



High-Tune: RenDeRer

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Overview

Monte Carlo rendering engine of scenes with inhomogeneous media

- ▶ Backward path-tracing algorithm [Villefranque et al., 2019]

Pure CPU solver

- ▶ C ANSI [Kernighan, 1988]
- ▶ GNU/Linux 64-bits
- ▶ Mixed parallelisation (OpenMP/OpenMPI)

Command line interface

UNIX philosophy in mind

Write programs that do one thing and do it well

- ▶ Write readable, simple, small, transparent and robust programs

Write programs that work together

- ▶ Use composition

Write programs that handle text streams...

- ▶ ... because that is a universal interface

"When in doubt, use brute force" (K. Thompson)

- ▶ Monte Carlo!

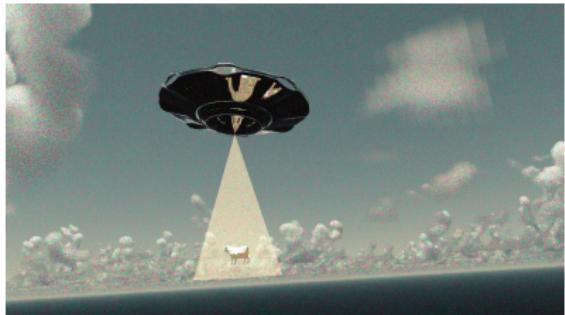
Libre Software [GPL, 2007]

Hacked by its community

- ▶ Study, modify and extend both data and code



(a) Spectral reflectivity support

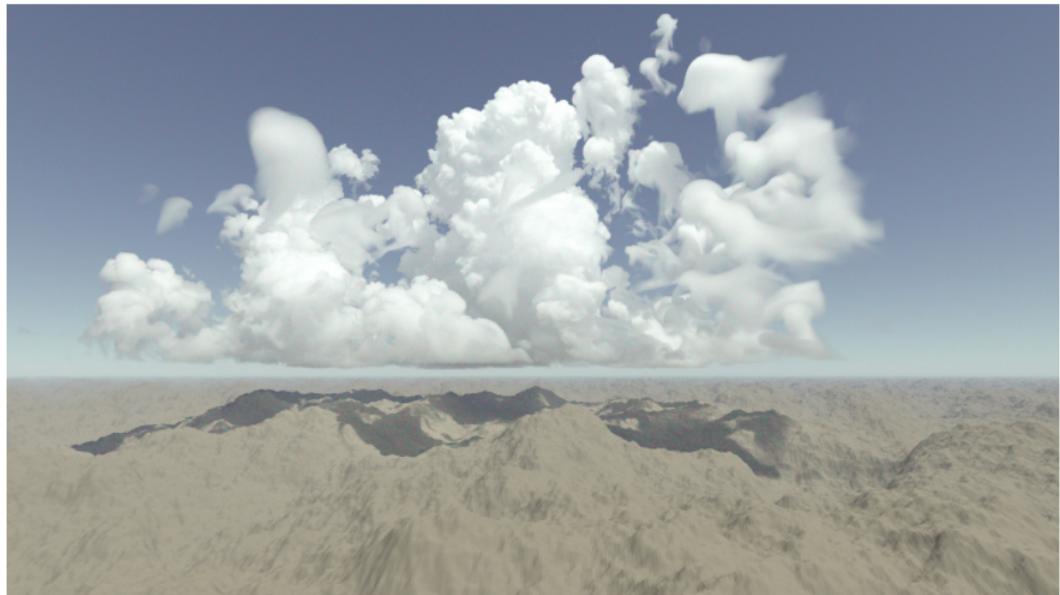


(b) Handling laser sheet [Sans et al., 2021]

History

2018, 2019¹

- ▶ Rendering of cloudy atmospheres [Villefranque et al., 2019]
- ▶ Arbitrary ground geometries



¹High-Tune ANR-16-CE01-0010 - CNRS/CNRM, CNRS/LMD, LAPLACE

History

2020²

- ▶ Infrared rendering of cloudy atmospheres
- ▶ Solve upward and downward atmospheric fluxes (visible & infrared)
- ▶ Ground geometry with spectral reflectivity [Kotthaus et al., 2013]



