

The physical parameterizations of LMDZ

LMDZ team

Laboratoire de Météorologie Dynamique / IPSL / CNRS /SU

LMDZ training, January 2024

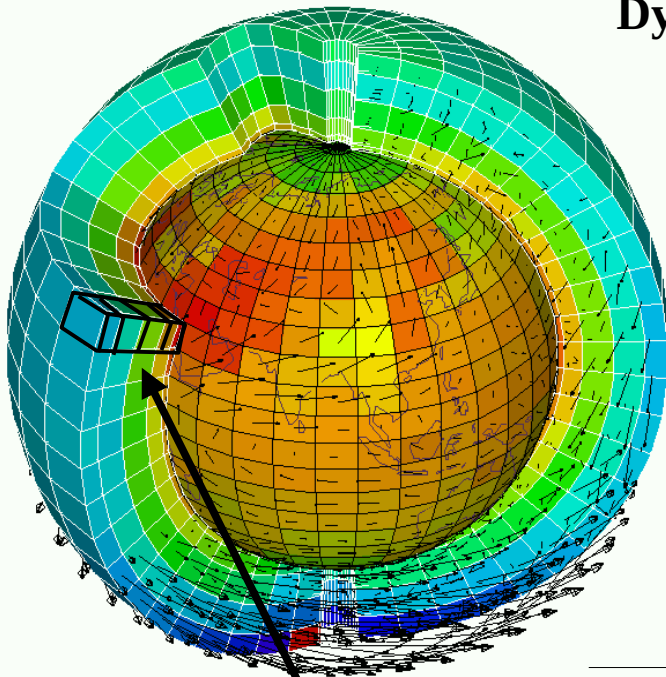
Part I : today (Frédéric presenting)

- Principles
- Radiation (clear sky)
- Reynolds decomposition
- Turbulent diffusion
- Mass flux representation of the convective boundary layer
- Subgrid scale orography
- Practice

Part II : tomorrow (Jean-Baptiste presenting)

- Convection
- Clouds

Dynamical core : primitive equations discretized on the sphere



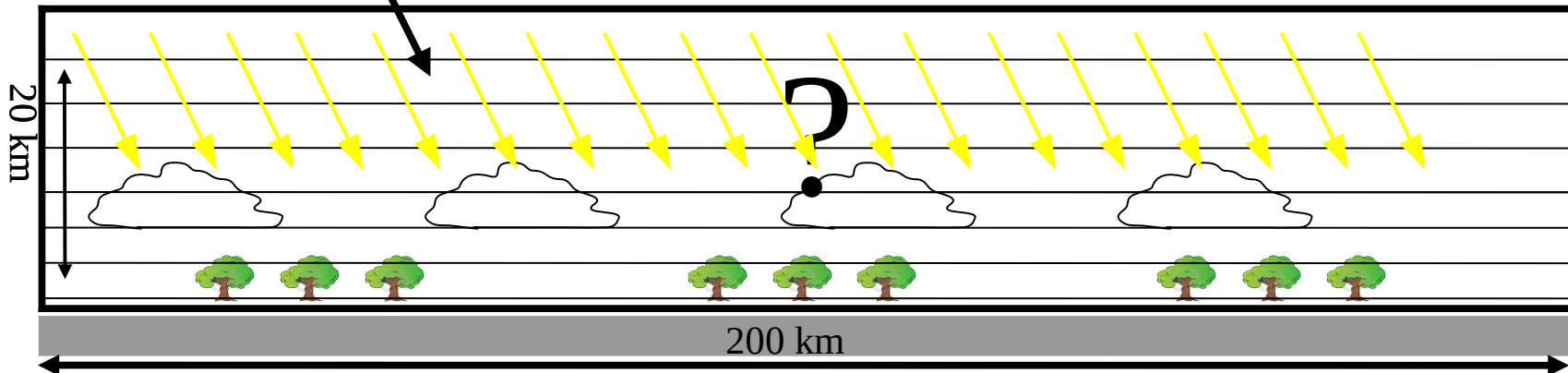
- Mass conservation

$$D\rho/Dt + \rho \operatorname{div}\underline{U} = 0$$
- Potential temperature conservation

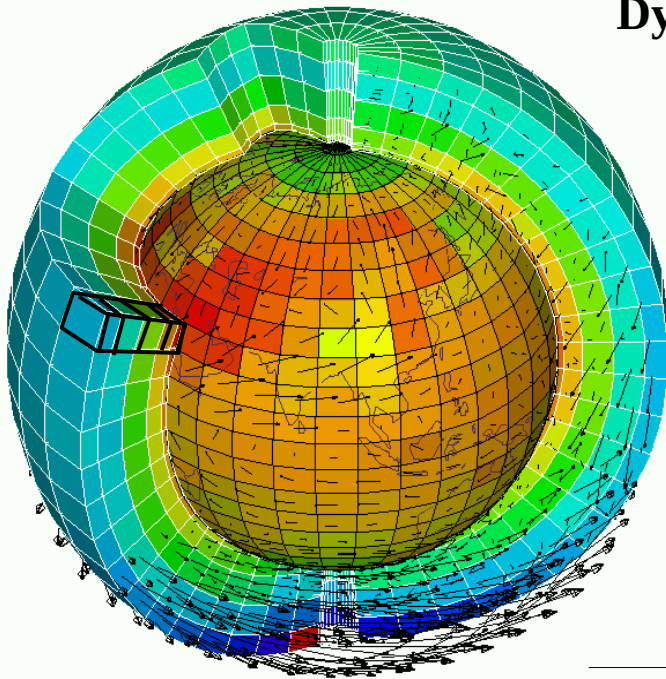
$$D\theta / Dt = Q / C_p (p_0/p)^\kappa$$
- Momentum conservation

$$D\underline{U}/Dt + (1/\rho) \operatorname{grad}p - g + 2 \underline{\Omega} \wedge \underline{U} = \underline{F}$$
- Secondary components conservation

$$Dq/Dt = Sq$$



Dynamical core : primitive equations discretized on the sphere



- Mass conservation
$$D\rho/Dt + \rho \operatorname{div}\underline{U} = 0$$
- Potential temperature conservation
$$D\theta / Dt = Q / C_p (p_0/p)^\kappa$$
- Momentum conservation
$$D\underline{U}/Dt + (1/\rho) \operatorname{grad}p - \mathbf{g} + 2 \underline{\Omega} \wedge \underline{U} = \underline{F}$$
- Secondary components conservation
$$Dq/Dt = Sq$$

Parameterizations purpose : account for the effect of processes non resolved by the dynamical core

→ **Traditional « source » terms in the equations**

- Q : Heating by radiative exchanges, thermal conduction (neglected), condensation, sublimation, **subgrid-scale motions (turbulence, clouds, convection)**
- F : Molecular viscosity (neglected), **subgrid-scale motions (turbulence, clouds, convection)**
- Sq : condensation/sublimation (q = water vapor or condensed), chemical reactions, photo-dissociation (ozone, chemical species), micro physics and scavenging (pollution aerosols, dust, ...), **subgrid-scale motions (turbulence, clouds, convection)**

Model tendencies

The integration of a given prognostic variable X ($T, \vec{v}(u, v, w), p, \rho, q_{vap}$) can be written as :

$$X_{t+\Delta t} = X_t + \left(\frac{\partial X}{\partial t} \right)_{\text{dyn}} \Delta t \text{ (dynamical core)} \quad (1)$$

$$+ \left(\frac{\partial X}{\partial t} \right)_{\text{param}} \Delta t \text{ (parameterizations)} \quad (2)$$

From model outputs

$$\text{temp}(t+\text{dtphys})-\text{temp}(t)=\text{dtdyn}+\text{dtphy}$$

$$\text{ovap}(t+\text{dtphys})-\text{ovap}(t)=\text{dqdyn}+\text{dqphy}$$

$$\text{vit}[u/v](t+\text{dtphys})-\text{vit}[u/v](t)=\text{dudyn}+\text{duphy}$$

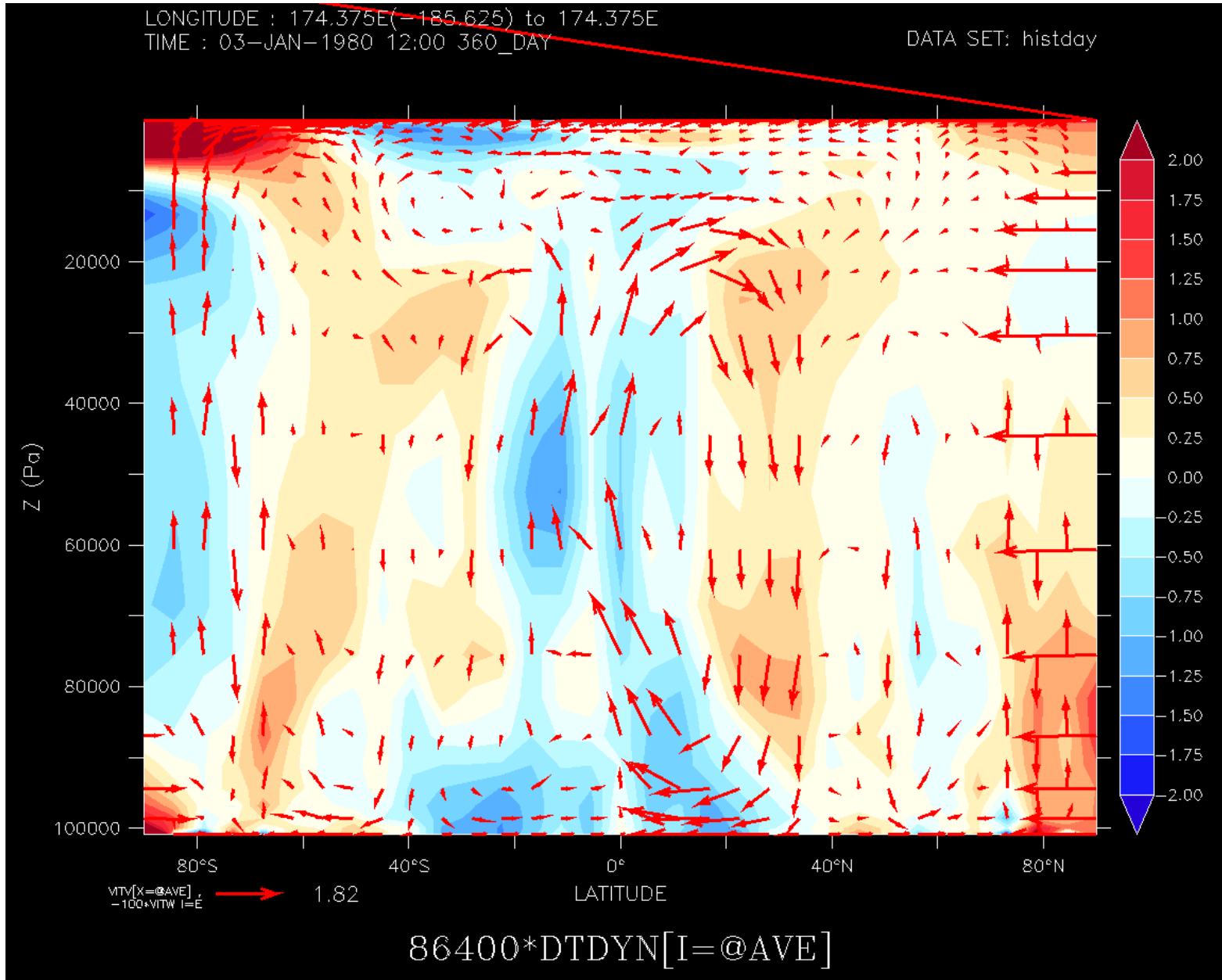
Physics time-step :

$$\text{dtphys}=\text{daysec}*\text{iphysic}/(\text{day_step}), \quad \text{day_sec}=86400$$

reg/l=3

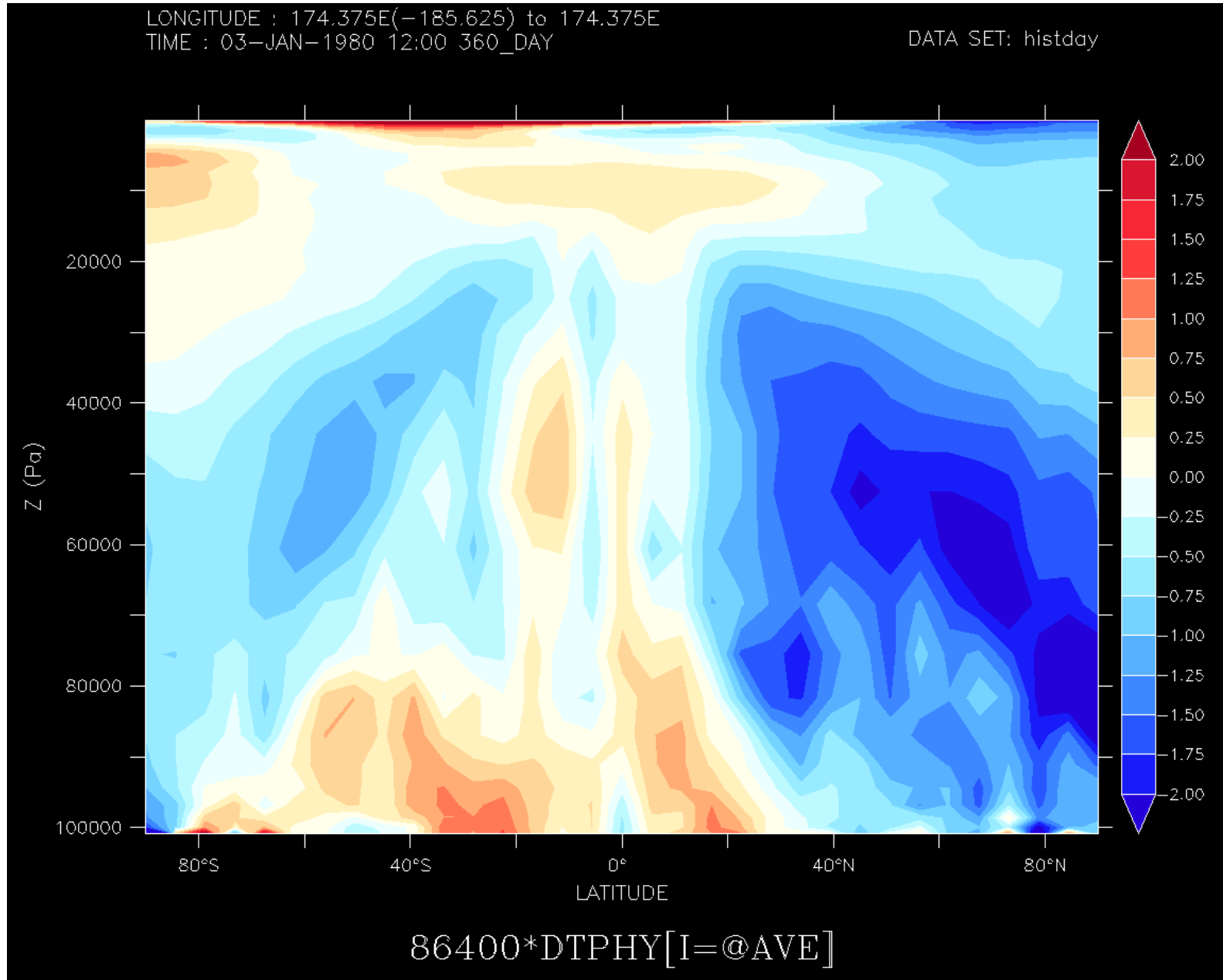
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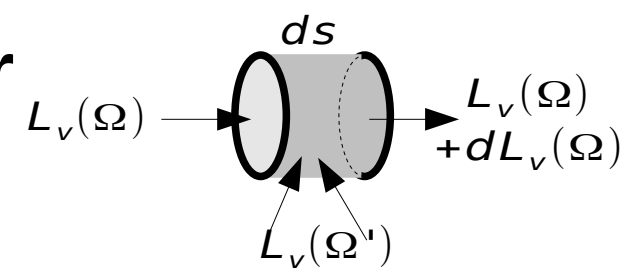


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Parameterization of radiative transfer



Radiative transfer : well known equations ...

Giving the evolution of luminance along a line of sight:

$$\frac{dL_v(\Omega)}{ds} = \underbrace{-\kappa_v L_v(\Omega)}_{\text{absorption}} + \underbrace{\kappa_v B_v(T)}_{\text{emission}} - \underbrace{\sigma_v L_v(\Omega)}_{\text{Scattered in other directions}} + \underbrace{\sigma_v \frac{1}{4\pi} \int_{4\pi} P(\Omega', \Omega) L_v(\Omega') d\Omega'}_{\text{Scattered from other directions}}$$

Computation of energy fluxes very costly

- should be integrated over all frequencies ν
- should be integrated on angles
- knowing radiative properties of scatterers and absorbers is a question by itself

Computing radiation for one full scene with reference methods for the spectral integration (line-by-line) and angular (discrete ordinates, Monte-Carlo) integrations, even for a plan parallel atmosphere without clouds, may take hours of CPU hours on super computers.

In LMDZ : using codes developed and used at ECMWF

3 codes avec des mots clé : **oldrad / rrtm / ecrad**

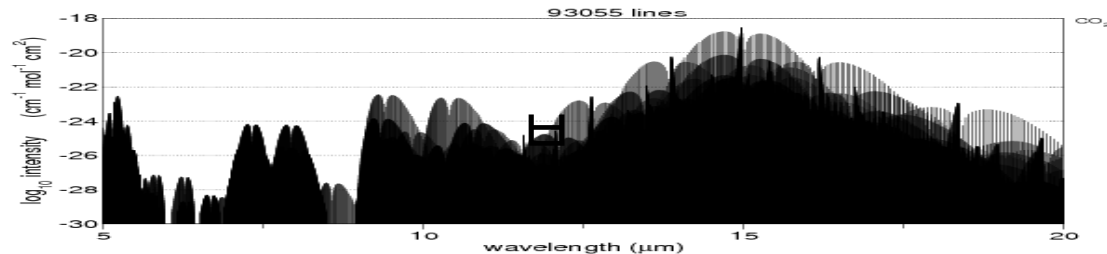
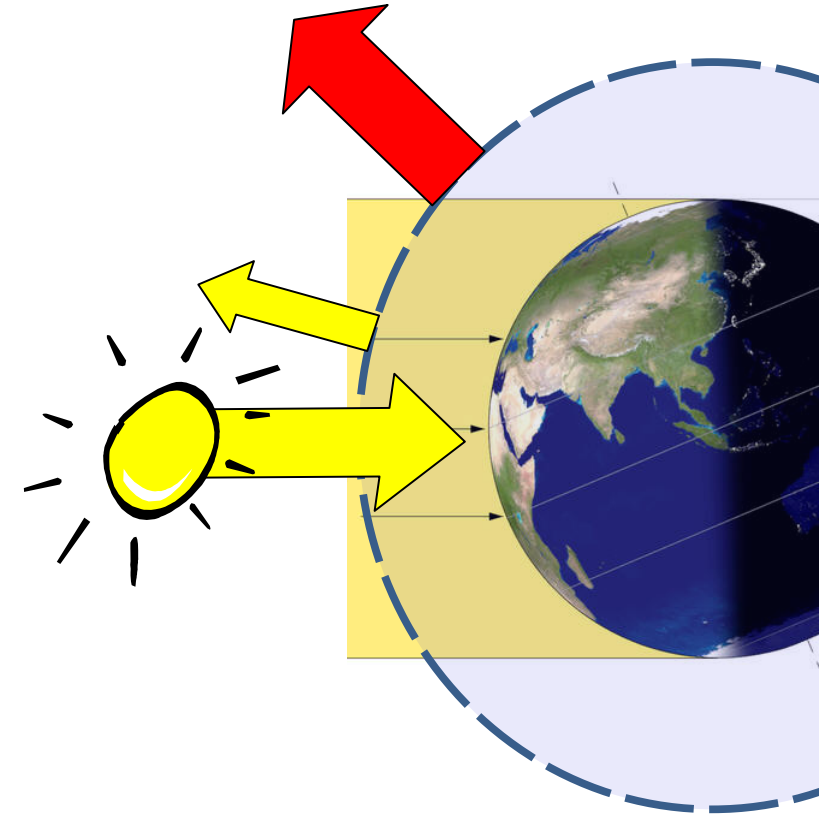
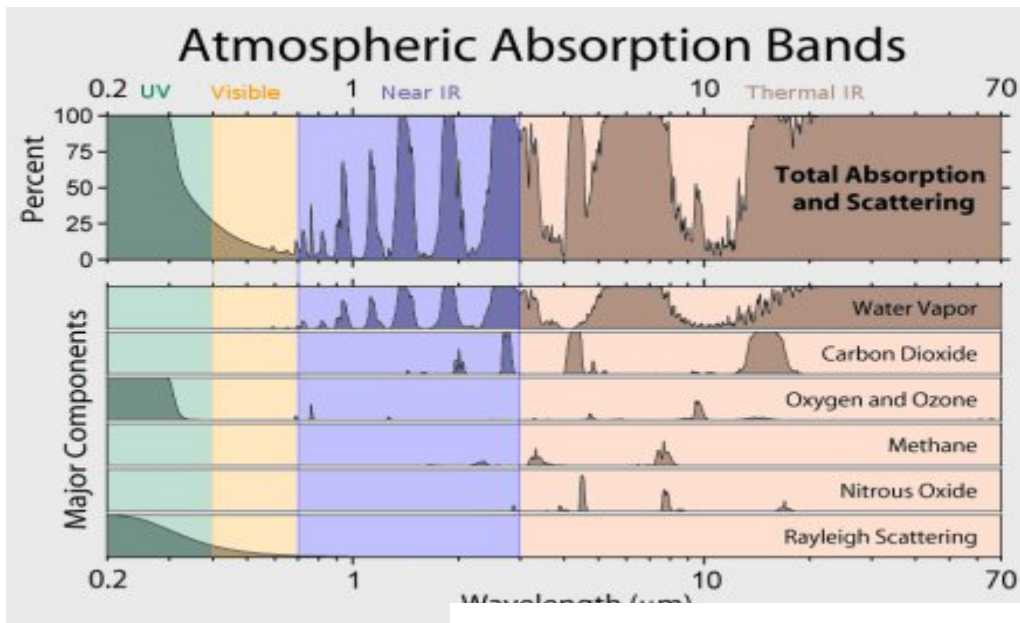
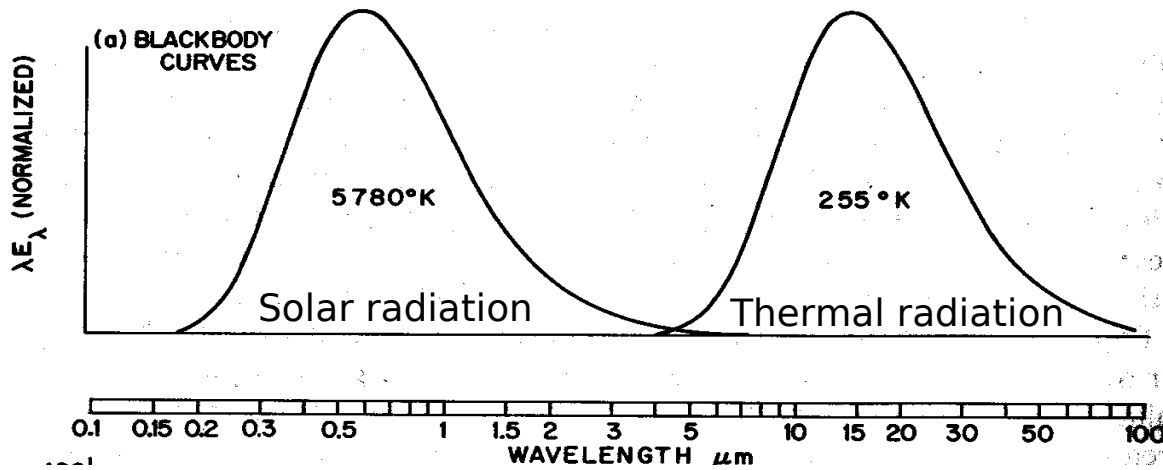
ECrad can be seen as a tool box with various options for :

