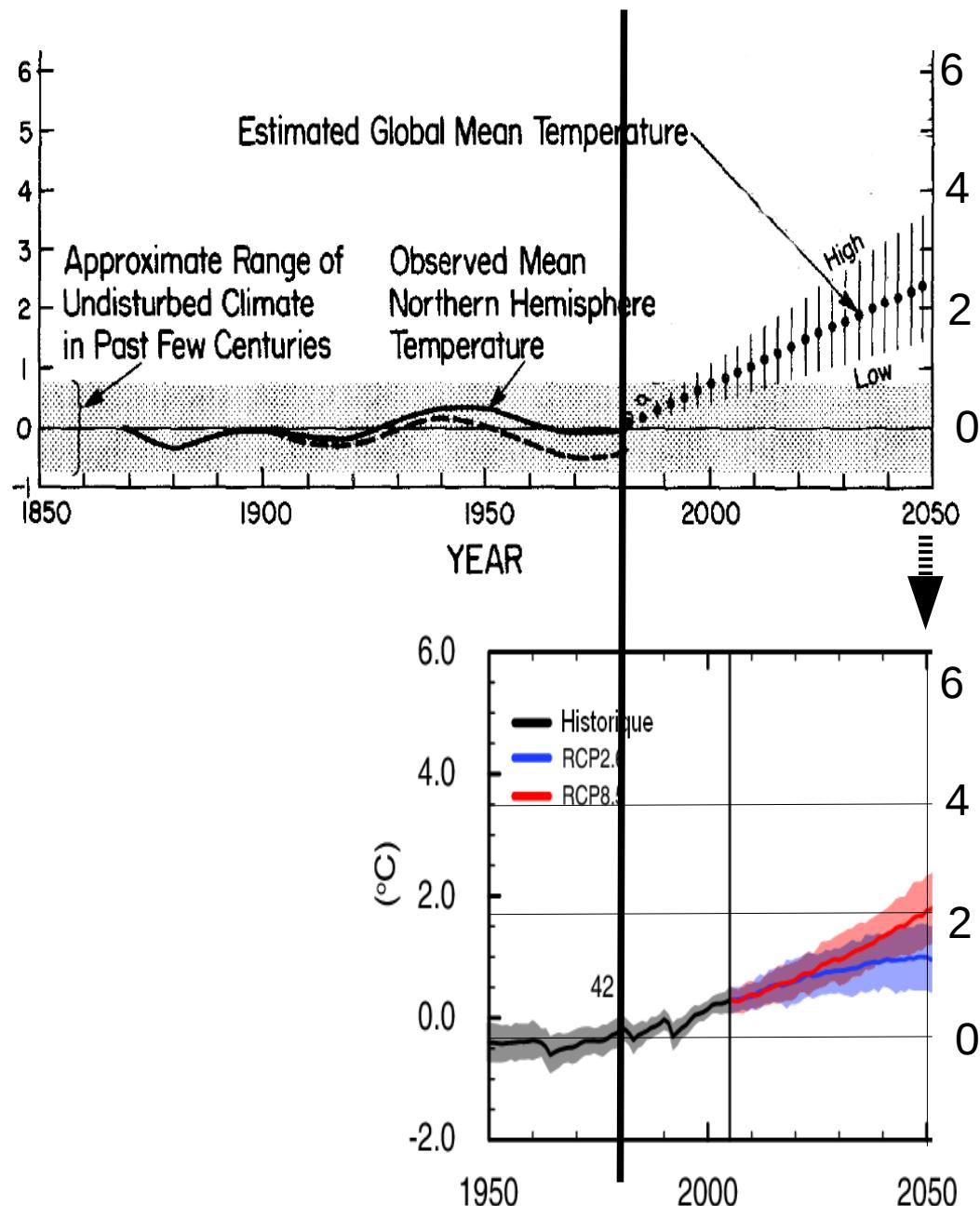
1897: S. Arrhenius:
première estimation du
rôle du CO₂

1937: G. Callendar:
nouvelle estimation du
rôle du CO₂

1988: Création du GIEC

1970-1980: Premières
projections climatiques avec des
modèles numériques

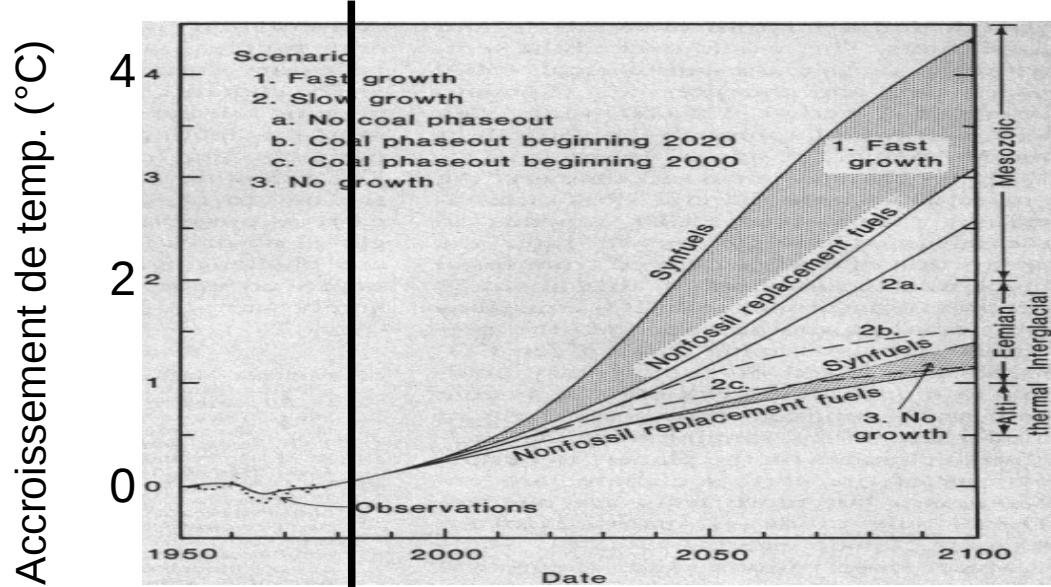
Premières projections climatiques alors que la température a peu augmenté