(a) CIE XYZ rendering

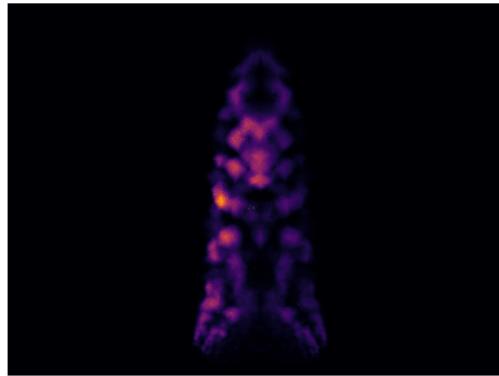


(b) Infrared rendering in $9 \mu\text{m}$ and $10 \mu\text{m}$

To come up

2020, 2021³

- ▶ Radiative transfer in soot aggregates illuminated by a laser [Sans et al., 2021]



Monochromatic shortwave rendering at 532 nm

2021⁴, ...

- ▶ Line sampling: a whole new class of radiative transfer algorithms

³Astoria ANR-18-CE05-0015 - ONERA, CNRS/RAPSOSEE, CORIA

⁴MCG-Rad ANR-18-CE46-0012 - IRIT, CNRS/LMD, LAPLACE

Data

Input data

- ▶ 1D atmospheric profile
- ▶ 3D cloud field
- ▶ Water droplets properties
- ▶ Ground geometry and materials

Output data

- ▶ Image of per pixel Monte Carlo estimations

Input data - 1D atmospheric profile

RRTM-G optical properties computed from an 1D profile of T, P, q

- ▶ Handle horizontal variation of q by pre-processing N water vapor concentrations
- ▶ Absorption coefficient (visible & infrared)
- ▶ Scattering coefficient (visible)

k-distributions

- ▶ Computed with ecRad [Hogan and Bozzo, 2016]
- ▶ Tabulated per x_{H_2O} to handle local variations
- ▶ Interpolated at runtime according to x_{H_2O}

Input data - 3D cloud field

Computed from Large Eddy Simulations
[Lafore et al., 1997, Lac et al., 2018]

Per cell data

- ▶ Water vapor mixing ratio in kg of water per m^3 of dry air
- ▶ Liquid water in suspension mixing ratio in kg of water per m^3 of dry air
- ▶ Pressure in Pascal
- ▶ Temperature in Kelvin

Input data - Water droplets properties

Computed from a Mie code [Mishchenko et al., 2002]

- ▶ Discretised over the visible & infrared spectral ranges
- ▶ Integrated onto one water droplet distribution
 - ▶ In htspk: lognormal distribution with an effective radius of $10 \mu\text{m}$ and a standard deviation of $0.1 \mu\text{m}$

Per wavelength data

- ▶ Massic absorption cross-sections
- ▶ Massic scattering cross-sections
- ▶ Asymmetric parameters of the equivalent Henyey-Greenstein phase function

Input data - Ground surface

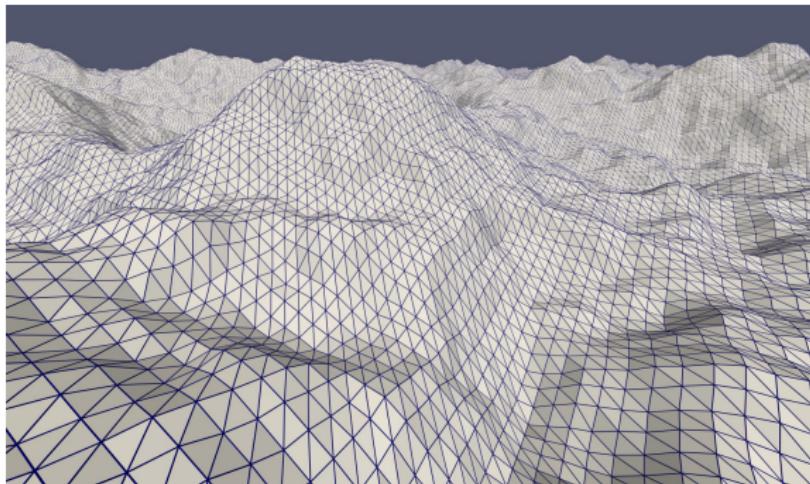
Arbitrary geometry stored in an obj-like format

Spectrally varying reflectivity (visible & infrared)