Spectral integration : **S/RRTM or ECCKD**

Solver : **McICA, Tripleclouds or Spartacus**

Spectral : separating radiation between solar and thermal infrared

Valid thanks to the linearity of the radiative transfer equation with respect to sources



Approaches for spectral integration

Emissivity / band models (code « Fouquart Morcrette, 1980)

$$\epsilon_{\Delta\nu}(z_1, z_2) = \frac{1}{\Delta\nu} \int_{\Delta\nu} \epsilon_\nu(z_1, z_2) d\nu$$

Loosing a fundamental property : $\epsilon_{\Delta\nu}(z_1, z_2) = \epsilon_{\Delta\nu}(z_1, z) \epsilon_{\Delta\nu}(z, z_2)$

Cost in N² instead of N, where N is the number of layers

K-distribution methods

Replacing the integration on ν by an integration on k .

$k(P, T)$ differ depending on the transition considered.

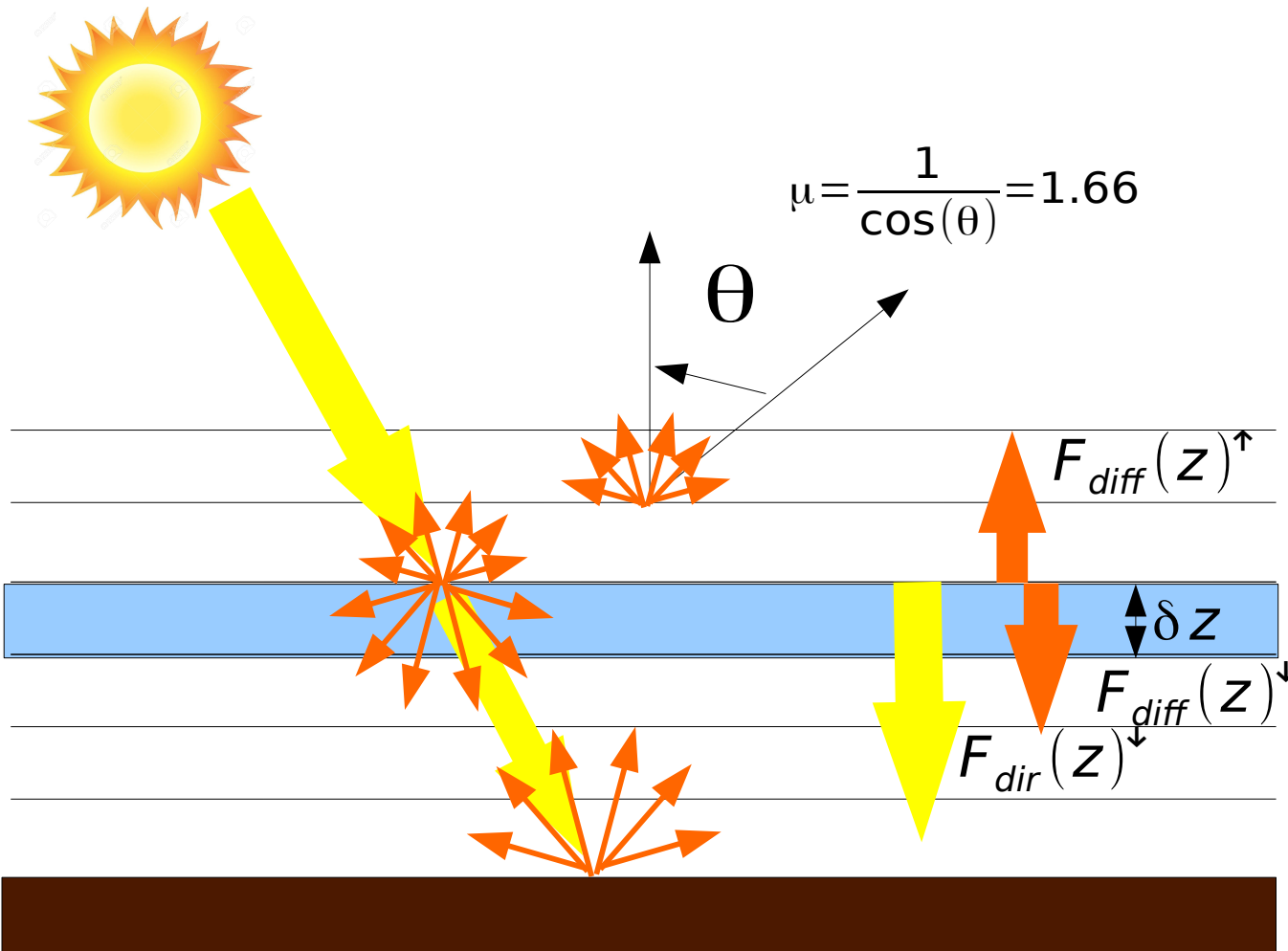
Name in .def files	SW	LW
oldrad	2 bands	6 bands
rrtm	6 bands	K-distributions (RRTM)
ecrad	K-distributions (SRTM) ECCKD	K-distributions (RRTM) ECCKD

Solar radiation : Direct radiation + 2-stream for diffuse radiation

Plane parallel approximation: homogeneous semi-infinite space

Upward and downward photons are grouped into two streams

Delta-Eddington approximation for scattering by strongly asymmetric phase functions



$$F^{\uparrow} = F_{diff}^{\uparrow} - F_{diff}^{\downarrow} - F_{dir}^{\downarrow}$$

$$F(z)^{\uparrow}$$

$$F(z - \delta z)^{\uparrow}$$

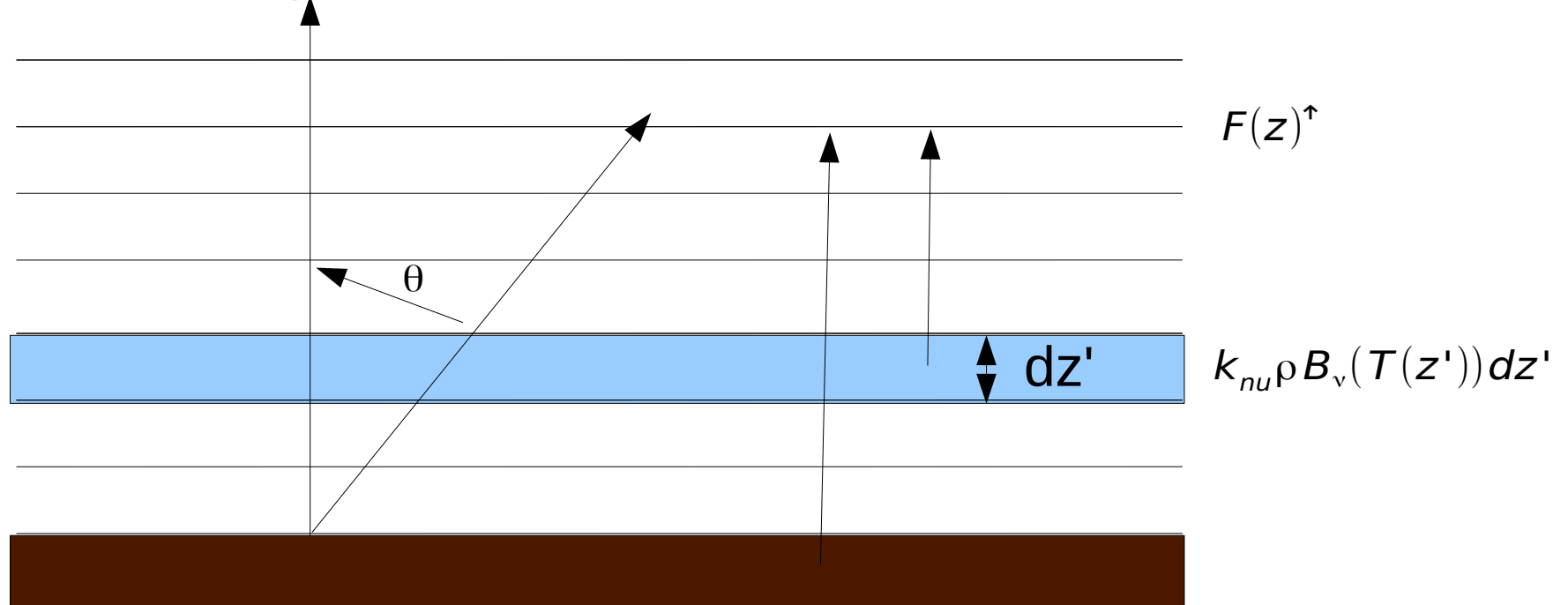
$$Q = \frac{F(z) - F(z - \delta z)}{C_p \rho \delta z}$$

Infrared, non-scattering case

Plane parallel approximation: homogeneous semi-infinite space

Diffuse" approximation

Up/down flux separation, 2-stream



$$\frac{\partial F(z)^\uparrow}{\partial z} = -k_{\nu} \rho \mu F(z)^\uparrow + k_{\nu} \rho \mu B_{\nu}(T)$$

$$\mu = \frac{1}{\cos(\theta)} = 1.66$$

$$\epsilon_{\nu}(z_1, z_2) = \exp\left[-\mu \int_{z_1}^{z_2} k_{\nu}(P, T) \rho dz\right]$$

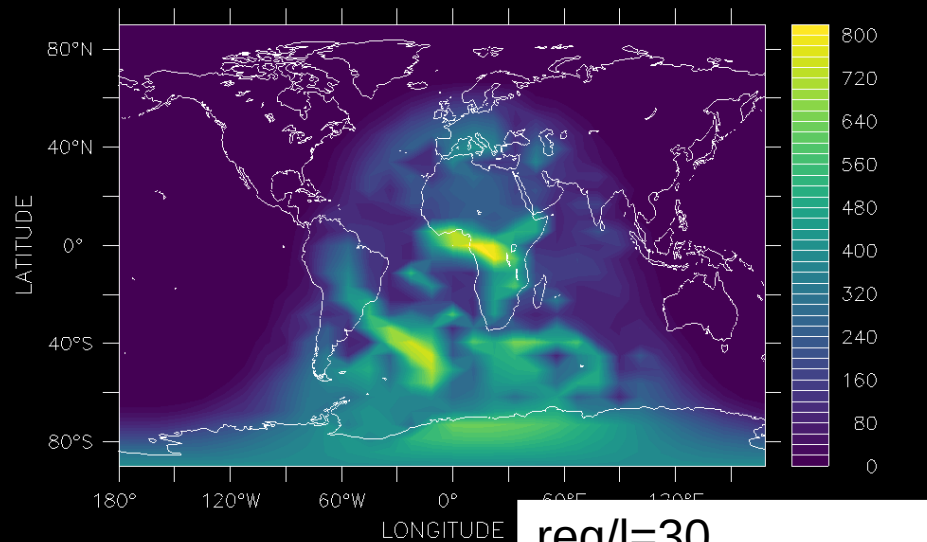
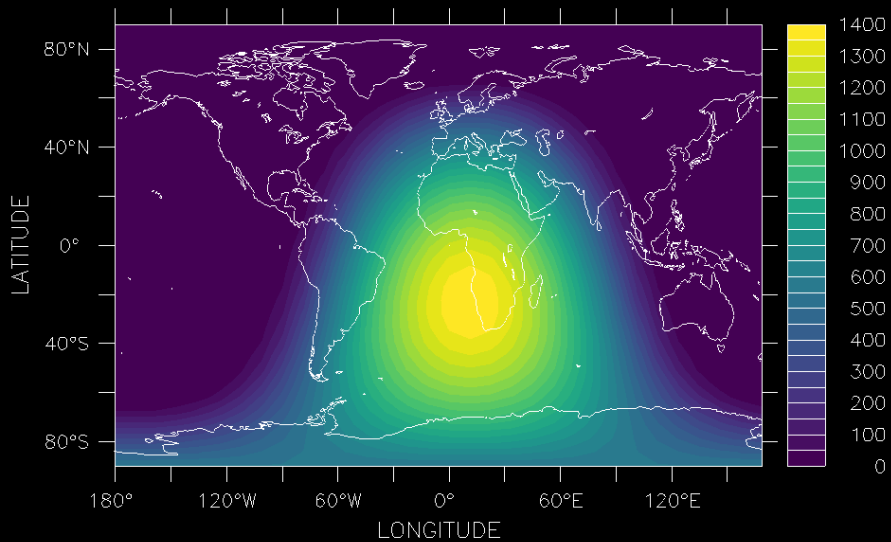
$$F(z)^\uparrow = B_{\nu}(T_s) \epsilon(0, z) + \int_0^z k_{\nu} \rho B_{\nu}(T(z')) \epsilon(z', z) dz'$$

$$F(z)^\uparrow = B_{\nu}(T_s) \epsilon(0, z) + \int_0^z B_{\nu}(T(z')) \frac{\partial \epsilon(z', z)}{\partial z'} dz'$$

$$Q = \frac{\partial T}{\partial z} = \frac{1}{\rho C_p} \frac{\partial F(z)^\uparrow}{\partial z}$$

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TIME : 03-JAN-1980 11:00 360_DAY DATA SET: histhf



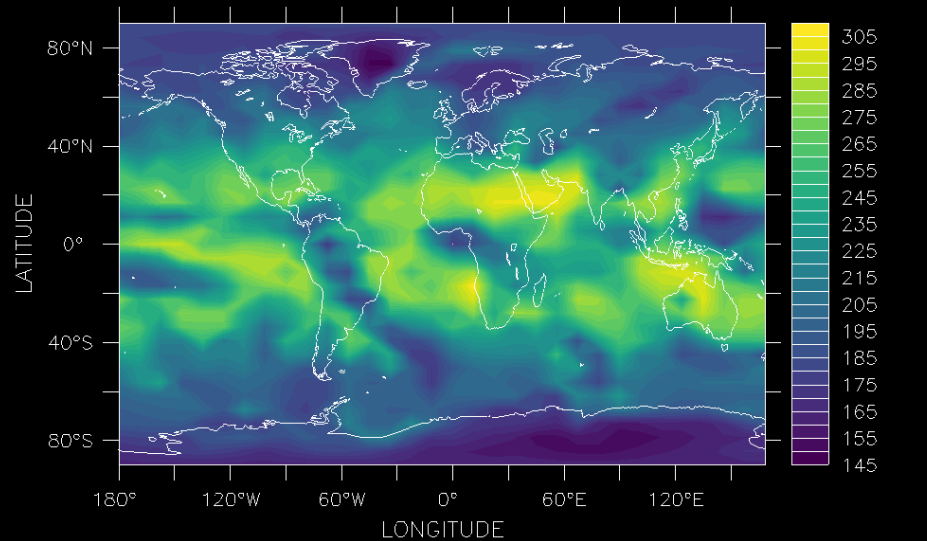
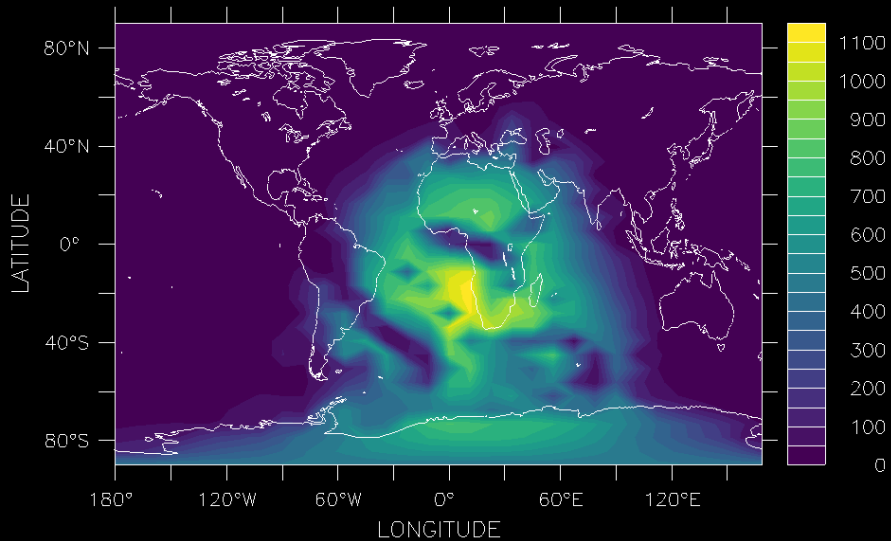
SWdn at TOA (W/m²)

SWup at TOA (W/m²)

reg/l=30
set v ul ; fill swdntoa ; go land
set v ur ; fill swuptoa ; go land
set v ll ; fill swdnsfc ; go land
set v lr ; fill topl ; go land

TIME : 03-JAN-1980 11:00 360_DAY DATA SET: histhf

TIME : 03-JAN-1980 11:00 360_DAY DATA SET: histhf



SWdn at surface (W/m²)