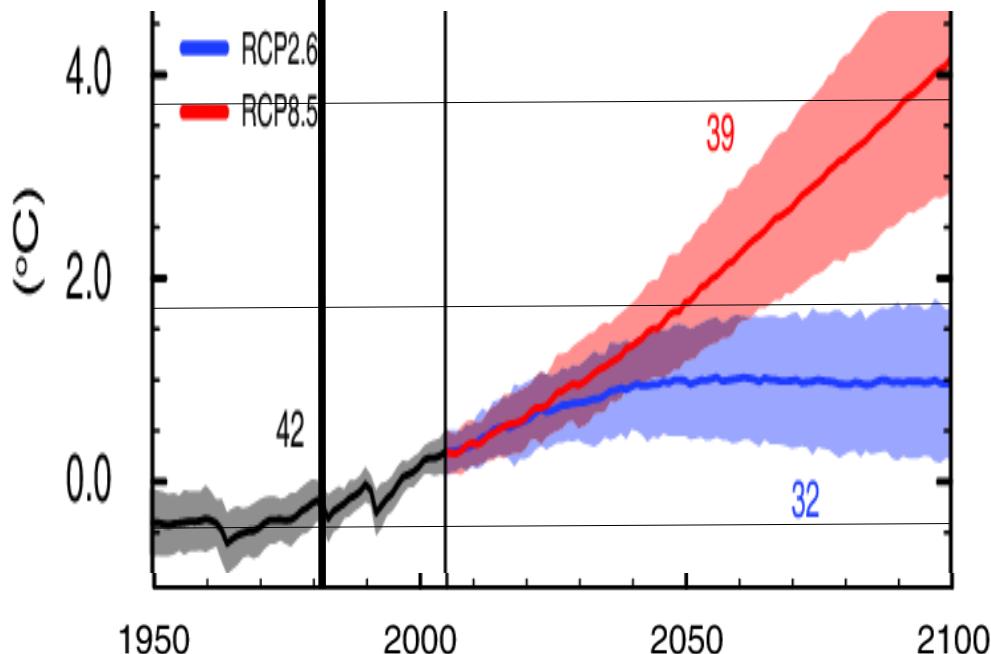
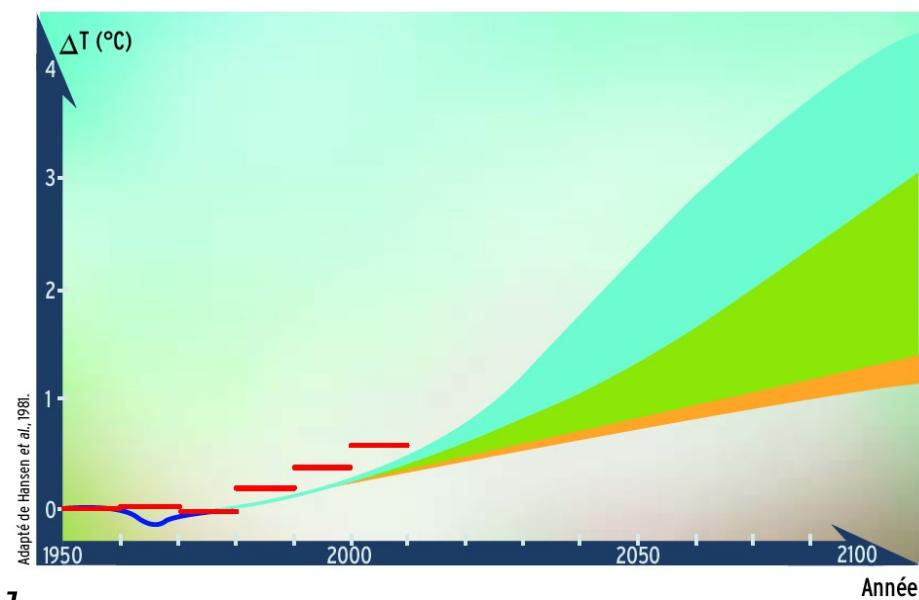
[Kellogg 1977]

[GIEC 2013]

Premières projections climatiques alors que la température a peu augmenté



[Hansen et al. 1981]



[GIEC 2013]

Conclusions

La construction et la confiance dans les modèles climatiques repose sur un ensemble méthodologique :

- Le rôle important jouée par **quelques lois physiques fondamentales**
- Les allers-retours permanents entre le **découpage** du système complet pour se focaliser sur une partie et l'**étude du système complet** (développement, évaluation, compréhension)
- L'analyse des résultats des modèles pour **comprendre** leurs fonctionnements : résultats **robustes versus dispersion**
- La comparaison aux changements observés

Changements climatiques futurs :

- Les questions évoluent et passent de l'**alerte** à la quantification et l'**anticipation des risques** associés
- Dans cette évolution, il y a un **saut** d'ordre de grandeurs sur les **exigences** vis-à-vis des modèles climatiques.
- Importance de la **représentation des processus** et de la **compréhension** des phénomènes climatiques



Merci de votre attention