- ▶ 2 types of BRDF: lambertian and specular

```
# List of vertices
v 0 0 367.196
v 19.5312 0 353.952
v 39.0625 0 347.371
v 58.5938 0 345.122
v 78.125 0 340.736
...
v 9980.47 10000 390.285
v 10000 10000 367.196

# List of triangles
usemtl air:rock
f 1 514 2
f 2 514 515
f 2 515 3
f 3 515 516
...
f 262655 263168 262656
f 262656 263168 263169
```



Output data - Rendered image

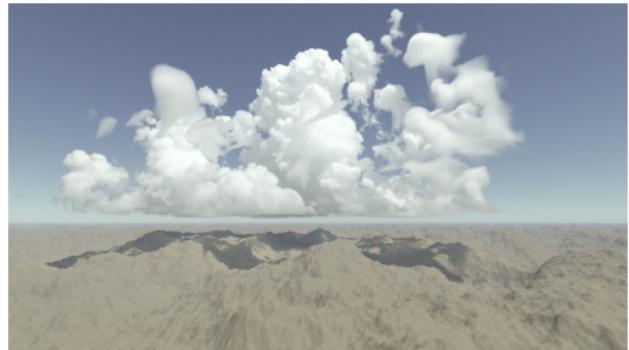
Array of per pixel Monte Carlo estimations

Camera image

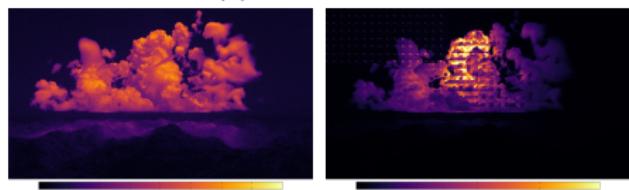
- ▶ CIE XYZ [CIE, 1931]
- ▶ Shortwave
- ▶ Longwave

Flux map

- ▶ Shortwave
- ▶ Longwave



(a) XYZ radiance



(b) X standard error

(c) Time per realisation

Beyond the renderer

Web site

- ▶ www.meso-star.com/projects/high-tune/high-tune.html

Reference documentation

- ▶ man pages

High-Tune: Starter Pack (htspk)

- ▶ Set of “ready for use” input data
- ▶ www.meso-star.com/projects/high-tune/starter-pack.html

High-Tune: Post Process (htpp)

- ▶ Post treatment of the generated images

-  (1931).

Commission International de l'Éclairage Proceedings.
Cambridge University Press, Cambridge.
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GNU General Public License.
<https://www.gnu.org/licenses/gpl.html>.
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filmicworlds.com/blog/filmic-tonemapping-operators.
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ECMWF Technical Memorandum number, 787.
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The C Programming Language.
Prentice Hall Professional Technical Reference, 2nd edition.
-  Kotthaus, S., Smith, T., Wooster, M., and Grimmond, S. (2013).

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The Meso-NH Atmospheric Simulation System. Part I: adiabatic formulation and control simulations.

Annales Geophysicae, 16(1):90–109.

-  Mishchenko, M. I., Travis, L. D., and Lacis, A. A. (2002).
Scattering, Absorption, and Emission of Light by Small Particles. Cambridge University Press, Cambridge.
-  Sans, M., Eymet, V., Villefranque, N., El Hafi, M., Fournier, R., and Forest, V. (2021).
Null-collision meshless Monte-Carlo - A new backward integral formulation designed for laser-source emission in absorbing/emitting inhomogeneous media.
<https://perso.imt-mines-albi.fr/~msans/>.
-  Villefranque, N., Fournier, R., Couvreux, F., Blanco, S., Cornet, C., Eymet, V., Forest, V., and Tregan, J.-M. (2019).
A path-tracing monte carlo library for 3-d radiative transfer in highly resolved cloudy atmospheres.
Journal of Advances in Modeling Earth Systems, 11(8):2449–2473.