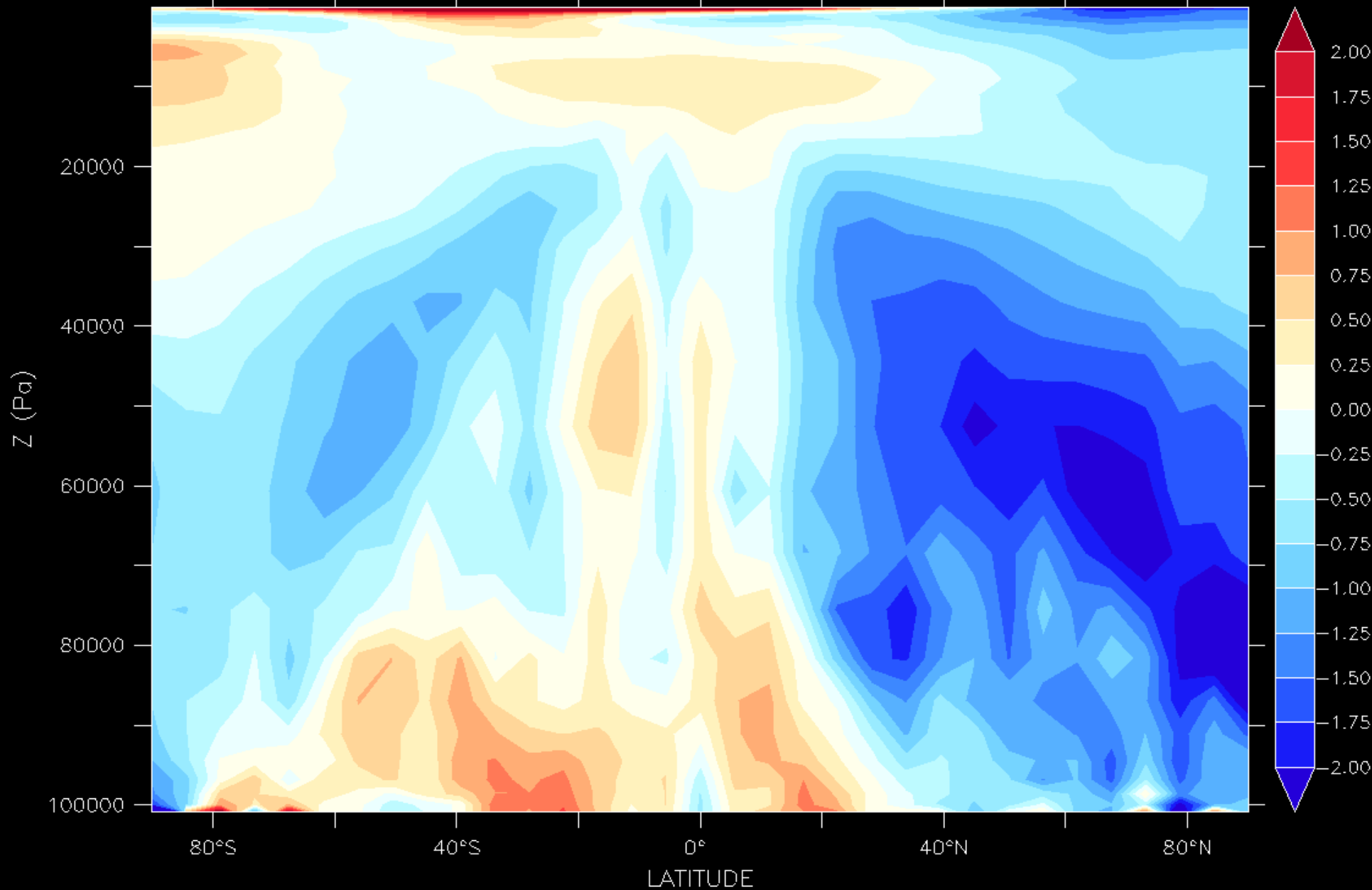
IR rad. at TOA (W/m²)

LONGITUDE : 174.375E(-185.625) to 174.375E
TIME : 03-JAN-1980 12:00 360_DAY

07-JAN-2024 19:19:01

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86400*dtphy[i=@ave]



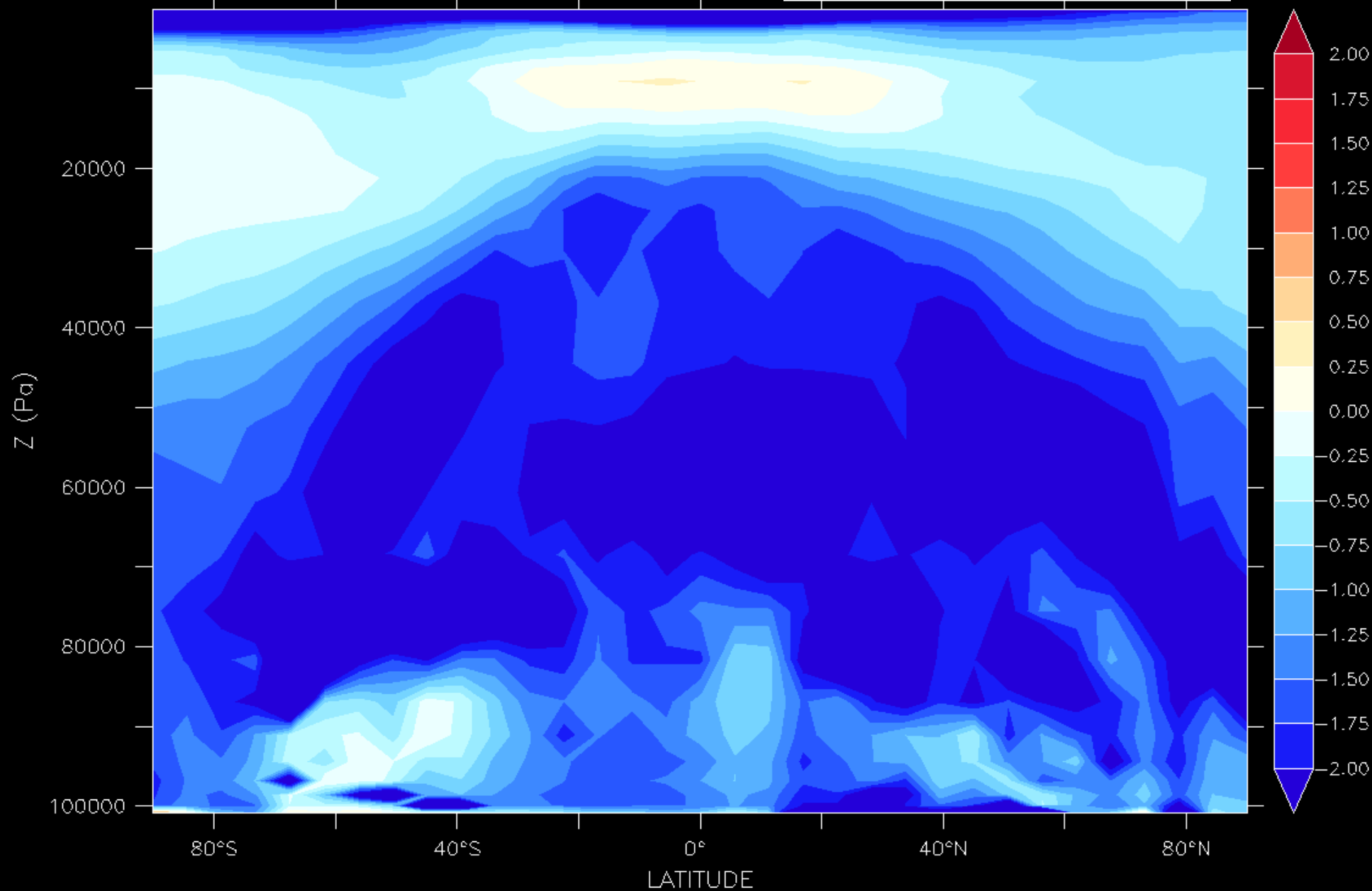
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LONGITUDE : 174.375E(-185.625) to 174.375E
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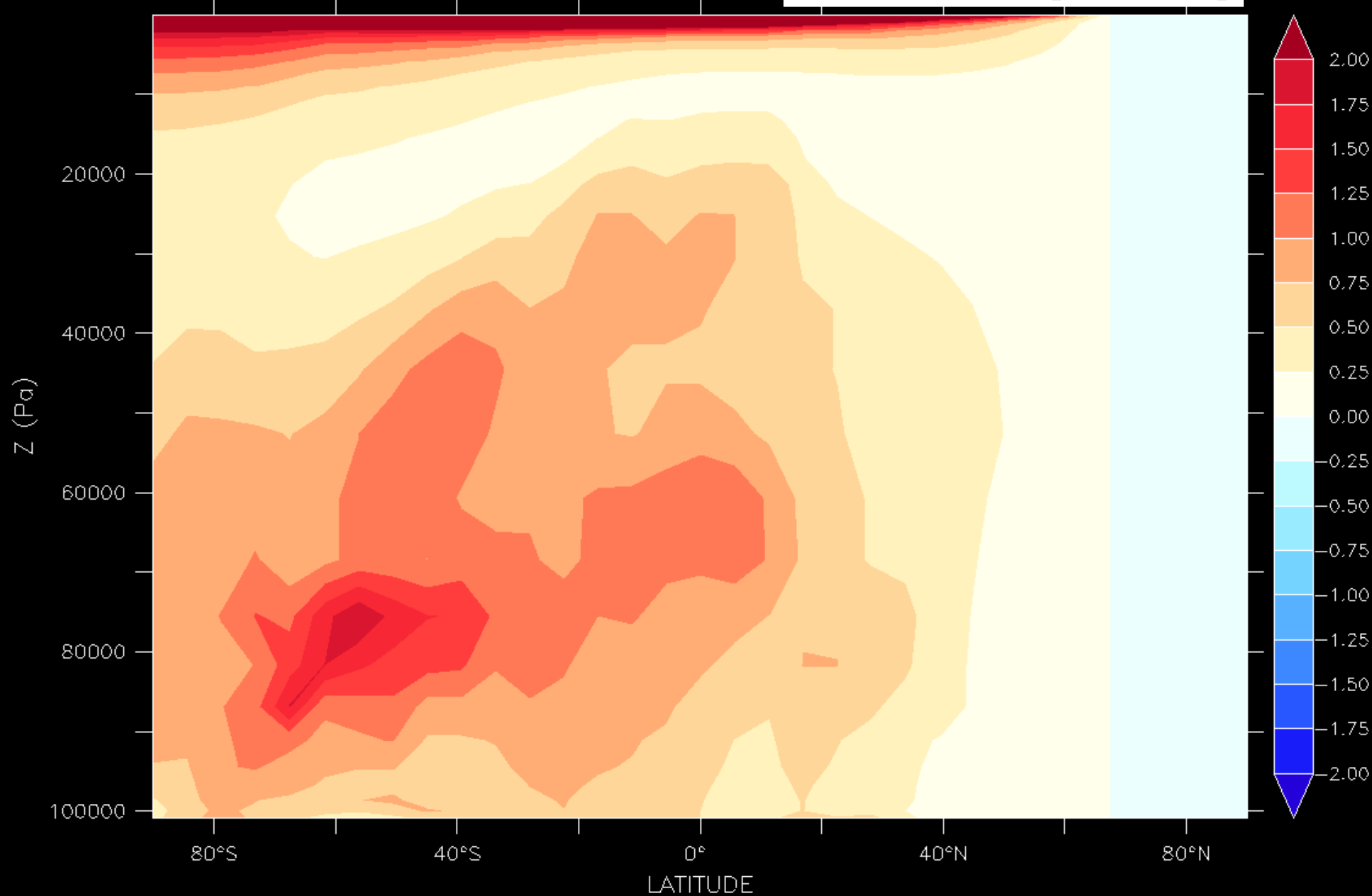
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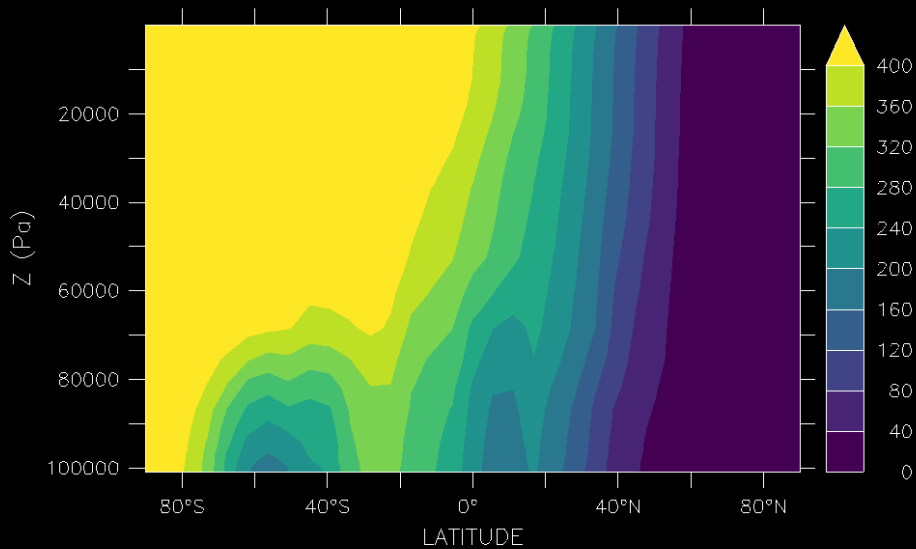
DATA SET: histday

86400*dtswr[i=@ave]



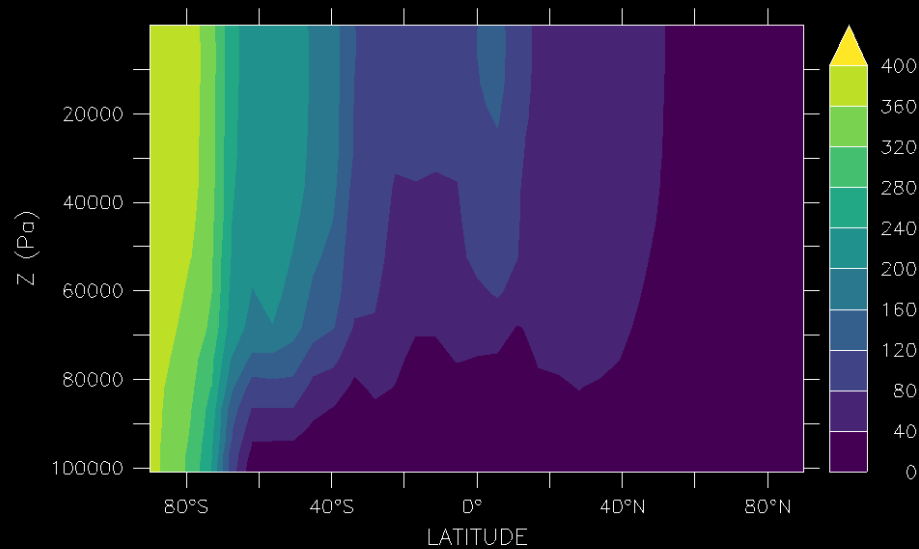
86400*DTSWR[I=@AVE]

LONGITUDE : 174.375E(-185.625) to 174.375E (averaged)
TIME : 03-JAN-1980 12:00 360_DAYDATA SET: histday



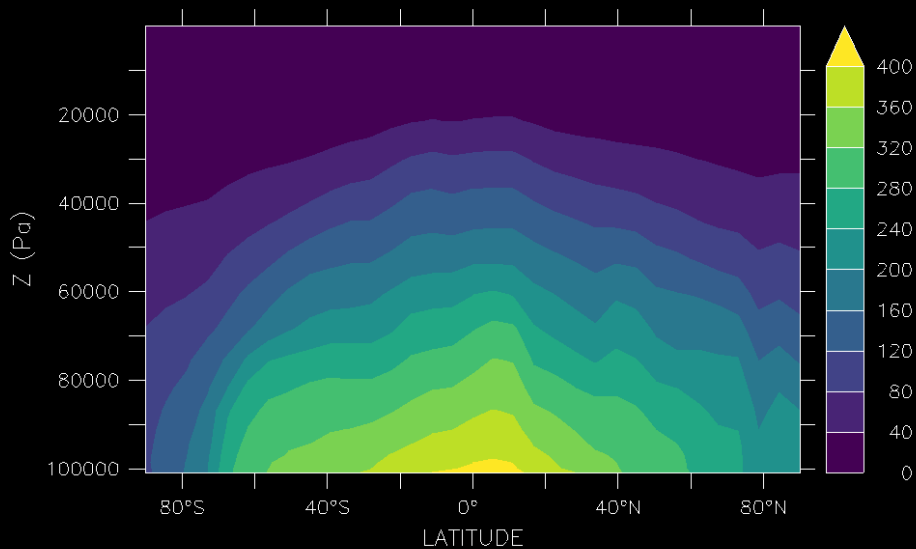
SW downward radiation ($W m^{-2}$)

LONGITUDE : 174.375E(-185.625) to 174.375E (averaged)
TIME : 03-JAN-1980 12:00 360_DAYDATA SET: histday



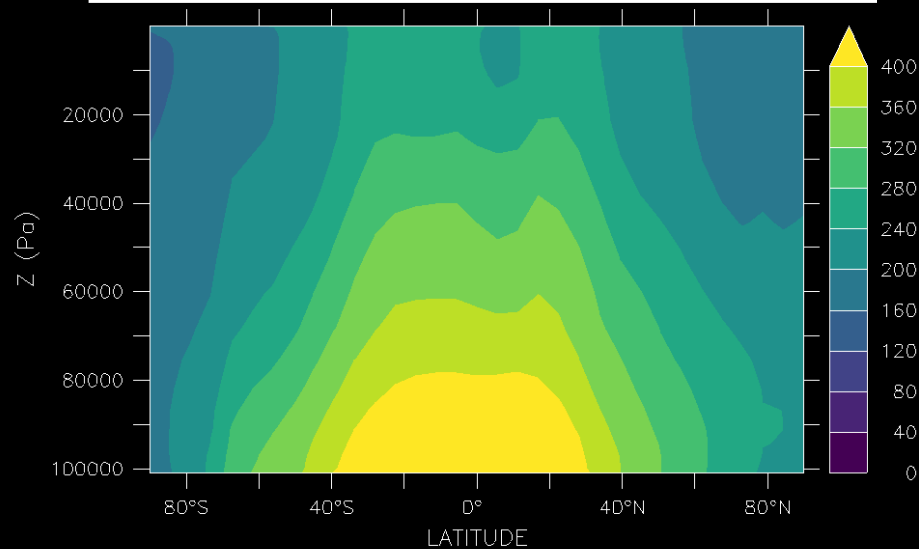
SW upward radiation ($W m^{-2}$)

LONGITUDE : 174.375E(-185.625) to 174.375E
TIME : 03-JAN-1980 12:00 360_DAYDATA SET: histday



$-1*RLD[I=@AVE]$

```
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set v ur ; fill/lev=(0,400,40)(Inf) rsu[i=@ave]  
set v ll ; fill/lev=(0,400,40)(Inf) -1*rls[i=@ave]  
set v lr ; fill/lev=(0,400,40)(Inf) rlu[i=@ave]
```



LW upward radiation ($W m^{-2}$)

Parameterization of subgrid-scale motions

- Reynolds decomposition
- Turbulence
- Boundary layer convection
- Deep convection
- Subgrid-scale orography

Based on the Reynolds decomposition between

- large-scale/resolved/explicit variables (dynamical core)
- subgrid-scale/unresolved/turbulent fluctuations (parameterizations)

Reynolds decomposition

\tilde{X} : "average" or "large scale" variable

$\bar{X} = \tilde{\rho} \mathbf{v} / \tilde{\rho}$: air mass weighted "average"

$X = \tilde{X} + X'$: X' , turbulent fluctuation

$$\begin{aligned} \Rightarrow \rho \tilde{\mathbf{v}} c &= \rho (\bar{\mathbf{v}} + \widetilde{\mathbf{v}'}) (\bar{c} + c') \\ &= \tilde{\rho} \bar{\mathbf{v}} \bar{c} + \tilde{\rho} \overline{\mathbf{v}' c'} \end{aligned}$$

$$\frac{\partial \rho c}{\partial t} + \widetilde{\text{div}(\rho \mathbf{v} c)} = 0 \quad \Rightarrow \quad \frac{\partial \tilde{\rho} \bar{c}}{\partial t} + \text{div}(\tilde{\rho} \bar{\mathbf{v}} \bar{c}) + \text{div}(\tilde{\rho} \overline{\mathbf{v}' c'}) = 0$$

$$\frac{\partial \bar{c}}{\partial t} + \mathbf{v} \cdot \mathbf{grad} \bar{c} = -\frac{1}{\rho} \text{div}(\rho \overline{\mathbf{v}' c'}) = -\frac{1}{\rho} \frac{\partial \overline{\rho w' c'}}{\partial z}$$

Parameterization of subgrid-scale motions

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$$Dq/Dt = Sq$$

$$\frac{\partial c}{\partial t} + \mathbf{v} \cdot \mathbf{grad} c = -\frac{1}{\rho} \text{div}(\rho \overline{\mathbf{v}' c'}) = -\frac{1}{\rho} \frac{\partial \overline{\rho w' c'}}{\partial z}$$

Turbulent diffusion : bases

**Boundary layer approximation (horizontal homogeneity)
+ eddy diffusion**

$$\overline{w'c'} = -K_z \frac{\partial c}{\partial z} \quad \longrightarrow \quad \frac{\partial c}{\partial t} = \frac{1}{\rho} \frac{\partial}{\partial z} \left(\rho K_z \frac{\partial c}{\partial z} \right)$$

- Analogy with molecular viscosity
(Brownian motion \leftrightarrow turbulence)
- Down-gradient fluxes.
- Turbulence acts as a "mixing"

Turbulent diffusion : Mellor et Yamada

Turbulent diffusivity K_z

- Prandtl (1925) mixing length : $K_z = l|\overline{w'}|$ or $K_z = l^2 \frac{\partial \|\mathbf{v}\|}{\partial z}$
- Accounting for static stability (Ex. Louis 1979)

$$K_z = f(Ri)l^2 \left| \frac{\partial \mathbf{v}}{\partial z} \right|, \text{ in } \quad \text{with } Ri = \frac{g}{\theta} \frac{\frac{\partial \theta}{\partial z}}{\left(\frac{\partial \mathbf{v}}{\partial z} \right)^2} \quad (1)$$

- Turbulent kinetic energy $\overline{w'^2} \simeq e = \frac{1}{2} \left[\overline{u'^2} + \overline{v'^2} + \overline{w'^2} \right]$

$$\frac{\partial e}{\partial t} = -\overline{w'u'} \frac{\partial u}{\partial z} - \overline{w'v'} \frac{\partial v}{\partial z} + \frac{g}{\theta} \overline{w'\theta'} - \frac{1}{\rho} \frac{\partial \overline{w'p'}}{\partial z} - \frac{\partial \overline{w'e}}{\partial z} - \epsilon$$

In LMDZ : Mellor and Yamada (Yamada 1983 version, see also Vignon et al. publications)

Turbulent diffusion : coupling with surface

$$\frac{\partial c}{\partial t} = -\frac{1}{\rho} \frac{\partial F_c(z)}{\partial z}$$

$$F_c(z > 0) = -K_z \rho \frac{\partial c}{\partial z}$$

At surface :

$F_c(z = 0)$: imposed or

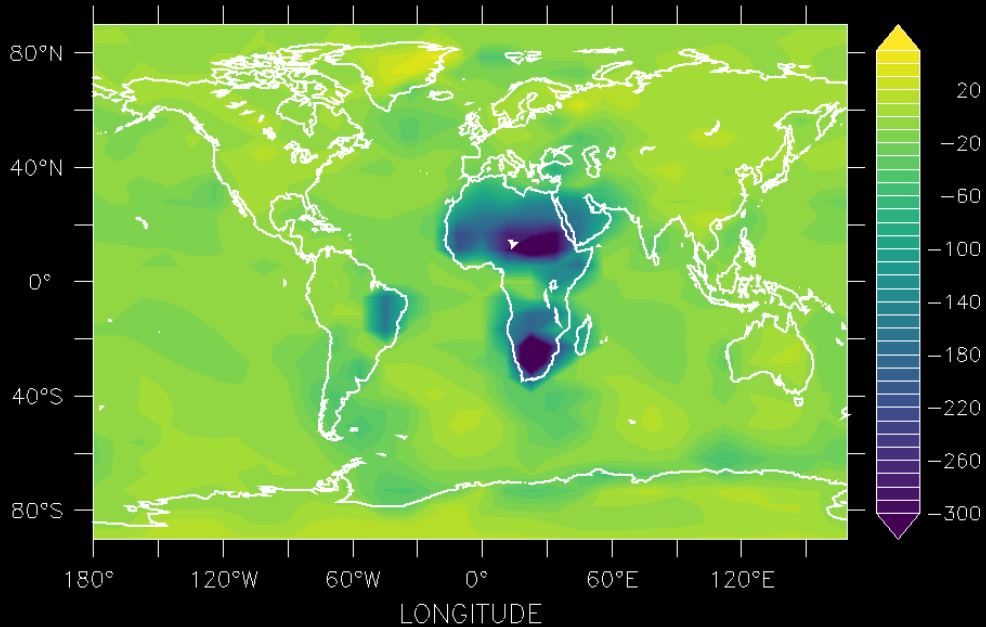
$$F_c(z = 0) = \rho C_d ||V|| (c_s - c_1)$$

Where c_s and c_1 are values of c at the surface and in the first model layer respectively

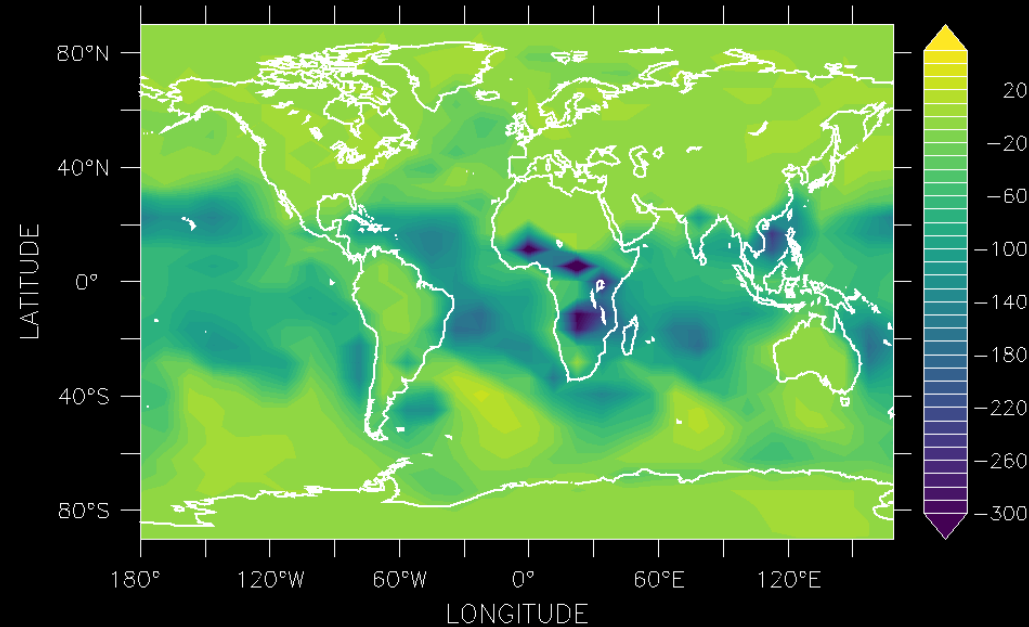
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set v lr ; fill/lev=(-Inf)(-300,50,10)(Inf) flat ; go land thick
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TIME : 03-JAN-1980 11:00 360_DAY DATA SET: histhf



Sensible heat flux (W/m²)



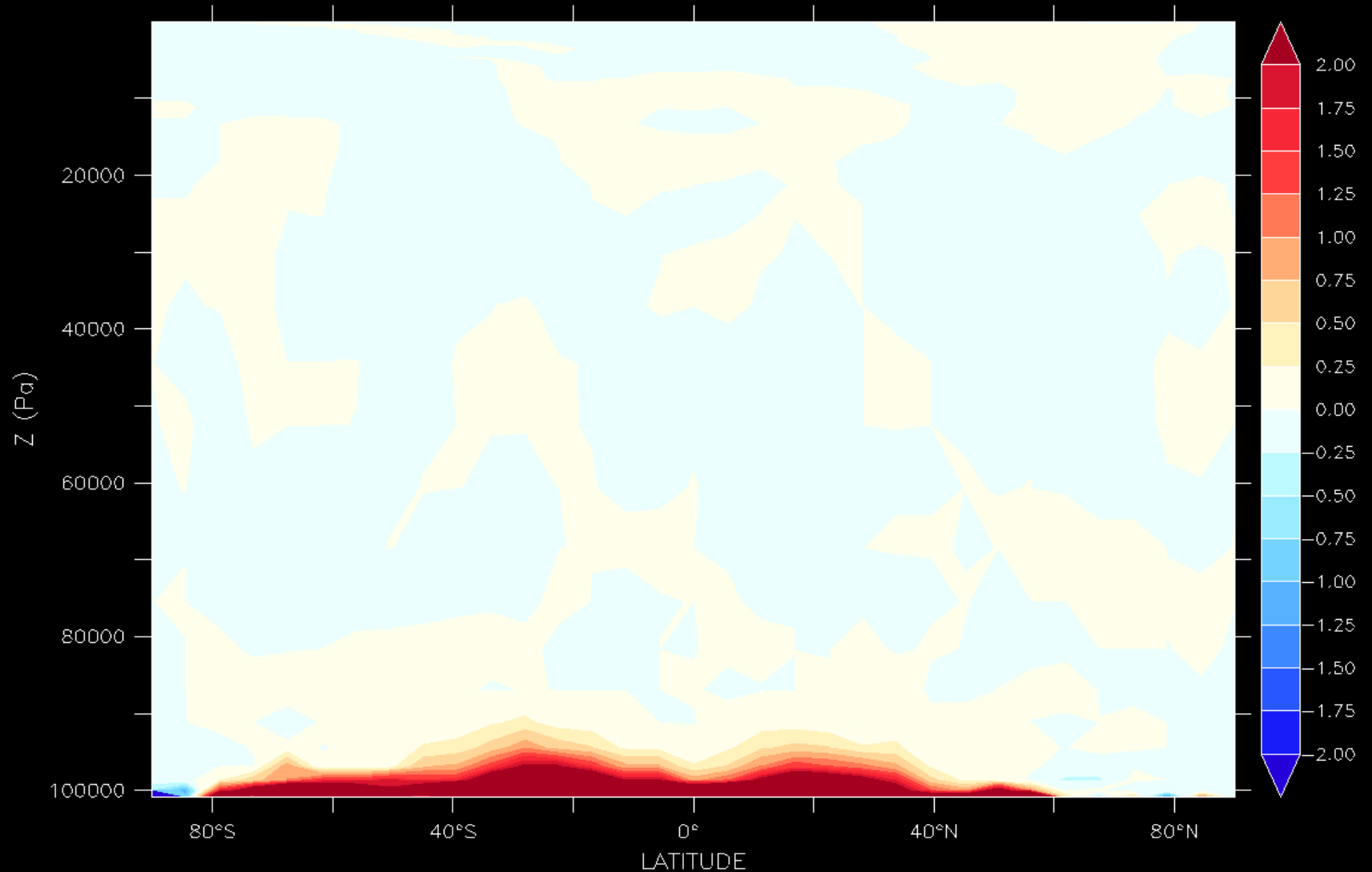
Latent heat flux (W/m²)

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07-JAN-2024 19:43:00

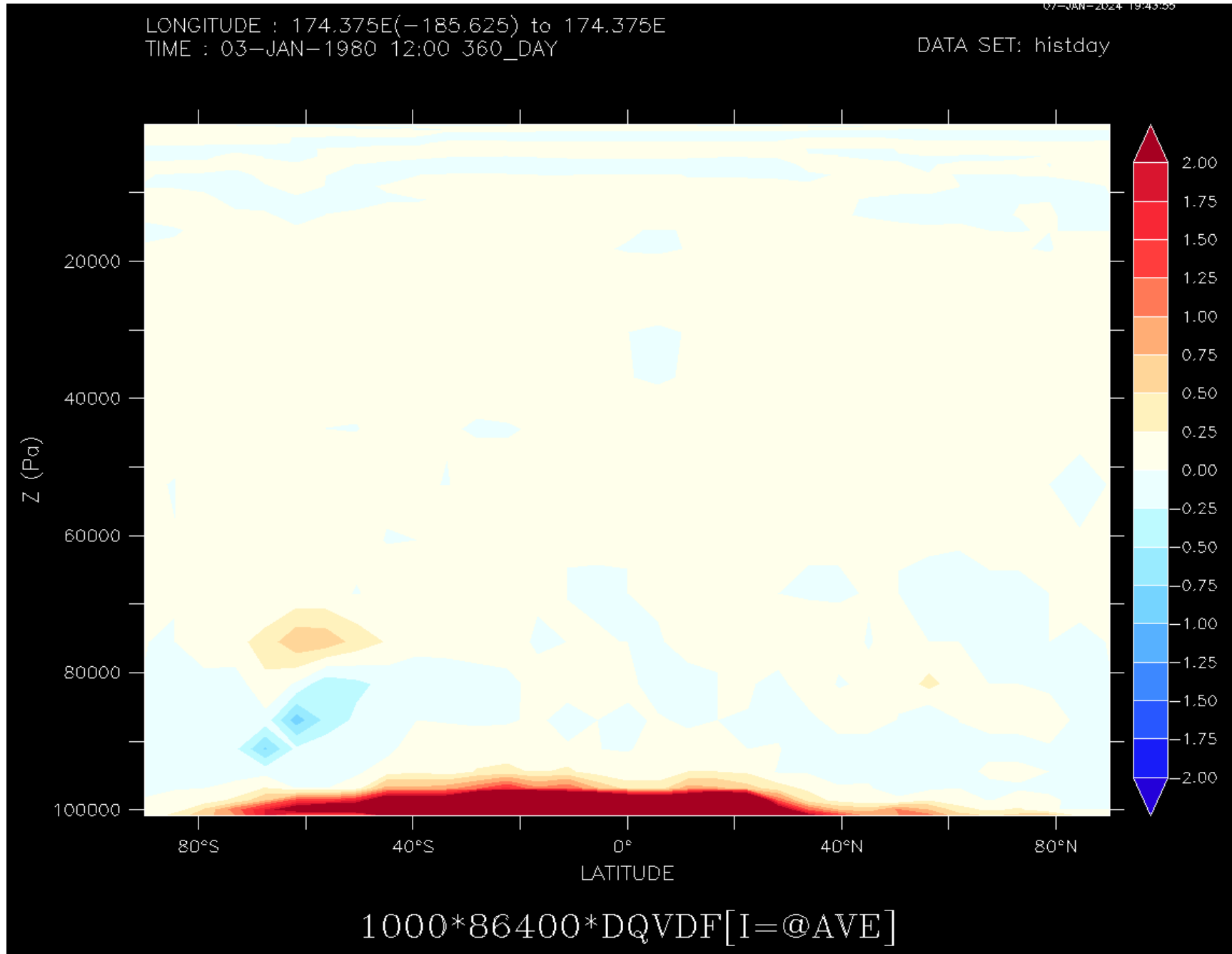
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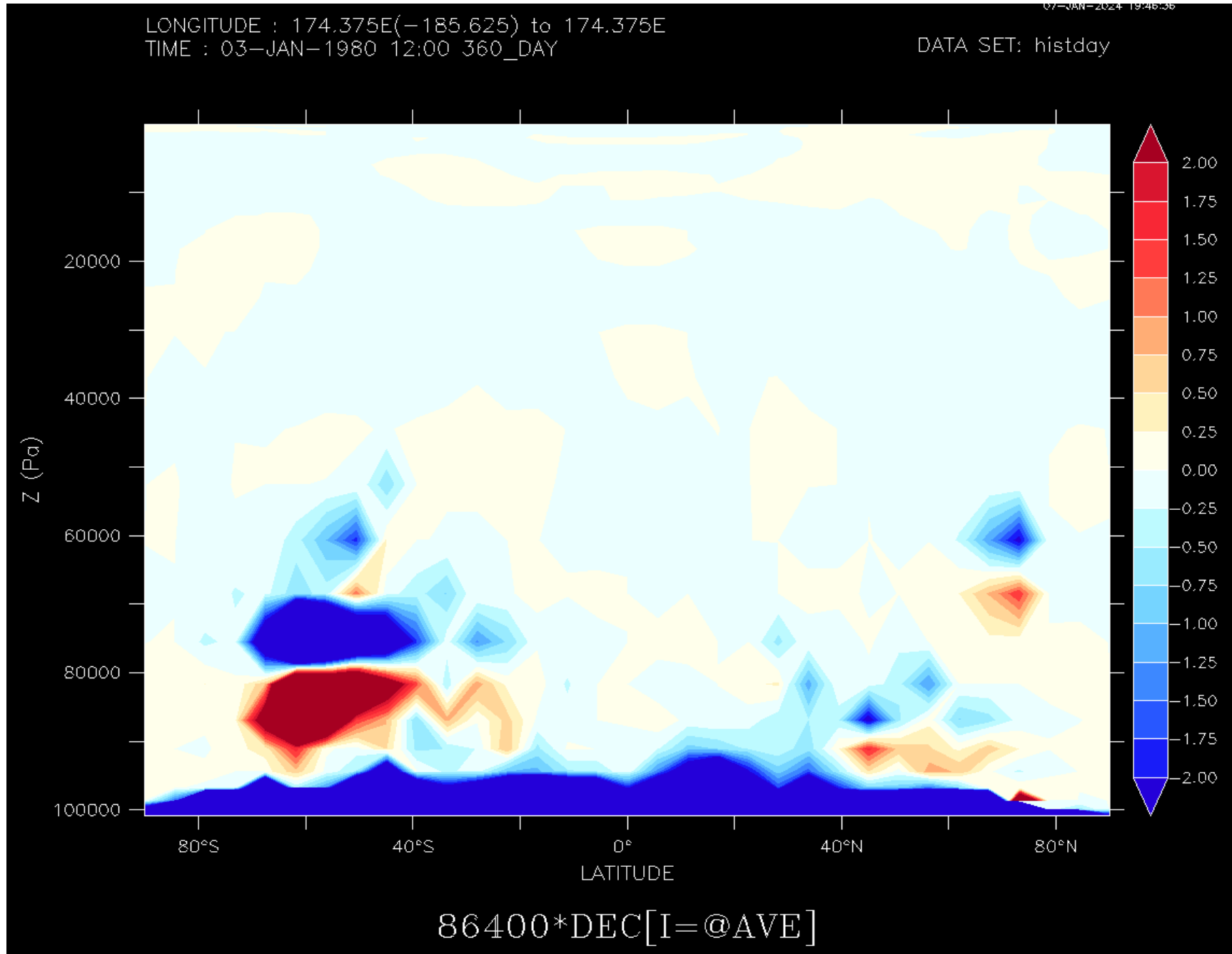
86400*DTVDF[I=@AVE]

fill/lev=(-Inf)(-2,2,0.25)(inf)/pal=blue_darkred 1000*86400*dqvdf[i=@ave]



let dec=vitu*duvdf+vitv*dvddf

yes? fill/lev=(-Inf)(-2,2,0.25)(inf)/pal=blue_darkred 86400*dec[i=@ave]