Converting a intensity spectral signal to a color

Rendering in the XYZ
colorimetry space [CIE, 1931]

$$X = \int_0^{+\infty} \bar{x}(\lambda) L_\lambda(\vec{x}, -\vec{u}) d\lambda \quad (1)$$

$$Y = \int_0^{+\infty} \bar{y}(\lambda) L_\lambda(\vec{x}, -\vec{u}) d\lambda \quad (2)$$

$$Z = \int_0^{+\infty} \bar{z}(\lambda) L_\lambda(\vec{x}, -\vec{u}) d\lambda \quad (3)$$

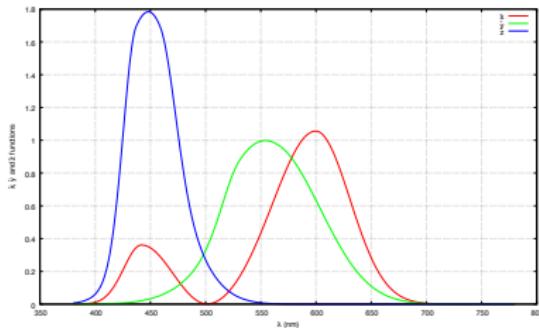


Figure: Functions $\bar{x}(\lambda)$, $\bar{y}(\lambda)$ et $\bar{z}(\lambda)$

XYZ → sRGB conversion : “High-Tune: Post-Process”

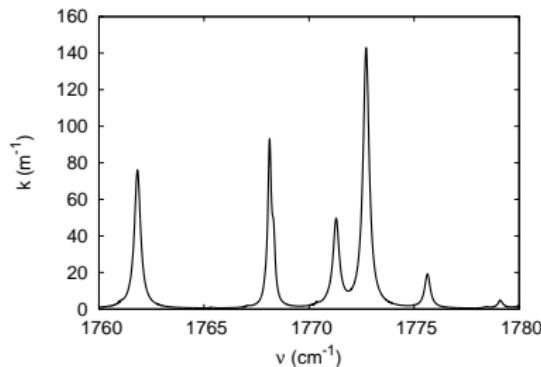
- ▶ <https://gitlab.com/meso-star/htpp.git>
- ▶ Tone mapping [Hable, 2010]
- ▶ XYZ → sRGB conversion : $C_{sRGB} = [M].C_{XYZ}$
- ▶ gamma correction

The k-distribution spectral model (1/4)

The spectral integration needs to be performed over the $[\nu_{min}, \nu_{max}]$ interval :

$$\frac{1}{\nu_{max} - \nu_{min}} \int_{\nu_{min}}^{\nu_{max}} A(\nu) d\nu \quad (4)$$

With A the monochromatic radiative quantity.

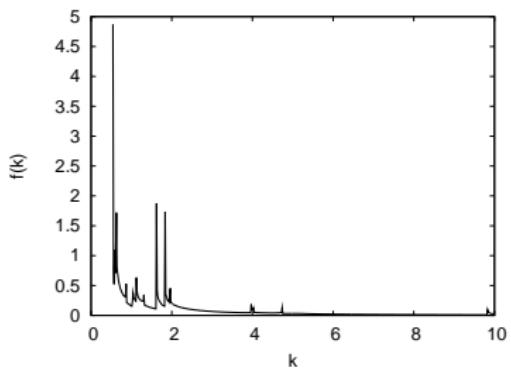
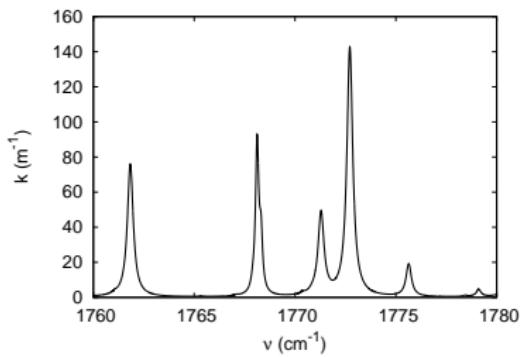


The k-distribution spectral model (2/4)

$$\frac{1}{\nu_{max} - \nu_{min}} \int_{\nu_{min}}^{\nu_{max}} A(\nu) d\nu = \int_0^{+\infty} A(k) f(k) dk \quad (5)$$