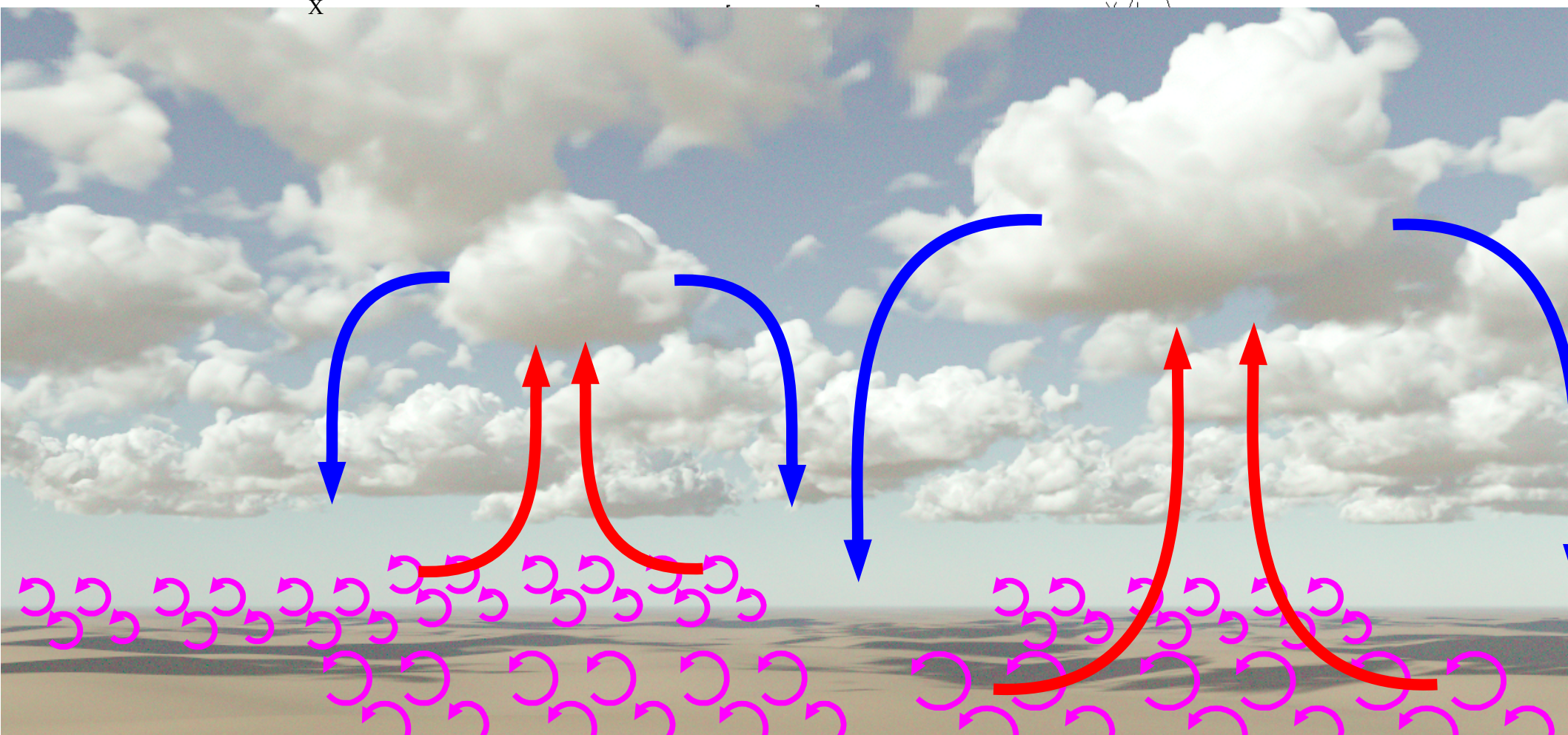
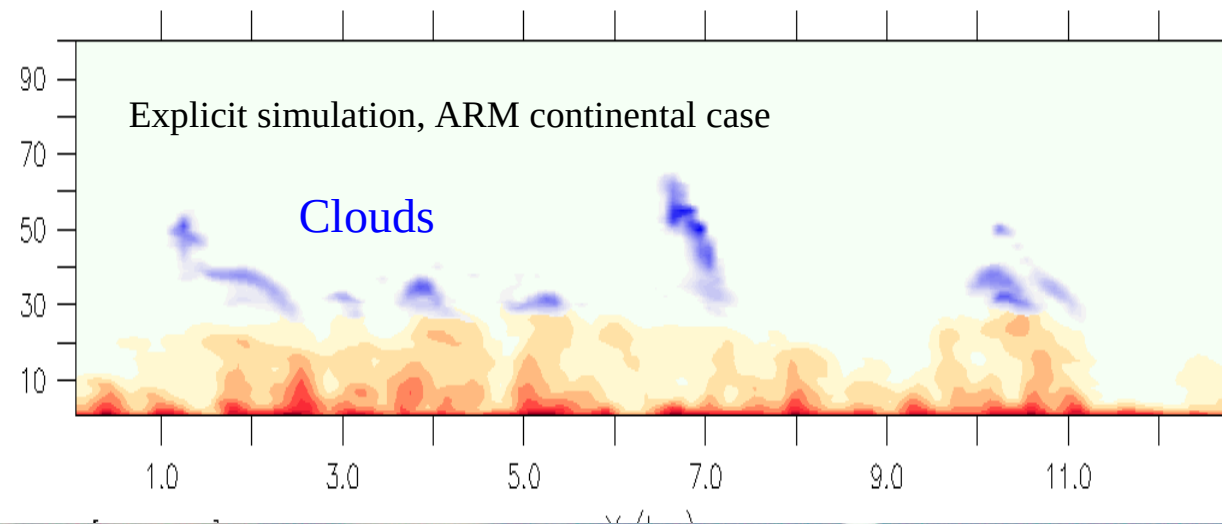
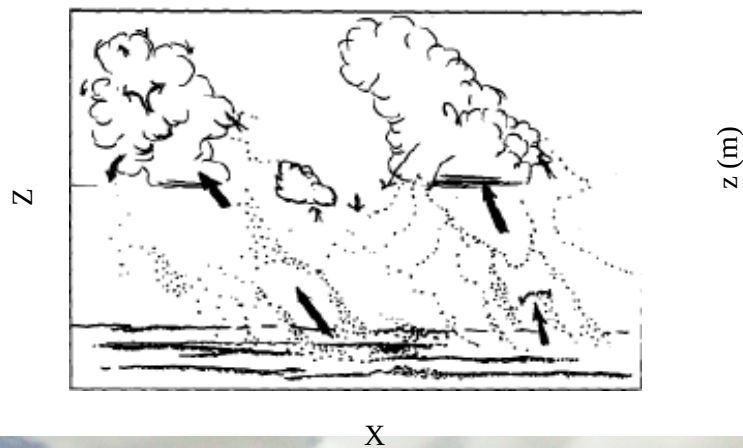
Parameterization of the convective boundary layer

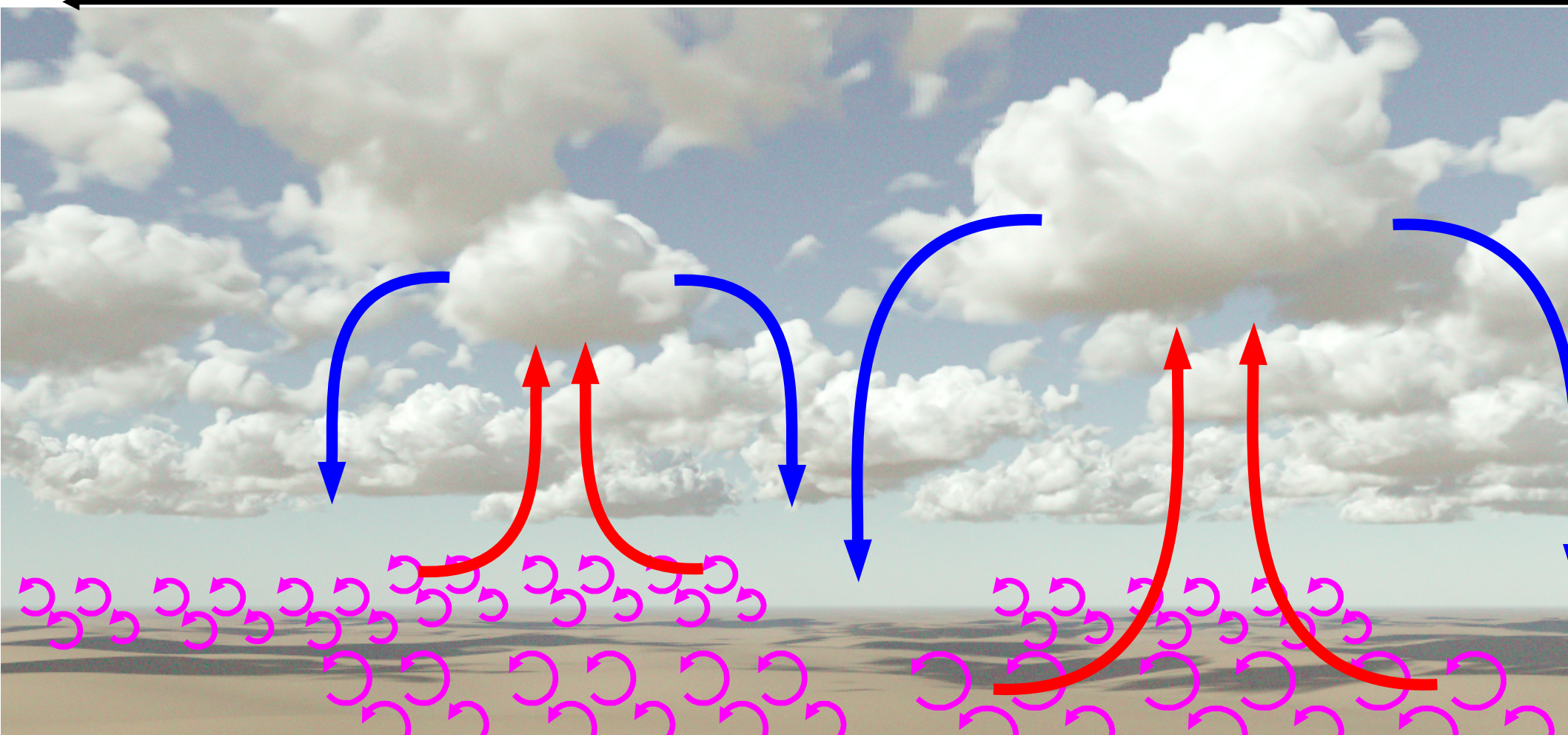
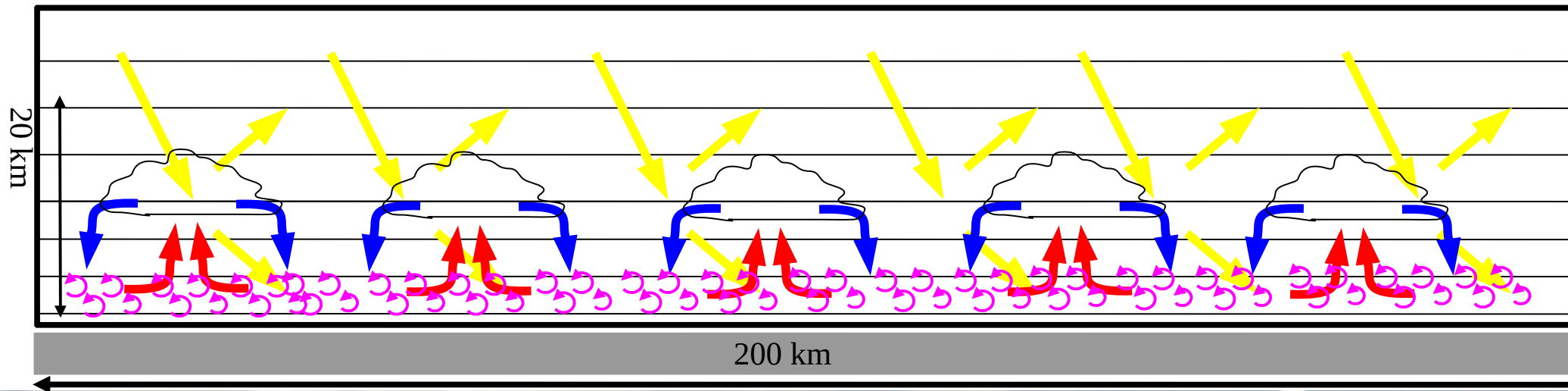
Based on high resolution explicit (LES) simulations

Movie entirely based on physics : MesoNH LES and physical rendering with htrdr (Villefranque)



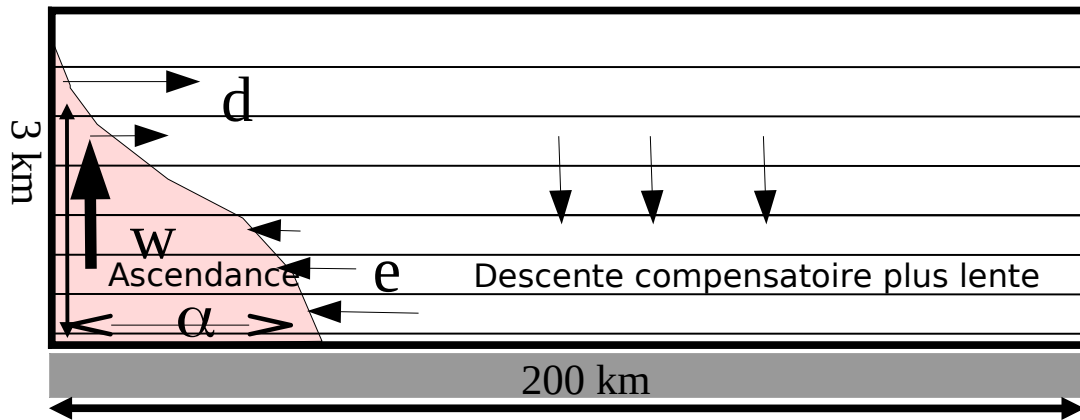
LeMone and Pennell, MWR, 1976





2. Couche limite convective

Le modèle du thermique



Variables internes de la paramétrisation :

w : vitesse moyenne des panaches ascendants

α : fraction de la surface couverte par les ascendances

e : taux d'entrée latérale d'air dans le panache (entraînement)

d : sorties d'air depuis le panache (déentraînement)

q_a : concentration du composant q dans l'ascendance

Terme source pour les équations explicites

$$S_q = -\frac{1}{\rho} \frac{\partial}{\partial z} \overline{\rho w' q'} = \frac{1}{\rho} \frac{\partial}{\partial z} \rho K_z \frac{\partial q}{\partial z} - \frac{1}{\rho} \frac{\partial}{\partial z} [\rho \alpha w (q_a - q)]$$

Diffusion turbulente

Transport par le modèle de panache

4 Paramètres libres :

$$a_1 = \frac{2}{3}, \beta_1 = 0.9, b = 0.002, c = 0.012 m^{-1}, d = 0.5$$

Conservation de la masse :

$$\frac{\partial f}{\partial z} = e - d \quad \text{avec } f = \alpha \rho w$$

Conservation de la masse du composant q

$$\frac{\partial f q_a}{\partial z} = e q - d q_a$$

Equation du mouvement

$$\frac{\partial f w}{\partial z} = -d w + \alpha \rho B$$

et la poussée d'Archimède

$$B = g \frac{\theta_{va} - \theta_v}{\theta_v}$$

$$e = f \max(0, \frac{\beta_1}{1 + \beta_1} (a_1 \frac{B}{w^2} - b))$$

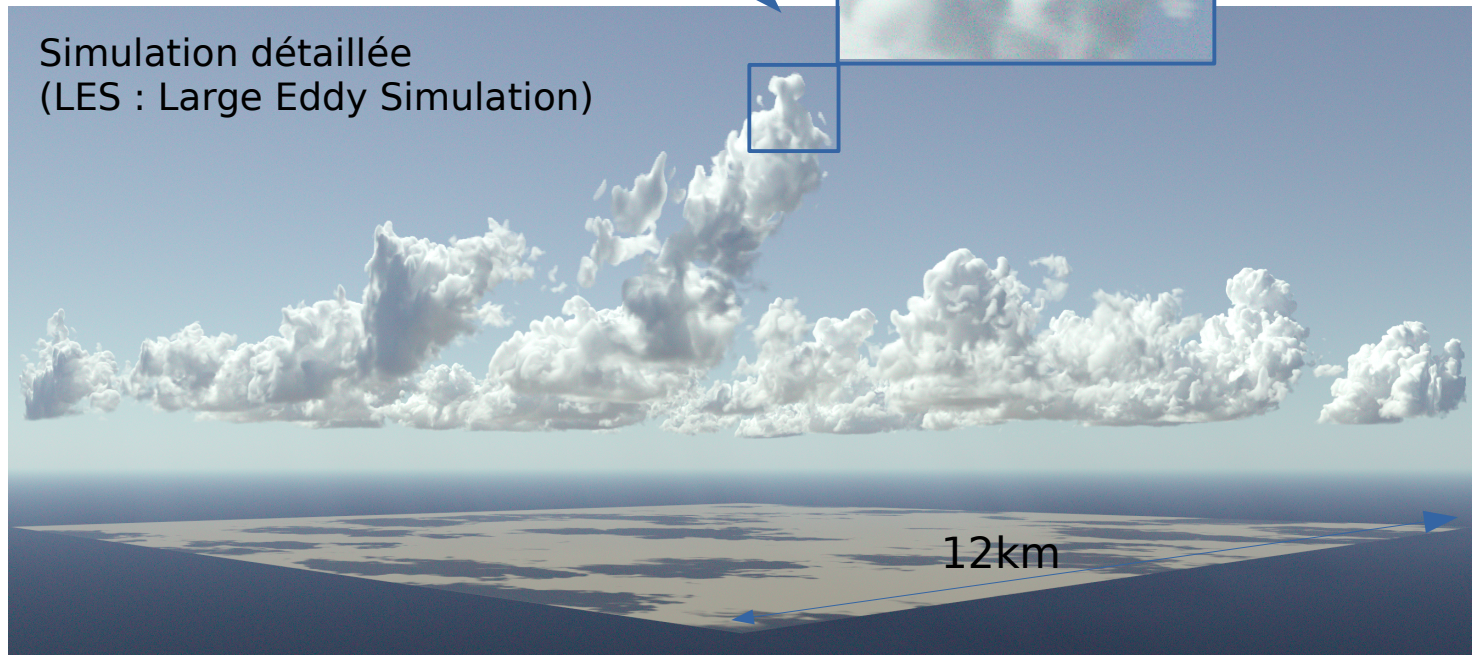
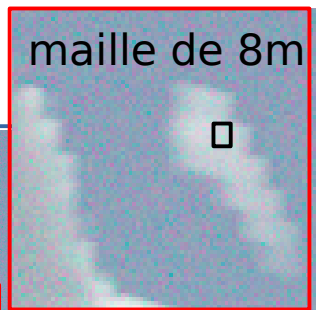
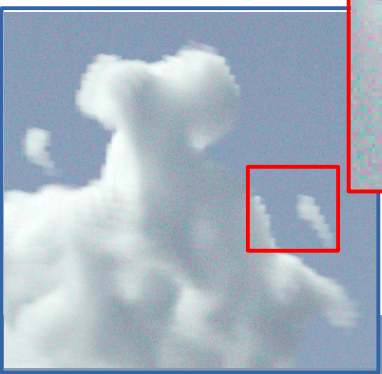
$$d = f \max(0, -\frac{a_1 \beta_1}{1 + \beta_1} \frac{B}{w^2} + c (\frac{q_a - q}{w^2})^d)$$

Etc ...



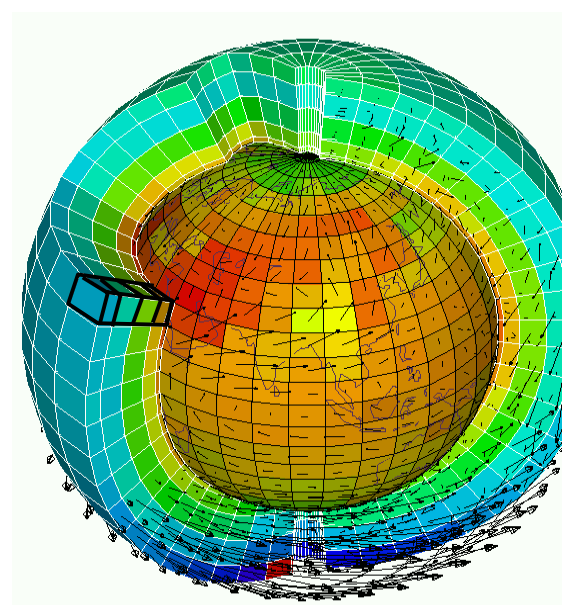
Campagne d'observation

Evaluation



Simulation détaillée
(LES : Large Eddy Simulation)

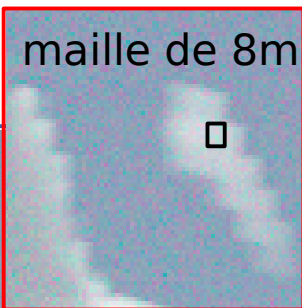
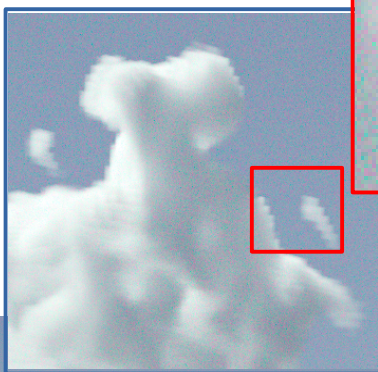
12km



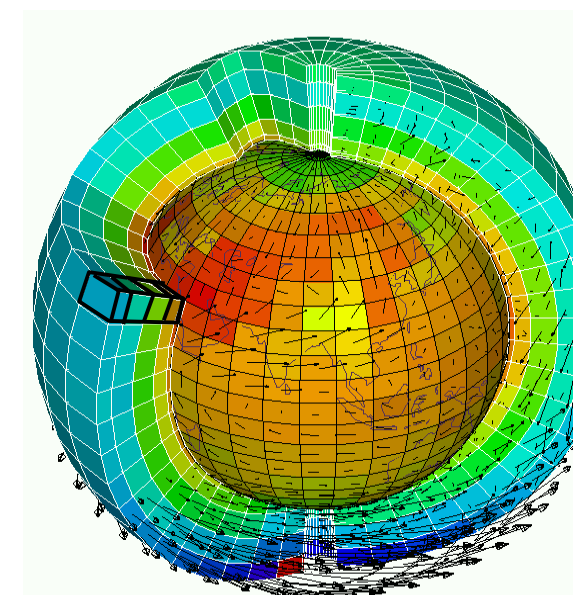
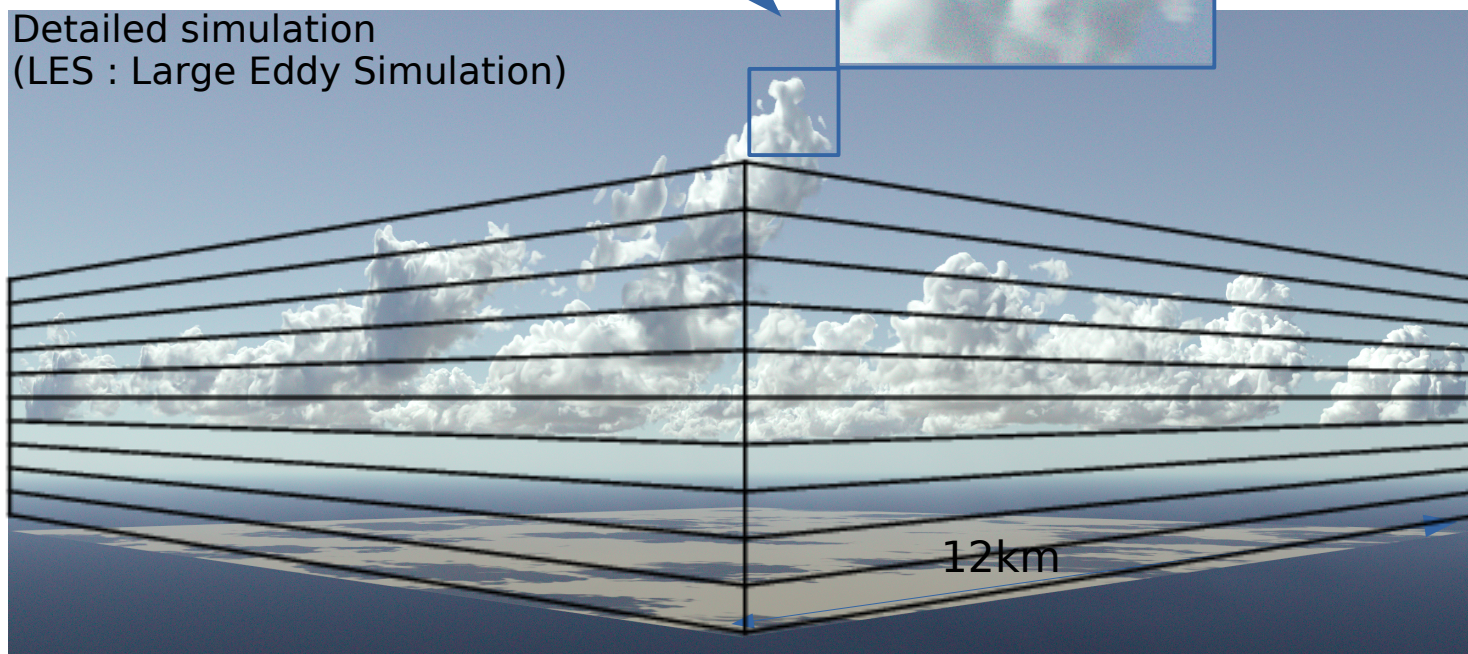


Field campaign experiment

Evaluation



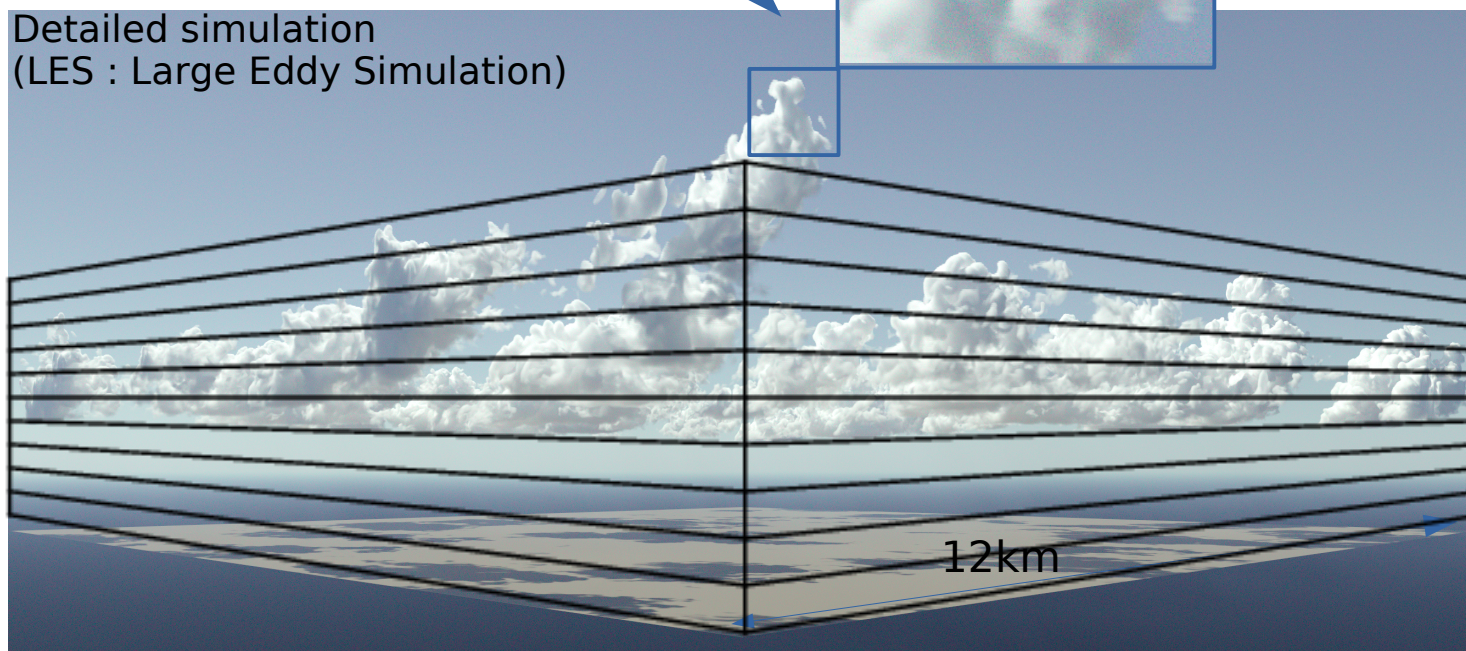
Detailed simulation
(LES : Large Eddy Simulation)



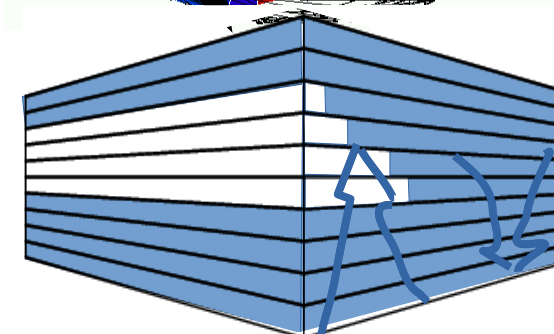
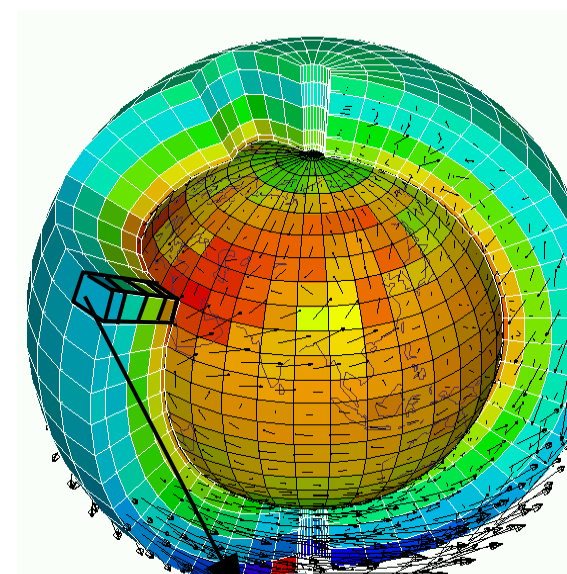
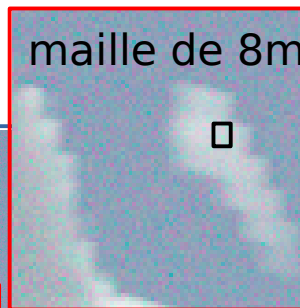


Field campaign experiment

Detailed simulation
(LES : Large Eddy Simulation)



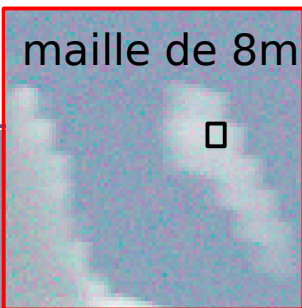
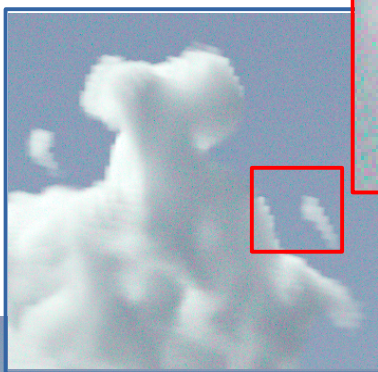
maille de 8m



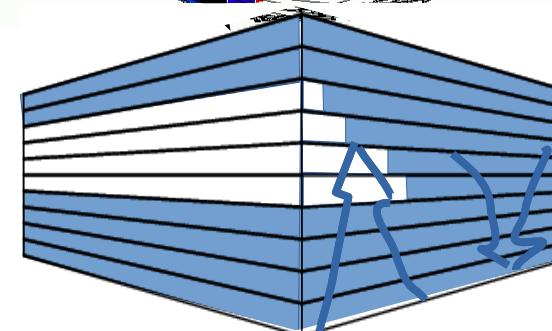
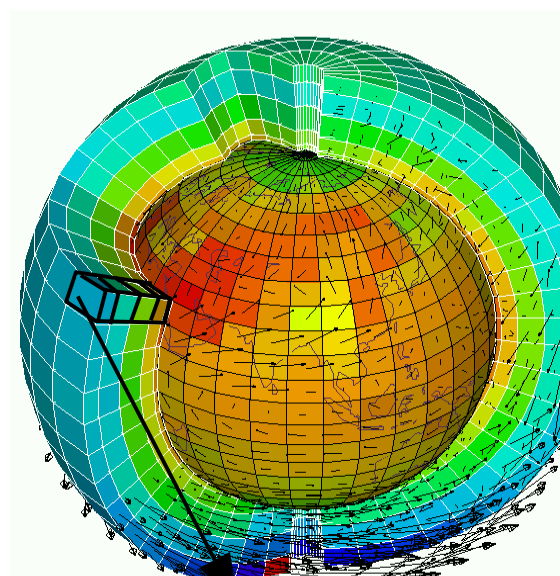
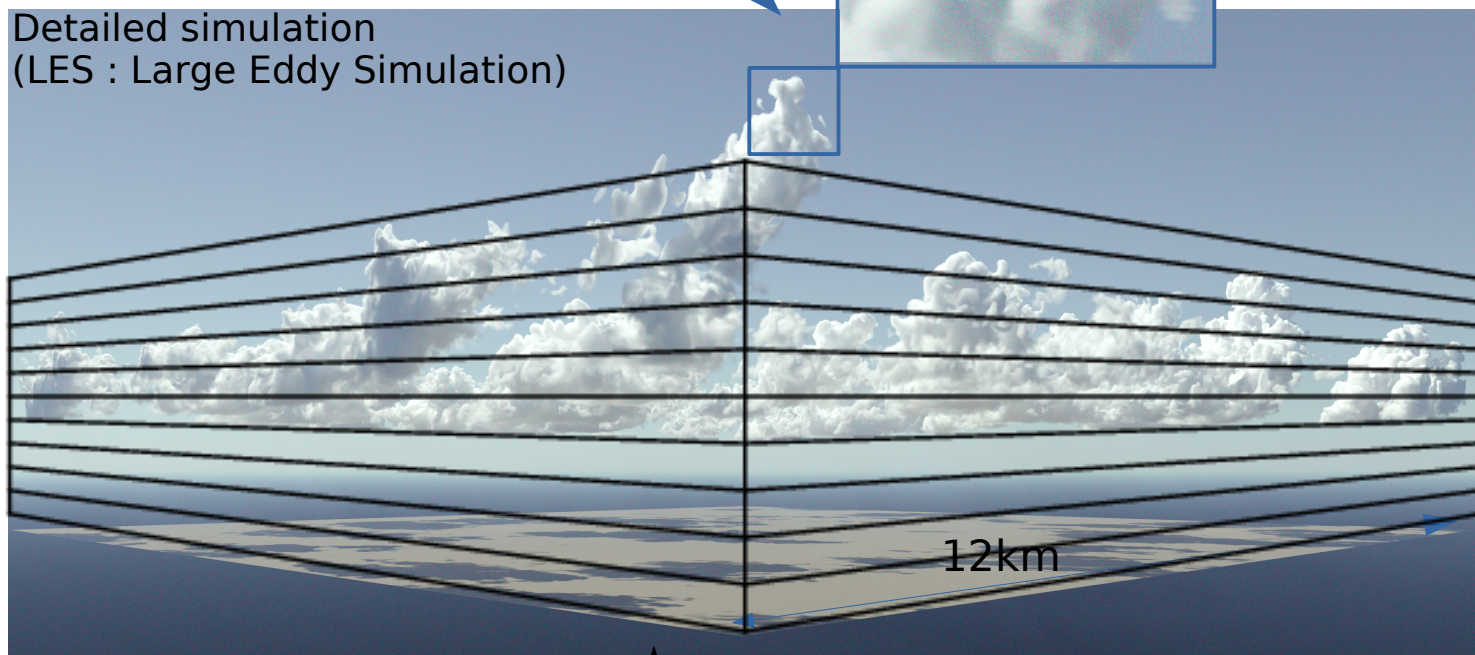
**Building a model of a
convectif plume and
associated clouds.
Trying to represent an
idealized mean cloud**



Evaluation

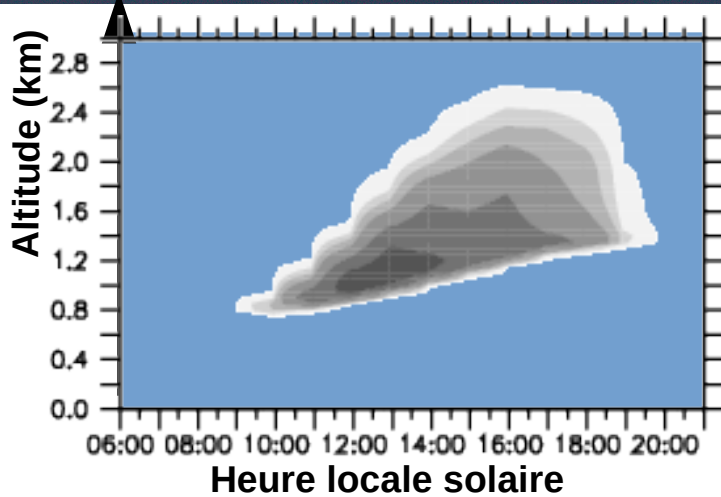


Detailed simulation
(LES : Large Eddy Simulation)

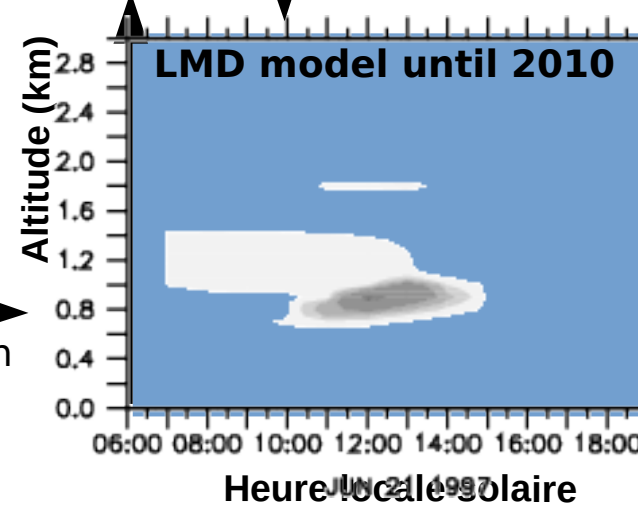


Computing the cloud fraction for each cell in one column of the GCM

Computing at each altitude the fraction of the horizontal domain covered by clouds. Also called : **the cloud fraction**



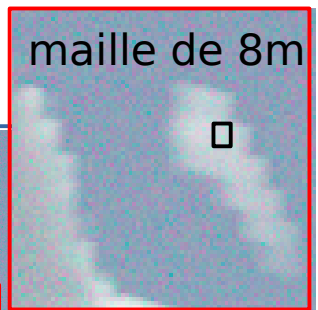
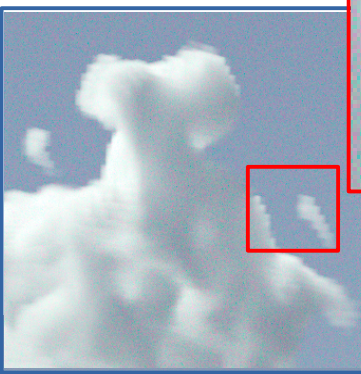
Evaluation



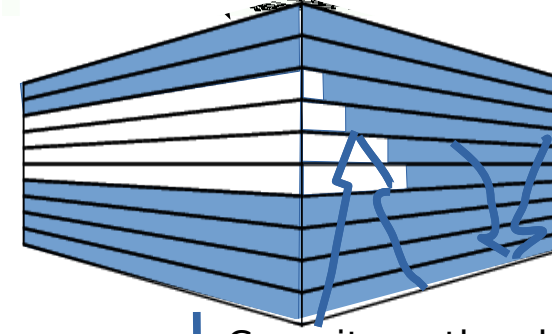
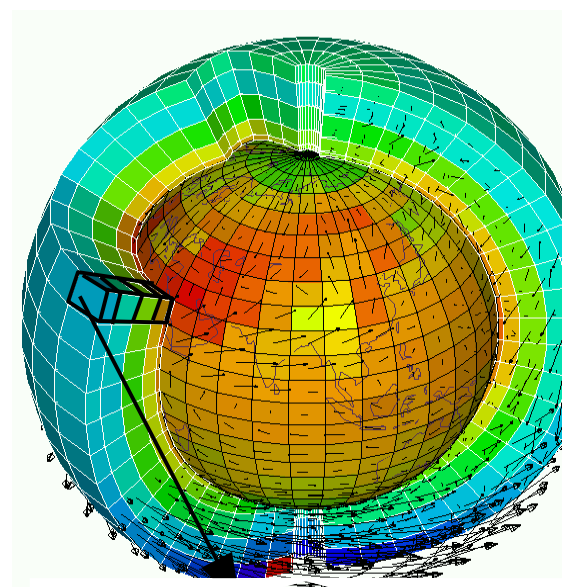
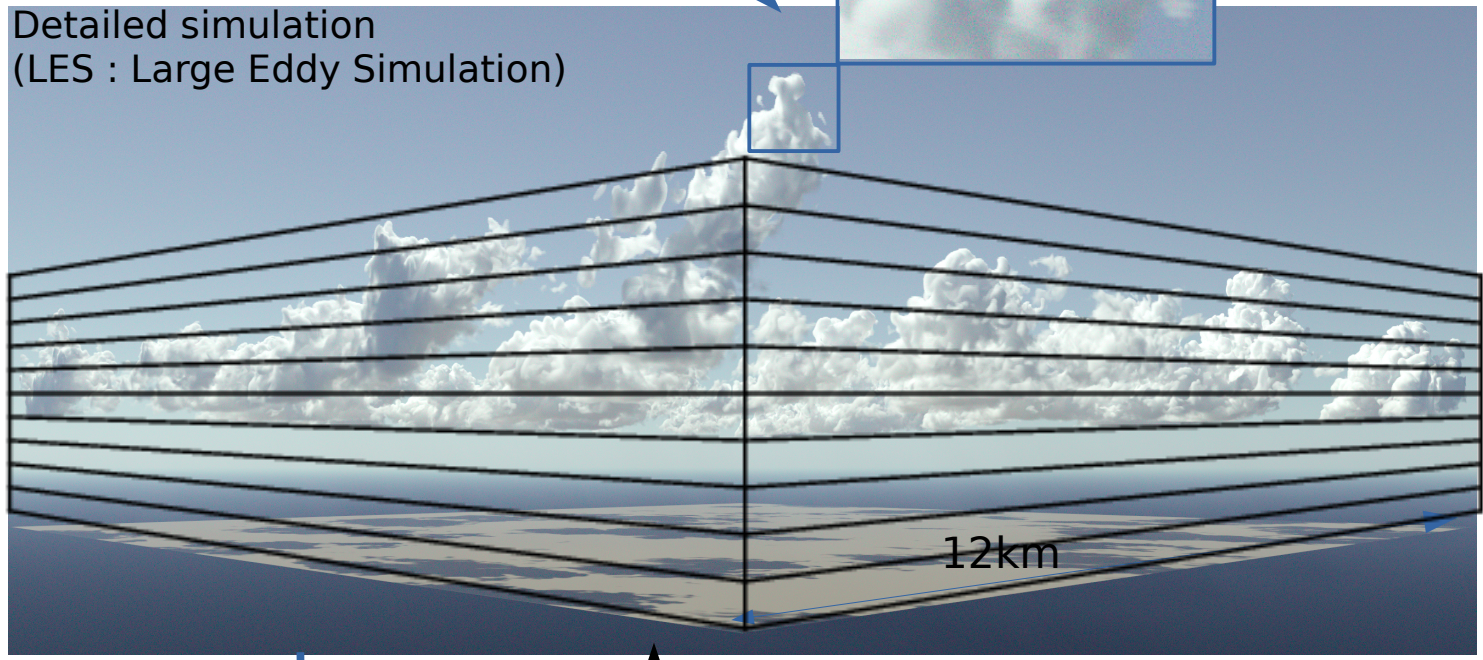
JUN 21 1997