$f(k)$: distribution function for $k(\nu)$

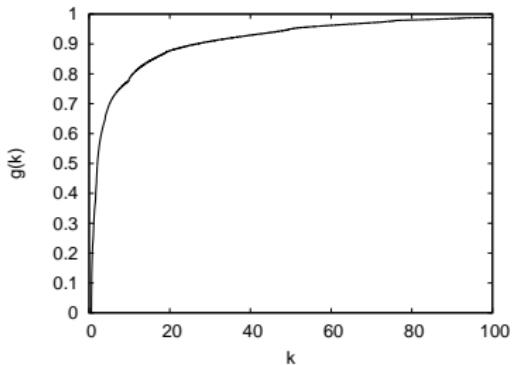
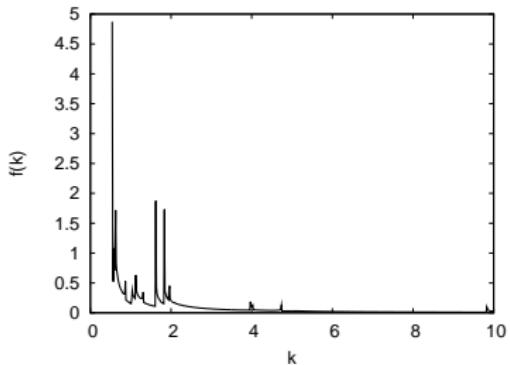
$f(k)dk$: fraction of $\nu_{max}-\nu_{min}$ where $k \in [k; k + dk]$



The k-distribution spectral model (3/4)

$$\frac{1}{\nu_{max} - \nu_{min}} \int_{\nu_{min}}^{\nu_{max}} A(\nu) d\nu = \int_0^{+\infty} A(k) f(k) dk = \int_0^1 A(k(g)) dg \quad (6)$$

with $g(k) = \int_0^k f(k') dk'$ the cumulative of $f(k)$



The k-distribution spectral model (4/4)

$$\frac{1}{\nu_{max} - \nu_{min}} \int_{\nu_{min}}^{\nu_{max}} A(\nu) d\nu = \int_0^1 A(k(g)) dg \approx \sum_{i=1}^N A(k(g_i)) \omega_i \quad (7)$$

g_i : quadrature abscissae

ω_i : quadrature weight

$k(g_i) \equiv k_i$: values of k

Advantage

Instead of performing thousands of monochromatic radiative transfer computations, only $N \approx 10$ independant computations are required for a spectral interval a few tenths of cm^{-1} wide.

Disadvantage

For a heterogeneous medium, a small error ($\approx 5\%$) is introduced: when T or P change, the $k(\nu)$ spectrum is modified and a given value of g_i/k_i no longer corresponds to the same pool of frequencies. The $k(\nu)$ spectrum is therefore supposed to be homothetic when thermodynamic conditions are modified, in order to assume a radiative transfer computation for a given k_i is equivalent to a series of monochromatic radiative transfer computations ("correlated-k" hypothesis, more or less relevant).



Using High-Tune: RenDeRer

Install

Prerequisites

- ▶ git and git-lfs
- ▶ GNU Compuler Collection (version ≥ 4.8)
- ▶ CMake (version ≥ 3)
- ▶ NetCDF library and headers (version ≥ 4)
- ▶ OpenMPI library and headers
- ▶ AsciiDoc to generate the man pages (optional)

Install htrdr

```
~ $ git clone -b High-Tune-0.6.1 \
  https://gitlab.com/meso-star/star-engine.git High-Tune-0.6.1
~ $ mkdir High-Tune-0.6.1/build; cd High-Tune-0.6.1/build
~/High-Tune-0.6.1/build $ cmake ../cmake
~/High-Tune-0.6.1/build $ make
```

Install htspk

```
~ $ wget https://www.meso-star.com/projects/high-tune/downloads/\
  High-Tune-Starter-Pack-0.6.0.tar.gz
~ $ tar xzvf High-Tune-Starter-Pack-0.6.0.tar.gz
```

Setup the working environment

Register htrdr against the current GNU/Bash shell

```
~ $ source ~/High-Tune-0.6.1/local/etc/high_tune.profile  
~ $ htrdr -h  
~ $ man htrdr  
~ $ export HTSPK=~/High-Tune-Starter-Pack-0.6.0
```

Create the working directory

```
~ $ mkdir Tuto  
~ $ cd Tuto  
~/Tuto $ echo "Hello, world!"
```

Clear sky rendering

First rendering

```
~/Tuto $ htrdr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/plane.mtls \
-g $HTSPK/models/plane.obj -R \
-i def=640x480:spp=128 \
-C pos=0,0,100:tgt=0,1,100 \
-o clear_sky.txt
```