Evaluation

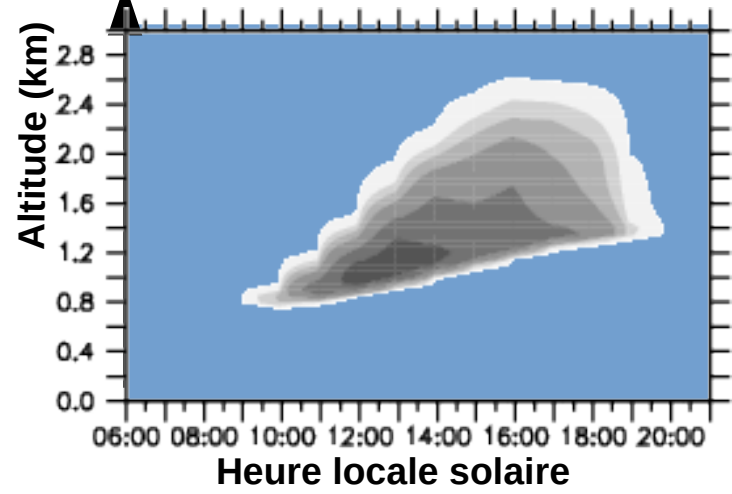


Detailed simulation
(LES : Large Eddy Simulation)

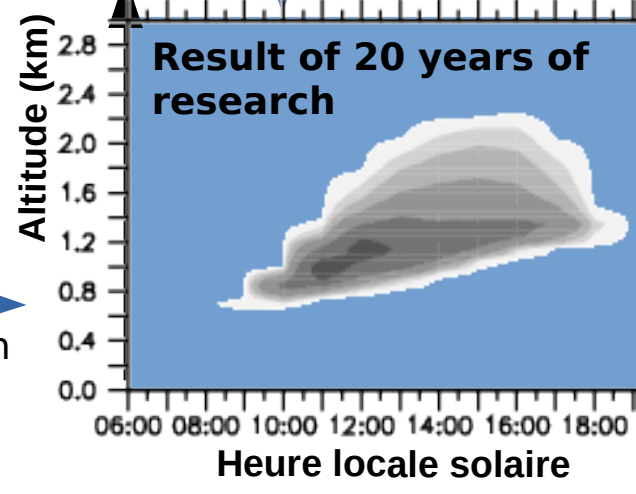


Computing the cloud fraction for each cell in one column of the GCM

Computing at each altitude the fraction of the horizontal domain covered by clouds. Also called : **the cloud fraction**



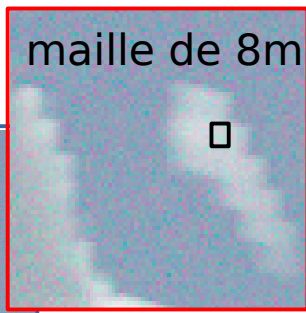
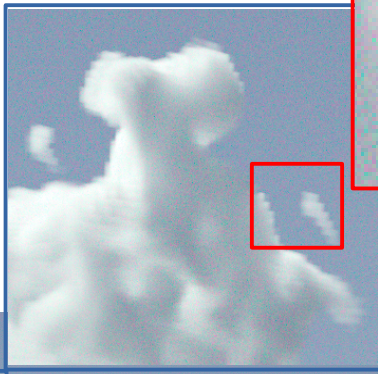
Evaluation



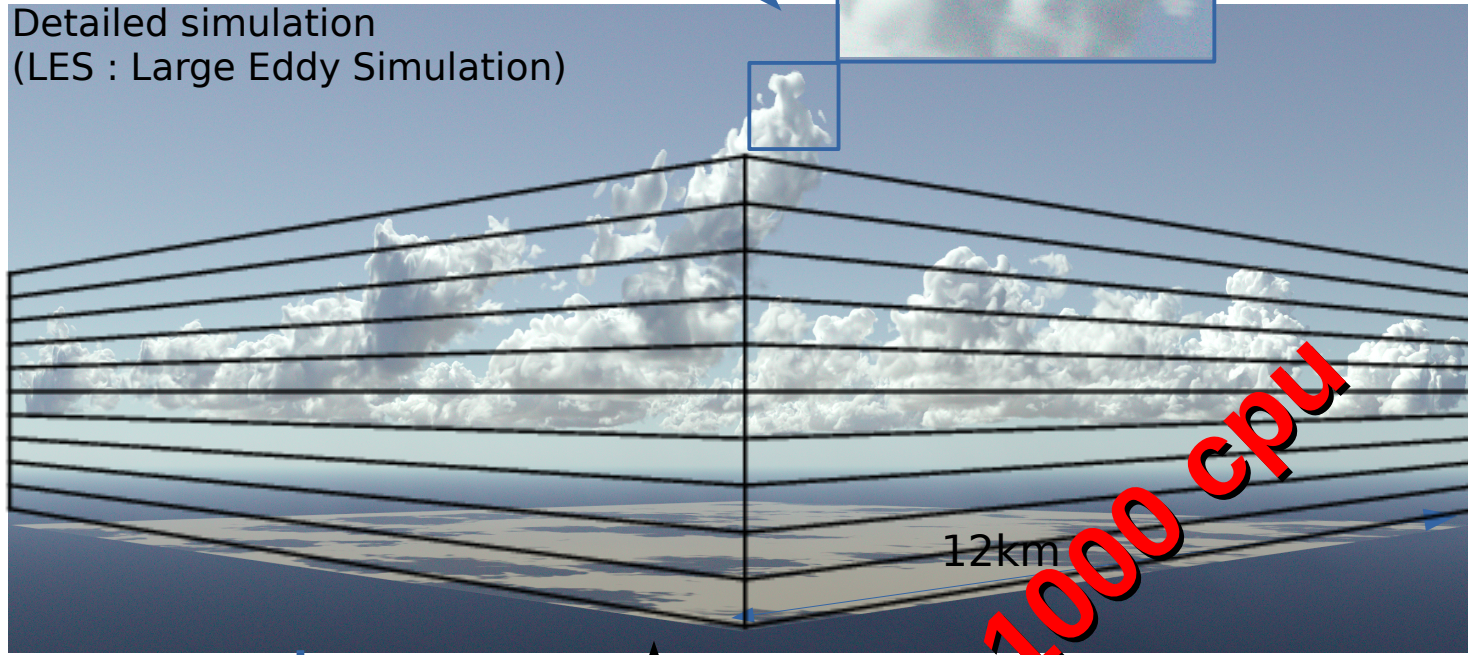


Field campaign experiment

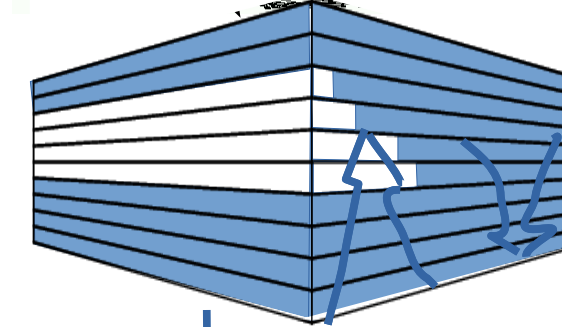
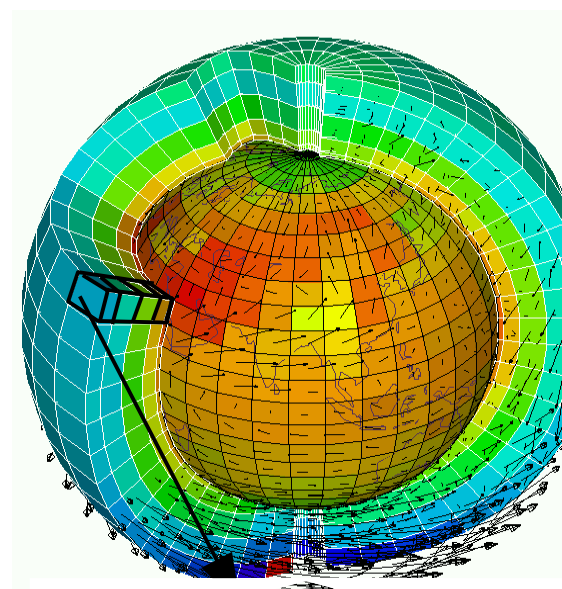
Evaluation



Detailed simulation (LES : Large Eddy Simulation)

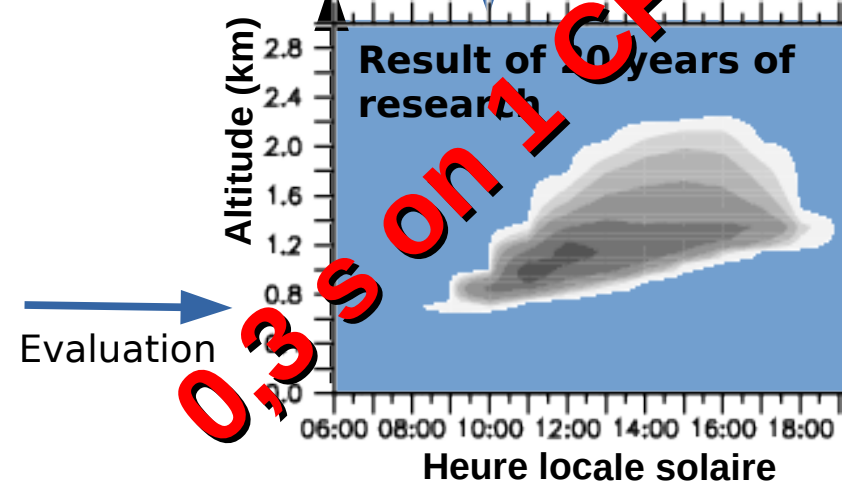
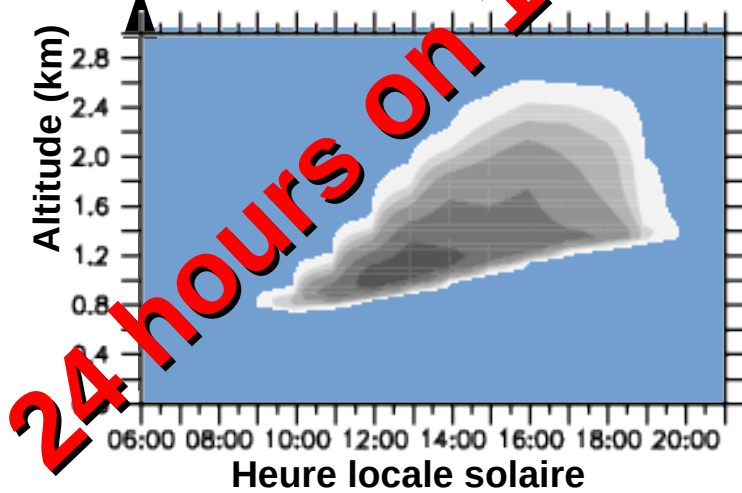


12km



Computing the cloud fraction for each cell in one column of the GCM

Computing at each altitude the fraction of the horizontal domain covered by clouds. Also called : **the cloud fraction**

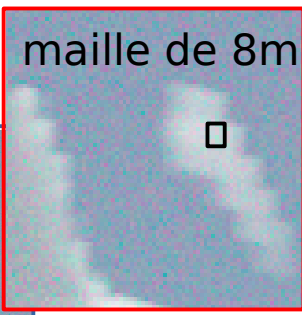
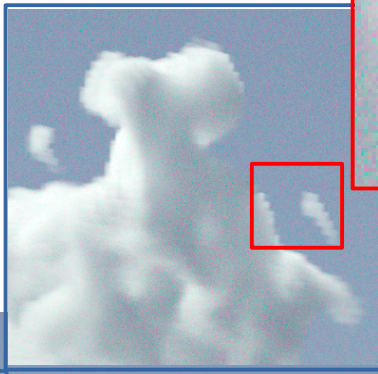


Evaluation

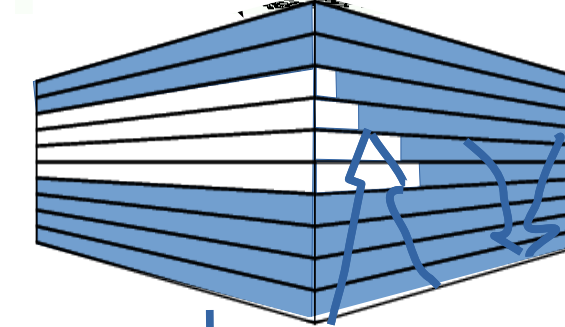
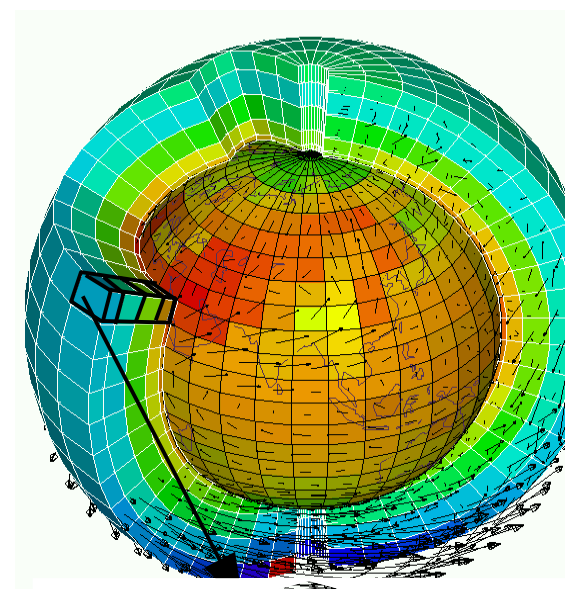
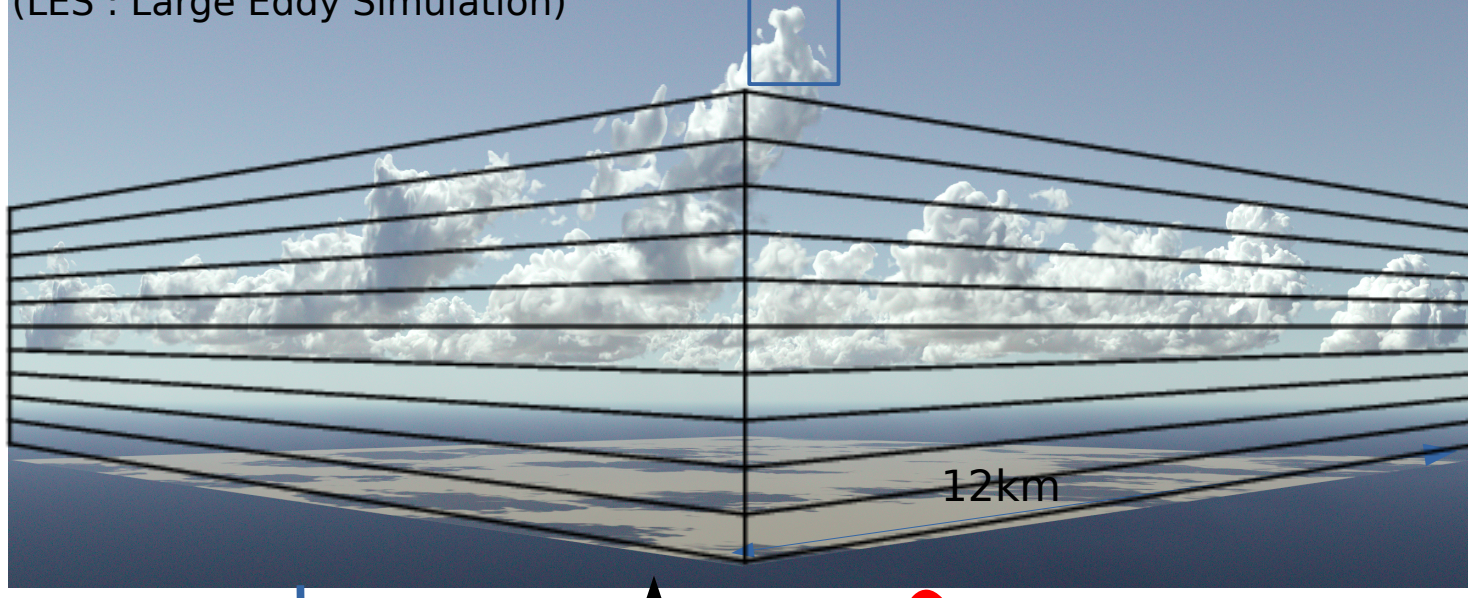


Field campaign experiment

Evaluation

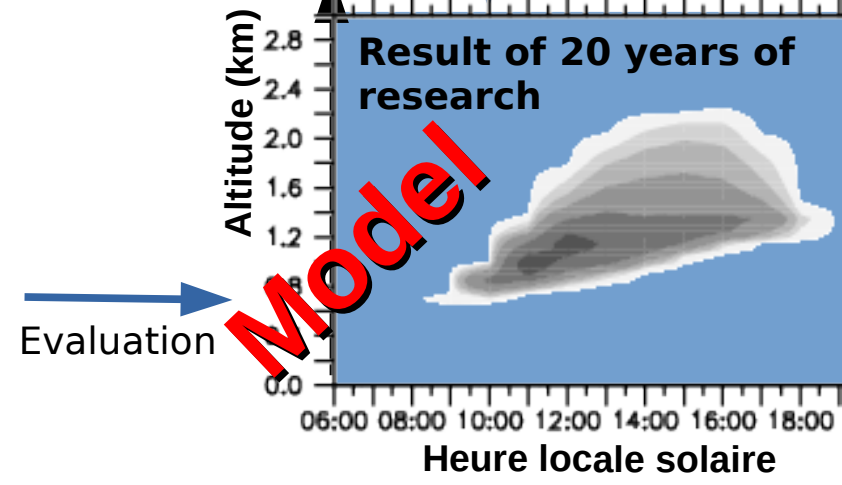
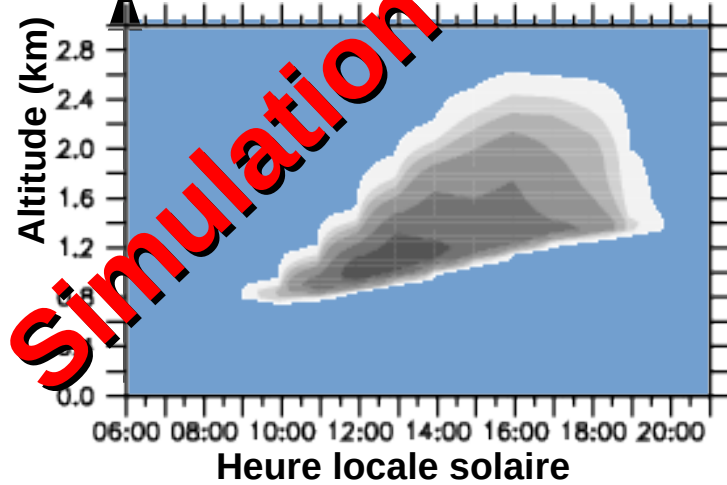


Detailed simulation (LES : Large Eddy Simulation)



Computing the cloud fraction for each cell in one column of the GCM

Computing at each altitude the fraction of the horizontal domain covered by clouds. Also called : **the cloud fraction**