Display the rendered image

```
~/Tuto $ htpp -h
~/Tuto $ man htpp
~/Tuto $ htpp clear_sky.txt | display -
~/Tuto $ htpp -o clear_sky.ppm clear_sky.txt
~/Tuto $ display clear_sky.ppm
~/Tuto $ htpp -i exposure=0.2 -o clear_sky_0.2.ppm clear_sky.txt
~/Tuto $ display clear_sky.ppm clear_sky_0.2.ppm
```

Overwrite the default sun position

```
~/Tuto $ htrdr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/plane.mtls \
-g $HTSPK/models/plane.obj -R \
-i def=640x480:spp=128 \
-C pos=0,0,100:tgt=0,1,100 \
-D 90,20 \
-o clear_sky.txt -f

~/Tuto $ htpp -i exposure=0.2 clear_sky.txt | display -
```

Update the camera

```
~/Tuto $ htrdr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/plane.mtls \
-g $HTSPK/models/plane.obj -R \
-i def=640x480:spp=128 \
-C pos=0,0,100:tgt=0,1,100.5:up=0,0,1:fov=60 \
-D 90,20 \
-o clear_sky.txt -f

~/Tuto $ htpp -i exposure=0.2 clear_sky.txt | display -
```

Cloud field rendering

```
~/Tuto $ htrdr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/plane.mtls \
-g $HTSPK/models/plane.obj -R \
-i def=640x480:spp=4 \
-C pos=0,0,100:tgt=0,1,100.5:up=0,0,1:fov=60 \
-D 90,20 \
-m $HTSPK/Mie_LUT_Cloud-2-10-0.010.nc \
-c $HTSPK/clouds/DZVAR.1.ARMCU.008.diaKCL.htcp \
-O DZVAR_octrees.cache \
-o sky_DZVAR.txt

~/Tuto $ htpp -i exposure=0.2 sky_DZVAR.txt | display -
```

Infinitely repeat the clouds

```
~/Tuto $ htrdr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/plane.mtls \
-g $HTSPK/models/plane.obj -R \
-i def=640x480:spp=4 \
-C pos=0,0,100:tgt=0,1,100.5:up=0,0,1:fov=60 \
-D 90,20 \
-m $HTSPK/Mie_LUT_Cloud-2-10-0.010.nc \
-c $HTSPK/clouds/DZVAR.1.ARMCU.008.diaKCL.htcp -r \
-O DZVAR_octrees.cache \
-o sky_DZVAR.txt -f

~/Tuto $ htpp -i exposure=0.2 sky_DZVAR.txt | display -
```

Dump the octrees

```
~/Tuto $ htrdr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/plane.mtls \
-g $HTSPK/models/plane.obj \
-m $HTSPK/Mie_LUT_Cloud-2-10-0.010.nc \
-c $HTSPK/clouds/DZVAR.1.ARMCU.008.diaKCL.htcp \
-O DZVAR_octrees.cache \
-d \
-o octrees.txt
```

```
~/Tuto $ csplit \
-f cloud_octree_ \
-b %02d.vtk \
-z \
--suppress-matched \
octrees.txt \
/^---$/ *
```

```
~/Tuto $ paraview cloud_octree_11.vtk
```

Update the optical thickness criterion

```
~/Tuto $ htldr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/plane.mtls \
-g $HTSPK/models/plane.obj -R \
-i def=640x480:spp=4 \
-C pos=0,0,100:tgt=0,1,100.5:up=0,0,1:fov=60 \
-D 90,20 \
-m $HTSPK/Mie_LUT_Cloud-2-10-0.010.nc \
-c $HTSPK/clouds/DZVAR.1.ARMCU.008.diaKCL.htcp -r \
-T 0.01 \
-O DZVAR_octrees_0.01.cache \
-o sky_DZVAR_0.01.txt

~/Tuto $ httpp -i exposure=0.2 sky_DZVAR_0.01.txt | display -
```

Update the ground surface

```
~/Tuto $ htrdr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/mountain.mtls \
-g $HTSPK/models/mountain.obj -R \
-i def=640x480:spp=4 \
-C pos=0,0,600:tgt=0,1,600.1:up=0,0,1:fov=60 \
-D 90,60 \
-m $HTSPK/Mie_LUT_Cloud-2-10-0.010.nc \
-c $HTSPK/clouds/DZVAR.1.ARMCU.008.diaKCL.htcp -r \
-O DZVAR_octrees.cache \
-o sky_DZVAR_mountain.txt

~/Tuto $ htpp -i exposure=0.2 sky_DZVAR_mountain.txt | display -
```