Evaluation

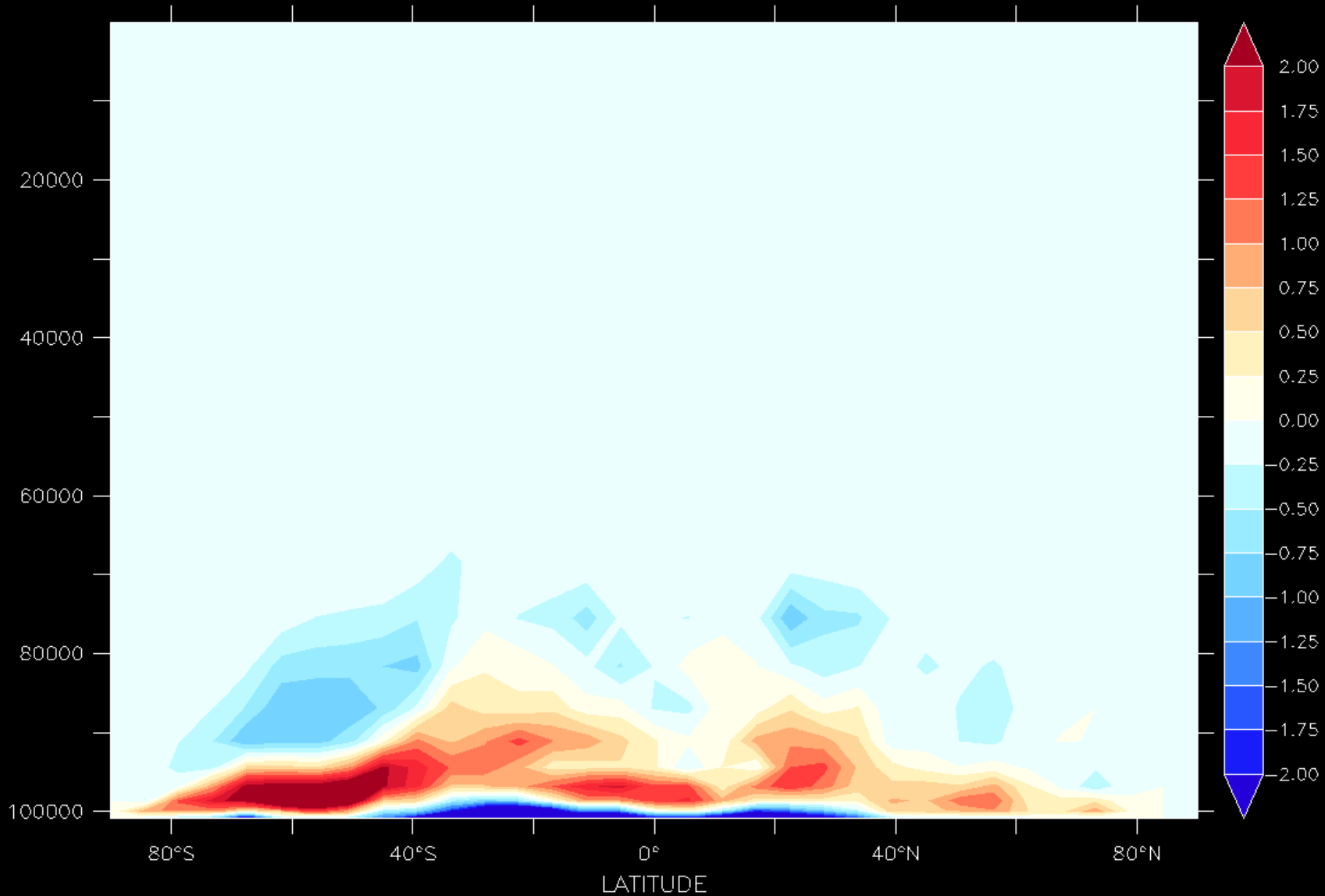
Result of 20 years of research

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LONGITUDE : 174.375E(-185.625) to 174.375E
TIME : 03-JAN-1980 12:00 360_DAY

DATA SET: histday



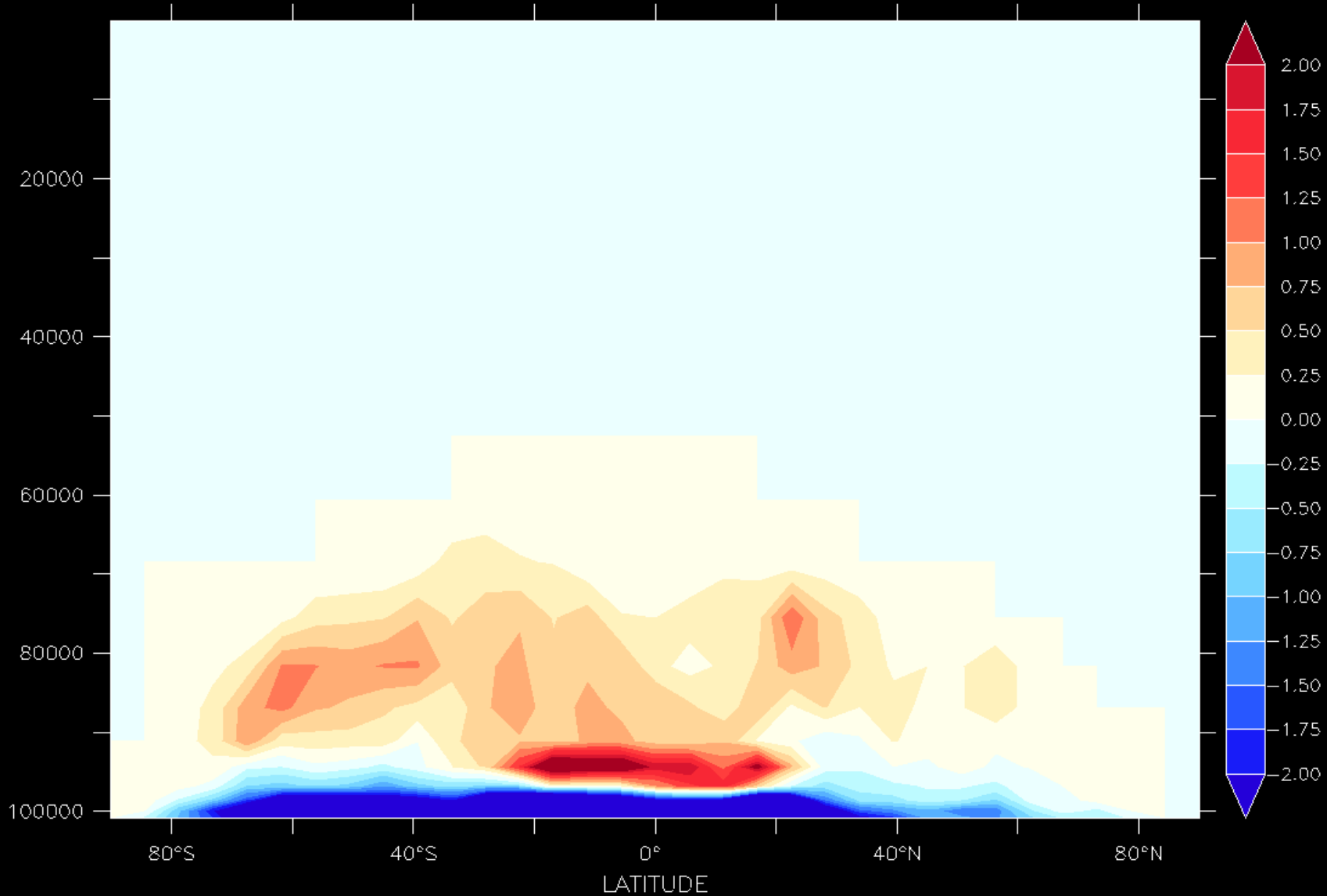
86400*DTTHE[I=@AVE]

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07-JAN-2024 19:58:14

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TIME : 03-JAN-1980 12:00 360_DAY

DATA SET: histday



1000*86400*DQTHE[I=@AVE]

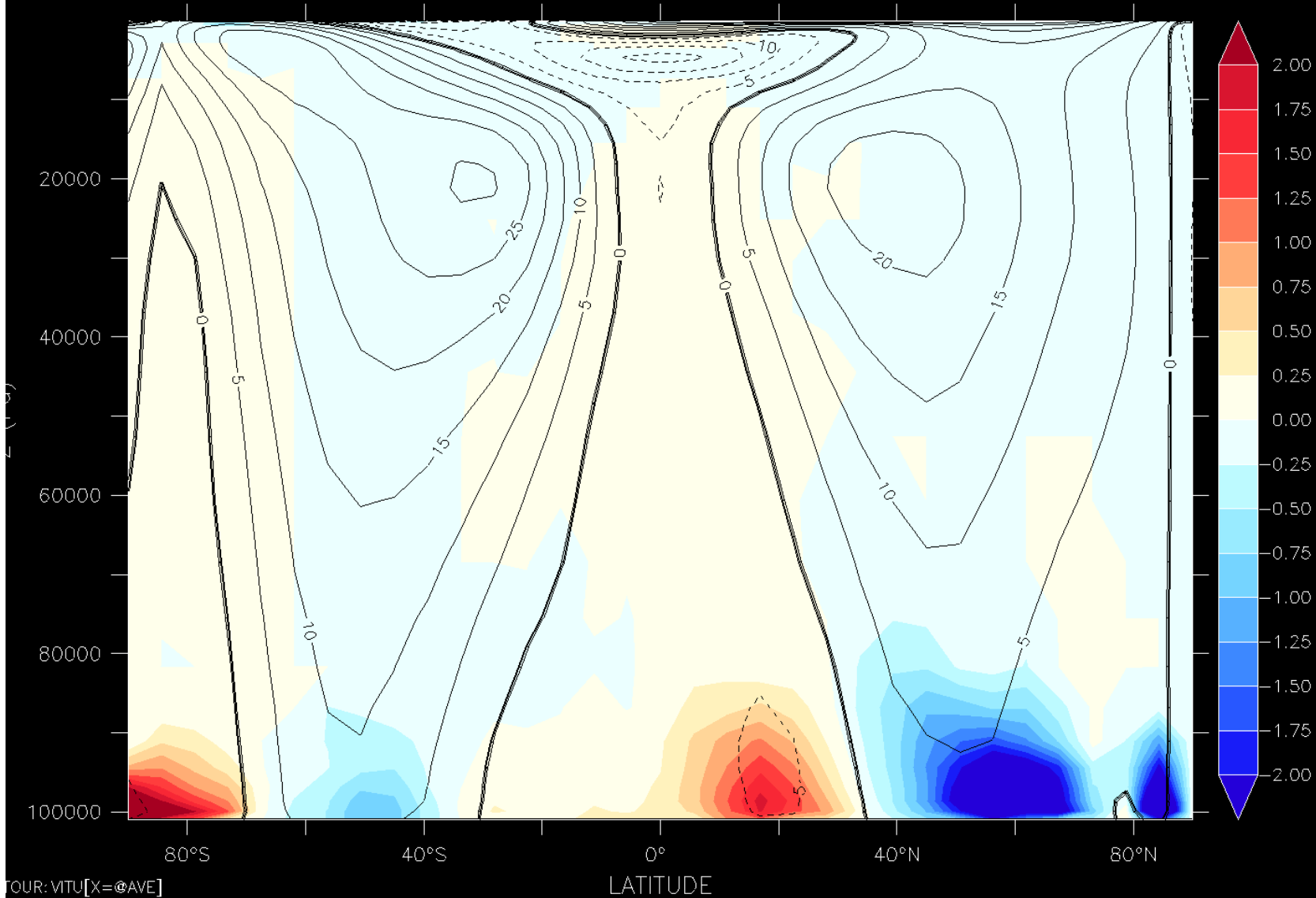
Parameterization of subgrid-scale orography

- Marine presentation

fill/lev=(-Inf)(-2,2,0.25)(inf)/pal=blue_darkred 86400*duoro[i=@ave]
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LONGITUDE : 174.375E(-185.625) to 174.375E
TIME : 03-JAN-1980 12:00 360_DAY

DATA SET: histday



86400*DUORO[I=@AVE]

Practice

**Try to analyse the diurnal cycle of temperature and humidity
Choose for instance a location in Sahel : 0W, 12N**

Temperature tendencies

Basic facts about parametrizations I

- Each parametrization : (1) works almost independently of the others ; (2) depends on vertical profiles of u , v , w , T , q and on some interface variables with the other parametrizations ; (3) ignores the spatial heterogeneities associated with the other processes (except for some processes in the deep convection scheme).
- The total tendency due to sub-grid processes is the sum of the tendencies due to each process :

$$\begin{aligned} S_T = (\partial_t T)_\varphi = & (\partial_t T)_{eva} + (\partial_t T)_{lsc} + (\partial_t T)_{diff\ turb} + (\partial_t T)_{conv} \\ & + (\partial_t T)_{wk} + (\partial_t T)_{Th} + (\partial_t T)_{ajs} \\ & + (\partial_t T)_{rad} + (\partial_t T)_{oro} + (\partial_t T)_{dissip} \end{aligned}$$

In the model, the total tendency of T for example is $\partial_t T_{dyn} + \partial_t T_{param} = dt_{dyn} + dt_{phy}$, where :

$$\begin{aligned} dt_{phy} = & dt_{eva} + dt_{lsc} + dt_{vdf} + dt_{con} + \\ & dt_{wak} + dt_{the} + dt_{ajs} + \\ & (dt_{swr} + dt_{lwr}) + (dt_{oro} + dt_{lif}) + (dt_{dis} + dt_{ec}) \end{aligned}$$

Output names
→ **Not the same as their name in the source code !**
physiq_mod.f90

Specific humidity tendencies

Basic facts about parametrizations II

- Similarly, the total tendency of a given tracer q writes :

$$S_q = (\partial_t q)_\varphi = (\partial_t q)_{\text{eva}} + (\partial_t q)_{\text{lsc}} + (\partial_t q)_{\text{diff turb}} + (\partial_t q)_{\text{conv}} \\ + (\partial_t q)_{\text{wk}} + (\partial_t q)_{\text{Th}} + (\partial_t q)_{\text{ajs}}$$

In the model, the total tendency of q is therefore

$\partial_t q_{\text{dyn}} + \partial_t q_{\text{param}} = dq_{\text{dyn}} + dq_{\text{phy}}$, where :

$dq_{\text{phy}} = dq_{\text{eva}} + dq_{\text{lsc}} + dq_{\text{vdf}} + dq_{\text{con}} + dq_{\text{wak}} + dq_{\text{the}} + dq_{\text{ajs}}$

Subroutine structure

physiq_mod.F90 structure - I

Initialization (once) : *conf_phys*, *phyetat0*,
phys_output_open

Beginning *change_srf_frac*, *solarlong*

Cloud water evap. *reevap*

Vertical diffusion (turbulent mixing) *pbl_surface*

Deep convection *conflx* (Tiedtke) or *concvl* (Emanuel)

Deep convection clouds *clouds_gno*

Density currents (wakes) *calwake*

Strato-cumulus *stratocu_if*

Thermal plumes *calltherm* and *ajsec* (sec = dry)

Large scale clouds *calcratqs*

Large scale and cumulus condensation *fisrtilp*

Diagnostic clouds for Tiedtke *diagcld1*

Aerosols *readaerosol_optic*

Cloud optical parameters *newmicro* or *nuage*

Radiative processes *radlwsr*

In blue : subroutines and instructions modifying state variables

physiq_mod.F90 structure - II

Orographic processes : drag *drag_noro_strato* or
drag_noro

Orographic processes : lift *lift_noro_strato* or
lift_noro

Orographic processes : Gravity Waves *hines_gwd* or
GWD_rando

**Axial components of angular momentum and
mountain torque** : *aaam_bud*

Cosp simulator *phys_cosp*

Tracers *phytrac*

Tracers off-line *phystokenc*

Water and energy transport *transp*

Outputs

Statistics

Output of final state (for restart) *phyredem*

Subroutine structure

physiq_mod.F90 structure - I

Initialization (once) : *conf_phys*, *phyetat0*,
phys_output_open

Beginning *change_srf_frac*, *solarlong*

Cloud water evap. *reevap*

Vertical diffusion (turbulent mixing) *pbl_surface*

Deep convection *conflx* (Tiedtke) or *concvl* (Emanuel)

Deep convection clouds *clouds_gno*

Density currents (wakes) *calwake*

Strato-cumulus *stratocu_if*

Thermal plumes *calltherm* and *ajsec* (sec = dry)

Large scale clouds *calcratqs*

Large scale and cumulus condensation *fisrtlpl*

Dagnostic clouds for Tiedtke *diagcld1*

Aerosols *readaerosol_optic*

Cloud optical parameters *newmicro* or *nuage*

Radiative processes *radlwsr*

In blue : subroutines and instructions modifying state variables

physiq_mod.F90 structure - II

Orographic processes : drag *drag_noro_strato* or *drag_noro*

Orographic processes : lift *lift_noro_strato* or *lift_noro*

Orographic processes : Gravity Waves *hines_gwd* or *GWD_rando*

Axial components of angular momentum and mountain torque : *aaam_bud*

Cosp simulator *phys_cosp*

Tracers *phytrac*

Tracers off-line *phystokenc*

Water and energy transport *transp*

Outputs

Statistics

Output of final state (for restart) *phyredem*

Effect of subrid-scale transport

Coupling with surface

Clouds and radiation

Today

Tomorrow

Radiation : output variables

Radiation I

Subroutine : radlsw

Tendencies :

dtswr, dtlwr Temperature tendencies due to solar radiation (SW = short wave) and thermal infra-red (LW = long wave)

The total radiative tendency is the sum of the SW and LW tendencies.

Other variables

- dtsw0 : clear sky SW tendency
- dtlw0 : clear sky LW tendency
- tops : net solar radiation at top of atmosphere (positive downward)
- topl : net infra-red radiation at top of atmosphere (positive upward)
- tops0, topl0 : same for clear sky
- sols : net solar radiation at surface (positive downward)
- soll : net infra-red radiation at surface (positive downward)
- sols0, soll0 : same for clear sky

New variables :

S[L]Wdn[up]TOA[SFC][clr] :

Short[Long]Wave

Downward[upward] radiative flux at

Top-Of-Atmosphere[Surface][clear-sky]

Cloud radiative effect (CRE) :

Old names : VAR - VAR0

New names : VAR - VARclr

Radiation : control parameters

In physiq.def (deepL translation)

```
#####  
#  
# Radiation  
#####  
#  
# activation of the new RRTM radiation code  
# 0: Old code and 1: RRTM (D=0)  
iflag_rrtm=1  
  
# Number of strips for SW. Set 2 if iflag_rrtm=0  
NSW=6
```

In config.def

```
#Radiative transfer code  
#*****  
# added this flag to activate/deactivate the radiation (MPL)  
# 0: no radiation. 1: radiation is activated (D=1).  
iflag_radia=1  
## Number of calls of radiation routines ( per day)  
nbapp_rad=24
```


Turbulent diffusion : output variables

Vertical diffusion

Subroutine : pbl_surface

Tendencies :

dtvdf, dqvdf, duvdf, dvvdf

Other variables

- sens : sensible heat flux at the surface (positive upward)
- evap : water vapour flux at the surface (positive upward)
- flat : latent heat flux at the surface (positive downward)
- taux, tauy : wind stress at the surface

Turbulent diffusion : control parameters

In physiq.def (deepL translation)

```
#####  
# Turbulent boundary layer  
#####  
  
# New version of Mellor and Yamada  
new_yamada4=y  
  
# Choice of numerical scheme for new_yamada4=y  
# 1 MAR diagram. Good for stable CL but destroys the stratoculus.  
# 5 MAR schema modified. Precaunise.  
yamada4_num=5  
  
# Stable boundary layer control flag  
iflag_corr_sta=4  
  
# min on the surface stability functions  
f_ri_cd_min=0.01  
  
# max of Ric for Kz. Larger decoupling for larger Ric.  
yamada4_ric=0.18  
  
# minimum mixing length for Kz  
lmixmin=0  
  
#shema of the surface layer (D:1, 1:LMD, 8:Mellor-Yamada)  
iflag_pbl=12  
  
# Thresholds for turbulent diffusion  
ksta_ter=1e-07  
ksta=1e-10
```

Radiation : input parameters

In physiq.def (deepL translation)

```
#####  
# Convective boundary layer / thermal model  
#####  
  
# Dry convection (D:0, 0:dry adjustment,=>1:thermal model)  
iflag_thermals=18  
  
# no splitting time for thermals  
# TURNS BUT POSES MORE PROBLEMS THAN IT SOLVES  
nsplit_thermals=1  
  
# tau_thermals to have a time constant on the thermals.  
# invalid  
tau_thermals=0  
  
# Flag controlling training and practice  
iflag_thermals_ed=8  
  
# We will look for the air at  $z * (1 + \text{fact\_thermals\_ed\_dz})$  to  
compute  
# training (A. Jam)  
fact_thermals_ed_dz=0.07
```

Thermal plume model : output variables

Thermals and dry adjustment

Subroutine : calltherm

Tendencies :

dtthe, dqthe, duthe, dvthe

Other variables

- dtajs : temperature tendency due to the sole dry adjustment
- dqajs : humidity tendency due to the sole dry adjustment
- a_th : fractional area of thermal plumes
- d_th : detrainment
- e_th : entrainment
- f_th : mass flux
- w_th : vertical velocity in the thermal plume (m/s, positive upward)
- q_th : total water content in the thermal plume
- zmax_th : altitude of the top of the thermal plume (m)
- f0_th : Thermal closure mass flux (kg/m².s)