Infrared rendering

```
~/Tuto $ htrdr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/mountain.mtls \
-g $HTSPK/models/mountain.obj -R \
-i def=640x480:spp=128 \
-C pos=0,0,600:tgt=0,1,600.1:up=0,0,1:fov=60 \
-D 90,60 \
-m $HTSPK/Mie_LUT_Cloud-2-10-0.010.nc \
-c $HTSPK/clouds/DZVAR.1.ARMCU.008.diaKCL.htcp -r \
-s lw=9000,10000:Tref=590 \
-O DZVAR_octrees_lw_9000nm_10000nm_590K.cache \
-o sky_DZVAR_mountain_lw.txt
```

Display heat map

Display expected value

```
~/Tuto $ htpp -v -m pixcpnt=0 sky_DZVAR_mountain_lw.txt | display -  
~/Tuto $ htpp -v -m pixcpnt=0:gnuplot -o heat_map.gp sky_DZVAR_mountain_lw.txt  
~/Tuto $ gnuplot heat_map.gp > heat_map.png  
~/Tuto $ display heat_map.png
```

Display standard error

```
~/Tuto $ htpp -v -m pixcpnt=1:gnuplot sky_DZVAR_mountain_lw.txt | gnuplot - | display -
```

Computing ascending flux in longwave

```
~/Tuto $ htldr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/plane.mtls \
-g $HTSPK/models/plane.obj -R \
-i def=512x512:spp=128 \
-D 0,90 \
-m $HTSPK/Mie_LUT_Cloud-2-10-0.010.nc \
-c $HTSPK/clouds/DZVAR.1.ARMCU.008.diaKCL.htcp -r \
-s lw=4000,10000 \
-O DZVAR_octrees_lw_4000nm_10000nm.cache \
-p pos=3200,3200,80000:tgt=3200,3200,0:up=0,1,0:sz=6400,6400 \
-o sky_DZVAR_plane_lwAscending_flux.txt

~/Tuto $ http -v -m default sky_DZVAR_plane_lwAscending_flux.txt | display -
```

Computing descending flux in longwave

```
~/Tuto $ htldr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/plane.mtls \
-g $HTSPK/models/plane.obj -R \
-i def=512x512:spp=128 \
-D 0,90 \
-m $HTSPK/Mie_LUT_Cloud-2-10-0.010.nc \
-c $HTSPK/clouds/DZVAR.1.ARMCU.008.diaKCL.htcp -r \
-s lw=4000,10000 \
-O DZVAR_octrees_lw_4000nm_10000nm.cache \
-p pos=3200,3200,1:tgt=3200,3200,2:up=0,1,0:sz=6400,6400 \
-o sky_DZVAR_plane_lw_descending_flux.txt

~/Tuto $ http -v -m default sky_DZVAR_plane_lw_descending_flux.txt | display -
```

Computing descending flux in shortwave

```
~/Tuto $ htldr -v \
-a $HTSPK/ecrad_opt_prop.txt \
-M $HTSPK/materials/plane.mtls \
-g $HTSPK/models/plane.obj -R \
-i def=512x512:spp=32 \
-D 0,90 \
-m $HTSPK/Mie_LUT_Cloud-2-10-0.010.nc \
-c $HTSPK/clouds/DZVAR.1.ARMCU.008.diaKCL.htcp -r \
-s sw=380,780 \
-O DZVAR_octrees_sw_380nm_780m.cache \
-p pos=3200,3200,1:tgt=3200,3200,2:up=0,1,0:sz=6400,6400 \
-o sky_DZVAR_plane_sw_descending_flux.txt

~/Tuto $ http -v -m range=0,900 sky_DZVAR_plane_sw_descending_flux.txt | display -
```

The ht-run.sh GNU/Bash script

Build a htrdr command line from an input file that describes the scene to render

```
~/Tuto $ cat $HTSPK/scenes/city
~/Tuto $ cat $HTSPK/ht-run.sh | more
~/Tuto $ bash $HTSPK/ht-run.sh $HTSPK/scenes/city
~/Tuto $ htpp -i exposure=0.2 city_1280x720x256.txt | display -
```



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